

Non-geometric states and Holevo information in AdS3/CFT2

Feng-Li Lin (NTNU)
@KIAS workshop 11/2018

based on 1806.07595, 1808.02873 & 1810.01258 with **Jiaju Zhang**(Milano-Bicocca) & **Wu-Zhong Guo**(NCTS)

Motivations

- Recently, there are lots of discussions on quantum thermalization in the context of AdS/CFT duality, inspired by the holographic entanglement.
- One of the key issues is how we can characterize the local (dis-)similarity between a typical pure state and thermal state, and how general it should be?
- The general principle/criterion is formulated as the **eigenstate thermalization hypothesis** (Srednicki) or **canonical typicality** (Popescu et al).
- On the bulk side, the story becomes how we can characterize the microstates of a black hole, or can we tell them locally?

- In this talk, I will discuss two issues on quantum thermalization in the context of AdS3/CFT2.
- We will examine $1/c$ behaviors through short-interval expansion of entanglement entropy.
- In CFT side, $1/c$ effects will fine-grain the microstates and lead to expected failure of ETH for integrable models.
- In the bulk side, $1/c \sim G$ effects encode quantum gravity corrections.

1. Are all CFT states geometric?

- A folklore in 3D gravity, all the geometries are described by the Banados metrics:

$$ds^2 = \frac{dy^2}{y^2} + \frac{L_\rho}{2} dz^2 + \frac{\bar{L}_\rho}{2} d\bar{z}^2 + \left(\frac{1}{y^2} + \frac{y^2}{4} L_\rho \bar{L}_\rho \right) dz d\bar{z},$$

- The holomorphic function $L_\rho(z)$ is related to the vev of boundary stress tensor:

$$\langle T(z) \rangle_\rho = -\frac{c}{12} L_\rho(z), \quad \langle \bar{T}(\bar{z}) \rangle_\rho = -\frac{c}{12} \bar{L}_\rho(\bar{z}).$$

- This seemingly suggests that every CFT state corresponds to a Banados metric.
- However, a simple example shows the opposite: **the linear supposition of the CFT quantum states**. This is because the bulk gravity is classical.
- *Can we formulate a general criterion for (non-)geometric states?*

Geometric state criterion

- By studying the (short-interval expansion of) entanglement entropy, we find some states have no bulk description a la Banados metric.
- This is based on following simple observation: **If a CFT state of order c stress tensor can be described by the Banados metric (order one in $1/c$ expansion), then**

$$S_{RT} \sim \frac{A}{4G} \sim O(c)$$

- The criterion for a CFT2 state to be bulk geometric, its entanglement/Renyi entropy calculated from CFT should be also order c at most in the large c limit.

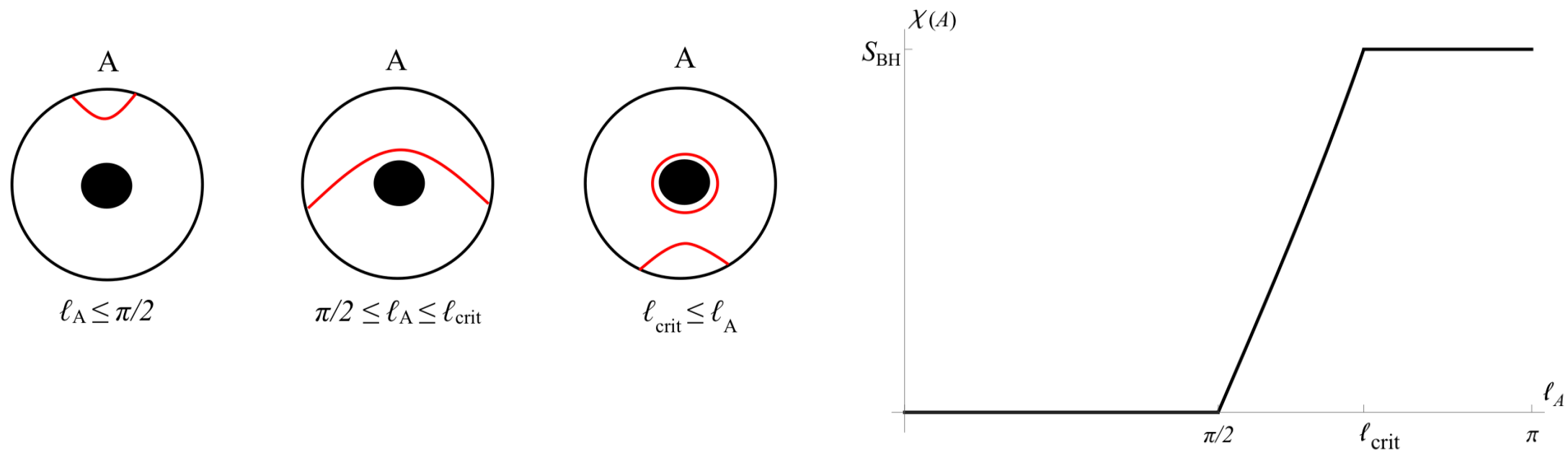
2. Can we locally distinguish the microstates of a black hole?

- By ETH or canonical typicality, a typical state locally looks like thermal state.
- In dual gravity, this means that black hole microstate geometry (aAdS metric) locally looks like BTZ black hole.
- Can we distinguish the black hole's microstates?
- In our previous works [1610.01362](#), [1703.08724](#) & [1708.05090](#), we showed that the EE of heavy primary states are the same as the EE of thermal state at leading order of c .
- Can we have quantify this (in-)distinguishability?

Holevo information

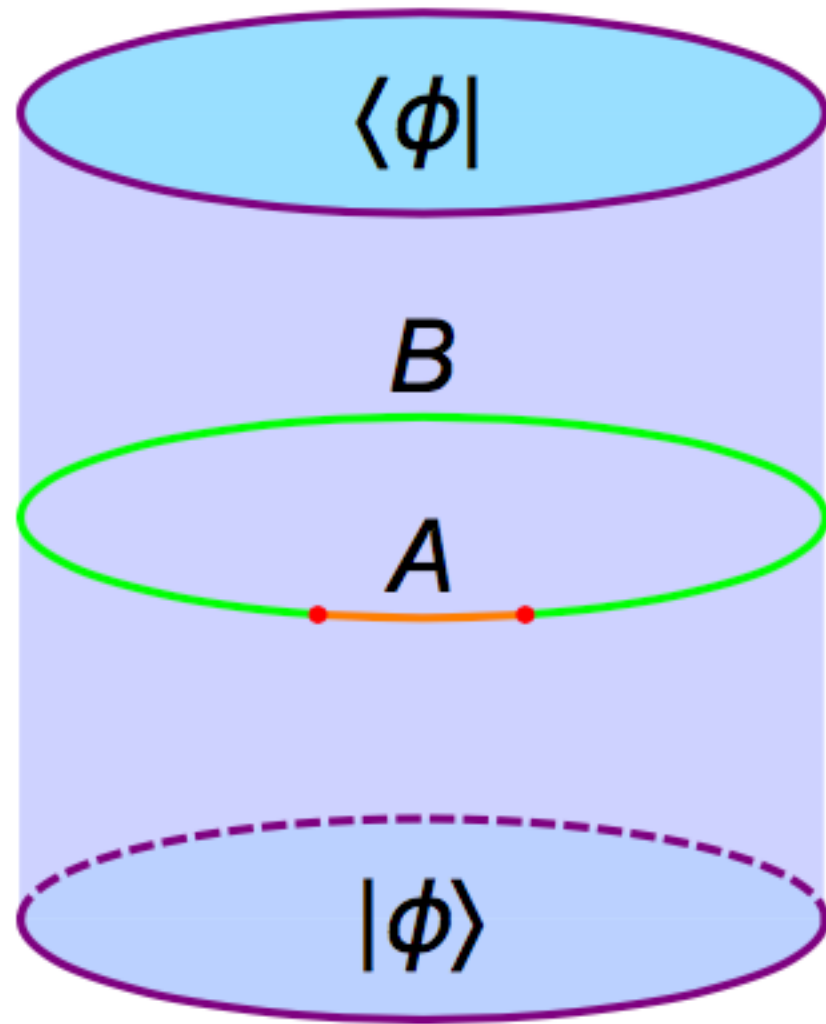
- If we encode a random variable X (with outcome probability denoted by p_i) into a quantum state $\rho = \sum_i p_i \rho_i$
- By measuring the quantum state to obtain the outcome Y , then the maximum of the mutual information $I(X;Y)$ is the Holevo information.
- Apply this to the local distinguishability of black hole microstates, the Holevo information is $\chi_A = S_A - \sum_i p_i S_{A,i}$
- χ_A is monotonically increasing with A .
- $0 \leq \chi_A \leq S_{BH} = S_{thermal} := - \sum_i p_i \log p_i$

- To calculate χ_A , the difficulty lies on how to take thermal average of EE for all CFT state.
- To bypass this, Bao & Ooguri take the facts in the large c limit: (i) light states are sparse, (ii) heavy states locally look thermal (not true for non-geometric states) (iii) verified by holographic EE



- One should expect the plateaux will be lifted by lifted by $1/c$ effect as I will show below.

Entanglement Entropy in 2D CFT



Renyi entropy

c.f. Calabrese & Cardy (2004)

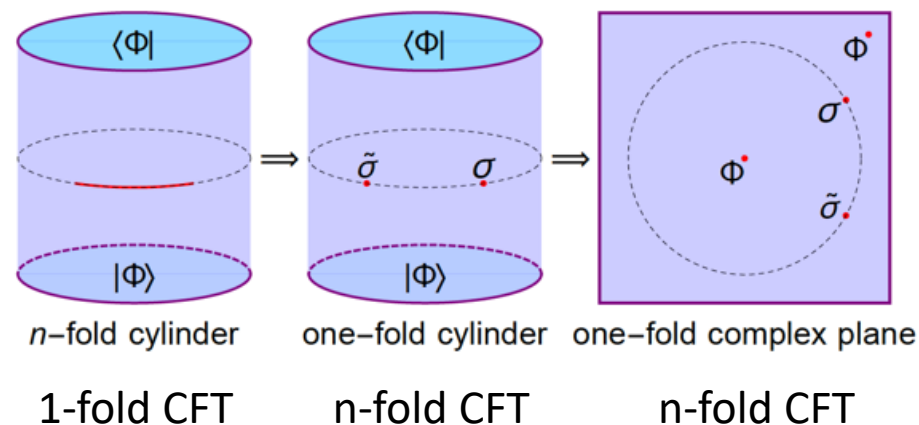
- Renyi entropy is defined as $S_n = -\frac{1}{n-1} \log \text{tr}_A \rho_A^n$,

- As $n \rightarrow 1$, we have entanglement entropy:

$$S = -\text{tr}_A(\rho_A \log \rho_A).$$

- For 2D CFT, it can be obtained by 2-point function of twist operators of n-fold CFT on Riemann surface \mathcal{R} :

$$\text{tr}_A \rho_A^n = \langle \mathcal{T}(\ell) \tilde{\mathcal{T}}(0) \rangle_{\mathcal{R}}, \quad h_{\mathcal{T}} = h_{\tilde{\mathcal{T}}} = \frac{c(n^2 - 1)}{24n}$$



OPE & Enumerating Quasi-primaries

$$\mathcal{T}(z)\tilde{\mathcal{T}}(w) = \frac{c_n}{(z-w)^{2h_\tau}} \sum_K d_K \sum_{p=0}^{\infty} \frac{c_K^p}{p!} (z-w)^{h_K+p} \partial^p \Phi_K(w),$$

$$c_K^p := \frac{C_{h_K+p-1}^p}{C_{2h_K+p-1}^p}, \quad d_K = \frac{1}{\alpha_K \ell^{h_K}} \lim_{z \rightarrow \infty} z^{2h_K} \langle \Phi_K(z) \rangle_{\mathcal{R}_{n,1}},$$

α_K = normalization of quasiprimary operator Φ_K

- Assuming translational symmetry, we have

$$\text{tr}_A \rho_A^n = \frac{c_n}{\ell^{2h_\tau}} \sum_{K \in \text{CFT}^n} d_K \ell^{h_K} \langle \Phi_K \rangle_{\mathcal{R}}$$

- Counting the quasiprimaries of n-fold CFT:

$$[x + (1-x)\text{tr} x^{L_0}]^n = 1 + nx^2 + \frac{n(n+1)}{2}x^4 + \frac{n(n^2+3n+8)}{6}x^6 + \frac{n(n+1)(n^2+5n+30)}{24}x^8 + \dots$$

$$\Phi_K^{j_1 \dots j_k} = \mathcal{X}_1^{j_1} \dots \mathcal{X}_k^{j_k} \Rightarrow \langle \mathcal{X}_1^{j_1} \dots \mathcal{X}_k^{j_k} \rangle_{\mathcal{R}} = \langle \mathcal{X}_1 \rangle_{\mathcal{R}} \dots \langle \mathcal{X}_k \rangle_{\mathcal{R}} \quad \text{with}$$

$$b_{\mathcal{X}_1 \dots \mathcal{X}_k} := \sum_{j_1, \dots, j_k} d_{\mathcal{X}_1 \dots \mathcal{X}_k}^{j_1 \dots j_k} \quad \text{with some constraints for } 0 \leq j_1, \dots, j_k \leq n-1.$$

level	quasiprimary	?	#	#
0	1	-	1	1
2	T	✓	n	n
4	\mathcal{A}	×	n	$\frac{n(n+1)}{2}$
	TT	✓	$\frac{n_2}{2}$	
6	B, D	×	$2n$	$\frac{n(n^2+3n+8)}{6}$
	TA	×	n_2	
	TTT	✓	$\frac{n_3}{6}$	
8	$\mathcal{E}, \mathcal{H}, \mathcal{I}$	×	$3n$	$\frac{n(n+1)(n^2+5n+30)}{24}$
	TB, TD	×	$2n_2$	
	AA	✓	$\frac{n_2}{2}$	
	TTA	✓	$\frac{n_3}{2}$	
	$TTTT$	✓	$\frac{n_4}{24}$	
9	$\mathcal{A}^{(9)}$	×	n	n
10	$\mathcal{A}^{(10,m)}$	×	$4n$	$\frac{n(n+1)(n+2)(n^2+7n+72)}{120}$
	$T\mathcal{E}, T\mathcal{H}, T\mathcal{I}$ AB, AD	×	$5n_2$	
	TTB, TTD	×	n_3	
	TAA	✓	$\frac{n_3}{2}$	
	$TTTA$	✓	$\frac{n_4}{6}$	
	$TTTT$	✓	$\frac{n_5}{120}$	

level	quasiprimary	?	#	#		
11	$\mathcal{A}^{(11,m)}$	×	$2n$	$n(n+1)$		
	$T\mathcal{A}^{(9)}$	×	n_2			
12	$\mathcal{A}^{(12,m)}$	×	$7n$	$\frac{n(n+3)(n^4+12n^3+169n^2+438n+640)}{720}$		
	$T\mathcal{A}^{(10,m)}$ $\mathcal{A}\mathcal{E}, \mathcal{A}\mathcal{H}, \mathcal{A}\mathcal{I}$ BD	×	$8n_2$			
	BB, DD	✓	n_2			
	TTE, TTH TTI	×	$\frac{3n_3}{2}$			
	TAB, TAD AAA	✓	$\frac{13n_3}{6}$			
	$TTTB, TTTD$ $TTAA$	✓	$\frac{7n_4}{12}$			
	$TTTTA$	✓	$\frac{n_5}{24}$			
	$TTTTTT$	✓	$\frac{n_6}{720}$			
	13	$\mathcal{A}^{(13,m)}$	×		$3n$	$\frac{n(n+1)(n+2)}{2}$
		$T\mathcal{A}^{(11,m)}$ $\mathcal{A}\mathcal{A}^{(9)}$	×		$3n_2$	
$TTA^{(9)}$		×	$\frac{n_3}{2}$			
...						

EE in small interval expansion

$$S_A^{(n)} = \frac{c(n+1)}{12n} \log \frac{\ell}{\epsilon} - \frac{1}{n-1} \log \left(1 + \sum_{k=1}^n \sum_{\{\mathcal{X}_1, \dots, \mathcal{X}_k\}} \ell^{h_{\mathcal{X}_1} + \dots + h_{\mathcal{X}_k}} b_{\mathcal{X}_1 \dots \mathcal{X}_k} \langle \mathcal{X}_1 \rangle_{\mathcal{R}} \cdots \langle \mathcal{X}_k \rangle_{\mathcal{R}} \right)$$

→ $S_A = \frac{c}{6} \log \frac{\ell}{\epsilon} + \sum_{k=1}^{\infty} \sum_{\{\mathcal{X}_1, \dots, \mathcal{X}_k\}} \ell^{h_{\mathcal{X}_1} + \dots + h_{\mathcal{X}_k}} a_{\mathcal{X}_1 \dots \mathcal{X}_k} \langle \mathcal{X}_1 \rangle_{\mathcal{R}} \cdots \langle \mathcal{X}_k \rangle_{\mathcal{R}}$ with $a_{\mathcal{X}_1 \dots \mathcal{X}_k} := - \lim_{n \rightarrow 1} \frac{b_{\mathcal{X}_1 \dots \mathcal{X}_k}}{n-1}$.

Explicitly,

$$\begin{aligned} S_A &= \frac{c}{6} \log \frac{\ell}{\epsilon} + \ell^2 a_T \langle T \rangle_{\rho} + \ell^4 a_{TT} \langle T \rangle_{\rho}^2 + \ell^6 a_{TTT} \langle T \rangle_{\rho}^3 \\ &+ \ell^8 (a_{AA} \langle A \rangle_{\rho}^2 + a_{TTA} \langle T \rangle_{\rho}^2 \langle A \rangle_{\rho} + a_{TTTT} \langle T \rangle_{\rho}^4) \\ &+ \ell^{10} (a_{TAA} \langle T \rangle_{\rho} \langle A \rangle_{\rho}^2 + a_{TTTA} \langle T \rangle_{\rho}^3 \langle A \rangle_{\rho} + a_{TTTTT} \langle T \rangle_{\rho}^5) \\ &+ \ell^{12} (a_{BB} \langle B \rangle_{\rho}^2 + a_{DD} \langle D \rangle_{\rho}^2 + a_{TAB} \langle T \rangle_{\rho} \langle A \rangle_{\rho} \langle B \rangle_{\rho} + a_{TAD} \langle T \rangle_{\rho} \langle A \rangle_{\rho} \langle D \rangle_{\rho} \\ &+ a_{AAA} \langle A \rangle_{\rho}^3 + a_{TTTB} \langle T \rangle_{\rho}^3 \langle B \rangle_{\rho} + a_{TTTD} \langle T \rangle_{\rho}^3 \langle D \rangle_{\rho} + a_{TTAA} \langle T \rangle_{\rho}^2 \langle A \rangle_{\rho}^2 \\ &+ a_{TTTTA} \langle T \rangle_{\rho}^4 \langle A \rangle_{\rho} + a_{TTTTTT} \langle T \rangle_{\rho}^6) + O(\ell^{14}). \end{aligned}$$

$$\begin{aligned} a_T &= -\frac{1}{6}, \quad a_{TT} = -\frac{1}{30c}, \quad a_{TTT} = -\frac{4}{315c^2}, \\ a_{AA} &= -\frac{1}{126c(5c+22)}, \quad a_{TTA} = \frac{1}{315c^2}, \quad a_{TTTT} = -\frac{c+8}{630c^3}, \\ a_{TAA} &= -\frac{16}{693c^2(5c+22)}, \quad a_{TTTA} = \frac{3}{385c^3}, \quad a_{TTTTT} = -\frac{16(c+5)}{3465c^4}, \\ a_{BB} &= -\frac{25}{123552c(70c+29)}, \quad a_{DD} = -\frac{70c+29}{18018c(2c-1)(5c+22)(7c+68)}, \\ a_{TAB} &= -\frac{10}{1287c^2(70c+29)}, \quad a_{TAD} = \frac{5}{3003c^2(5c+22)}, \quad a_{AAA} = \frac{4(5c+64)}{3003c^2(5c+22)^2}, \\ a_{TTTB} &= \frac{5(14c+43)}{9009c^3(70c+29)}, \quad a_{TTTD} = -\frac{2}{9009c^3}, \quad a_{TTAA} = -\frac{585c+10804}{90090c^3(5c+22)}, \\ a_{TTTTA} &= \frac{2(33c+784)}{45045c^4}, \quad a_{TTTTTT} = -\frac{2(11c^2+380c+1480)}{45045c^5}. \end{aligned}$$

EEs

- For the ground state of a CFT on a circle of size L:

$$\langle T \rangle_L = \frac{\pi^2 c}{6L^2}, \quad \langle \mathcal{A} \rangle_L = \frac{\pi^4 c(5c + 22)}{180L^4}, \quad \langle \mathcal{B} \rangle_L = -\frac{62\pi^6 c}{525L^6}, \quad \langle \mathcal{D} \rangle_L = \frac{\pi^6 c(2c - 1)(5c + 22)(7c + 68)}{216(70c + 29)L^6}$$

→
$$S_A = \frac{c}{6} \log \frac{\ell}{\epsilon} - \frac{\pi^2 c \ell^2}{36L^2} - \frac{\pi^4 c \ell^4}{1080L^4} - \frac{\pi^6 c \ell^6}{17010L^6} - \frac{\pi^8 c \ell^8}{226800L^8} - \frac{\pi^{10} c \ell^{10}}{2806650L^{10}} - \frac{691\pi^{12} c \ell^{12}}{22986463500L^{12}} + O(\ell^{14})$$

- For a low temperature state on a circle: $q := e^{-2\pi\beta/L} \ll 1$

$$\langle T \rangle_{L,q} = \frac{\pi^2}{6L^2} [c - 48q^2 - 72q^3 - 144q^4 + O(q^5)],$$

$$\langle \mathcal{A} \rangle_{L,q} = \frac{\pi^4}{180L^4} [c(5c + 22) + 480(5c + 22)q^2 + 2160(5c + 22)q^3 + 30240(c + 6)q^4 + O(q^5)],$$

$$\langle \mathcal{B} \rangle_{L,q} = -\frac{2\pi^6}{525L^6} [31c - 1008(120c + 1)q^2 - 1512(720c + 161)q^3 - 3024(1640c + 841)q^4 + O(q^5)],$$

$$\langle \mathcal{D} \rangle_{L,q} = \frac{\pi^6(2c - 1)(7c + 68)}{216(70c + 29)L^6} [c(5c + 22) + 1584(5c + 22)q^2 + 6696(5c + 22)q^3 + 432(215c - 638)q^4 + O(q^5)]$$

→
$$S_A = \frac{c}{6} \log \frac{\ell}{\epsilon} + \left[-\frac{c}{36} + \frac{4q^2}{3} + 2q^3 + 4q^4 + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^2 + \left[-\frac{c}{1080} + \frac{4q^2}{45} + \frac{2q^3}{15} + \frac{4(c-8)q^4}{15c} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^4 + \left[-\frac{c}{17010} + \frac{8q^2}{945} + \frac{4q^3}{315} + \frac{8(c-16)q^4}{315c} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^6 + \left[-\frac{c}{226800} + \frac{4q^2}{4725} + \frac{2q^3}{1575} - \frac{4(159c + 728)q^4}{1575c} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^8 + \left[-\frac{c}{2806650} + \frac{8q^2}{93555} + \frac{4q^3}{31185} - \frac{104(295c + 1312)q^4}{155925c} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^{10} + \left[-\frac{691c}{22986463500} + \frac{5528q^2}{638512875} + \frac{2764q^3}{212837625} - \frac{8(21728429c + 15283768)q^4}{212837625c} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^{12} + O(\ell^{14})$$

Geometric vs Non-geometric

Conditions on expectation values of quasiprimaries

- As shown, the expectation values of the quasiprimaries dictates the entanglement/Renyi entropy.
- Requiring the Renyi entropy of A in a state ρ to be at most order c yield the following constraints:

$$\langle T \rangle_\rho = c\alpha(w) + \beta(w) + \frac{\gamma(w)}{c} + O\left(\frac{1}{c^2}\right),$$

$$\langle \mathcal{A} \rangle_\rho = c^2\alpha(w)^2 + c\delta(w) + \epsilon(w) + O\left(\frac{1}{c}\right),$$

$$\langle \mathcal{B} \rangle_\rho = c^2\left[\alpha'(w)^2 - \frac{4}{5}\alpha(w)\alpha''(w)\right] + c\zeta(w) + O(c^0),$$

$$\langle \mathcal{D} \rangle_\rho = c^3\alpha(w)^3 + 3c^2\alpha(w)[\delta(w) - \alpha(w)\beta(w)] + c\eta(w) + O(c^0),$$

$$\langle \mathcal{E} \rangle_\rho = c^2\left\{\alpha''(w)^2 + \frac{10}{63}[\alpha(w)\alpha^{(4)}(w) - 7\alpha'(w)\alpha^{(3)}(w)]\right\} + O(c),$$

$$\begin{aligned} \langle \mathcal{H} \rangle_\rho = c^3\alpha(w)\left[\alpha'(w)^2 - \frac{4}{5}\alpha(w)\alpha''(w)\right] + c^2\left[-\alpha'(w)^2\beta(w) - 2\alpha(w)\alpha'(w)\beta'(w) + \frac{4}{5}\alpha(w)^2\beta''(w) \right. \\ \left. + \frac{8}{5}\alpha(w)\alpha''(w)\beta(w) + \alpha'(w)\delta'(w) - \frac{4}{5}\alpha''(w)\delta(w) - \frac{2}{9}\alpha(w)\delta''(w) + \frac{13}{9}\alpha(w)\zeta(w)\right] + O(c), \end{aligned}$$

$$\begin{aligned} \langle \mathcal{I} \rangle_\rho = c^4\alpha(w)^4 + 2c^3\alpha(w)^2[3\delta(w) - 4\alpha(w)\beta(w)] + c^2[12\alpha(w)^2\beta(w)^2 + 4\alpha(w)^3\gamma(w) \\ - 12\alpha(w)\beta(w)\delta(w) + 3\delta(w)^2 - 6\alpha(w)^2\epsilon(w) + 4\alpha(w)\eta(w)] + O(c) \end{aligned}$$

c.f. $n \rightarrow 1$ for EE:

$$\langle T \rangle_\rho = c\alpha(w) + O(c^0),$$

$$\langle \mathcal{A} \rangle_\rho = c^2\alpha(w)^2 + O(c).$$

with $\alpha(w)$, $\beta(w)$, $\gamma(w)$, $\delta(w)$, $\epsilon(w)$, $\zeta(w)$, $\eta(w)$ being arbitrary order c^0 holomorphic functions.

Examples

- Heavy primary states are geometrical as can be checked straightforwardly.
- Thermal states (dual to BTZ) satisfy the constraints as expected. The above twos are consistent with the ETH.
- A more nontrivial example is the “coherent” CFT state, which is dual to a bulk moving particle:
$$\rho_{\phi(z_0)} = \frac{1}{\alpha_{\phi}} \left(\frac{1 - z_0 \bar{z}_0}{\bar{z}_0} \right)^{2h_{\phi}} \phi(z_0) |0\rangle \langle 0| \phi(1/\bar{z}_0).$$
- This coherent state example suggests existence of some kind of correspondence principle for the geometric states.

Correspondence Principle

- We look for the quantum observables (e.g., KdV currents) whose one-point functions w.r.t. geometric states obey the classical counterpart of the equations of motion.

Quantum KdV equations: $\dot{T} = \frac{1-c}{6}T'''' - 3(TT)' = -\frac{5c+22}{30}T'''' - 3\mathcal{A}'.$



$$J_{2k}^\rho(w) \equiv \lim_{c \rightarrow \infty} \frac{6^k}{c^k} \langle J_{2k}(w) \rangle_\rho$$

$$Q_{2k-1} = \int_0^L \frac{dw}{L} J_{2k}(w)$$

$$J_2 = T, \quad J_4 = \mathcal{A} + \frac{3}{10}T'', \quad J_6 = \mathcal{D} - \frac{25(2c+7)(7c+68)}{108(70c+29)}\mathcal{B} - \frac{2c-23}{108}\mathcal{A}'' - \frac{c-14}{280}T^{(4)},$$

$$J_8 = \mathcal{I} - \frac{5(3c+46)(350c^2+2315c-361)}{39(1050c^2+3305c-251)}\mathcal{H} + \frac{49(3c+46)(5c^2+46c+99)}{2860(105c+11)}\mathcal{E} + \frac{5c+94}{234}\mathcal{D}''$$

$$- \frac{490c^3+6713c^2+6887c+53032}{1404(70c+29)}\mathcal{B}'' + \frac{(c+7)(7c+89)}{11880}\mathcal{A}^{(4)} + \frac{7c^2+68c-315}{9072}T^{(6)}.$$

Classical KdV equations: $\dot{U} = U''' + 6UU'.$

$$J_2^\rho = U, \quad J_4^\rho = U^2, \quad J_6^\rho = U^3 - \frac{1}{2}U'^2,$$

$$J_8^\rho = U^4 + U^2U'' + \frac{1}{5}UU^{(4)}.$$

non-examples

- As the classical gravity cannot reflect the quantum linear superposition principle, we expect the superpositions of the primary states are non-geometric. This is indeed the case.
- Besides, we find the following descendant states are non-geometric:

$$|\phi^{(m)}\rangle \text{ with } h_\phi + m \sim O(c),$$

$$|\tilde{\phi}\rangle \text{ with } h_\phi \sim O(c),$$

$$|\tilde{\phi}^{(m)}\rangle \text{ with } h_\phi + m \sim O(c),$$

$$|T^{(m)}\rangle \text{ with } m \sim O(c),$$

$$|\mathcal{A}^{(m)}\rangle \text{ with } m \sim O(c).$$

$$\tilde{\phi} := (T\phi) - \frac{3}{2(h_\phi + 1)}\phi''$$

Implications

- In perturbative quantum gravity, the metric is corrected by $G \sim 1/c$ correction, and thus we will not expect these effect will turn the non-geometric states into the geometric ones.
- By Cardy's formula, we find that the fraction of primary states is negligible in the thermodynamical limit. If most of the descendant states are non-geometric, they will be also non-thermal. This will then implies the canonical typicality of quantum thermalization fails in the large c 2D CFTs.

Holevo information of microstates

Thermal average of EEs

- To find the $1/c$ correction to the Holevo information, we need to find the thermal average of EEs over all states (inclu. primaries and descendants).
- In the small interval expansion of EE, this can be done by finding the thermal average of one-point functions (and its powers) of quasiprimaries. This relies on the following:

$$\sum_i p_i \langle \mathcal{X} \rangle_{\rho_i} = \langle \mathcal{X} \rangle_{\beta}, \quad \mathcal{X} = T, \mathcal{A},$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^r X_i = \left(\frac{2\pi}{L} \right)^r \frac{\partial_{\beta}^r (\exp^{\frac{\pi c L}{12\beta}} \sum_i p_i X_i)}{\exp^{\frac{\pi c L}{12\beta}}}, \quad \beta \ll L$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^r = \left(\frac{2\pi i}{L} \right)^{2r} \frac{(q \partial_q)^r Z(q)}{Z(q)}, \quad \text{with } Z(q) = q^{-\frac{c}{24}} [1 + q^2 + q^3 + 2q^4 + O(q^5)], \quad q := e^{-\frac{2\pi\beta}{L}} \ll 1$$

$$\sum_i p_i \langle \mathcal{X} \rangle_{\rho_i} \langle \mathcal{Y} \rangle_{\rho_i} = \frac{1}{L} \int_{-L/2}^{L/2} dx \langle \mathcal{X}(x) \mathcal{Y}(0) \rangle_{\beta} \iff [T, \mathcal{A}] = 0.$$

High T case:

$$\sum_i p_i \langle T \rangle_{\rho_i} = -\frac{\pi^2 c}{6\beta^2}, \quad \sum_i p_i \langle T \rangle_{\rho_i}^2 = \frac{\pi^3 c(\pi c L + 24\beta)}{36\beta^4 L}, \quad \sum_i p_i \langle T \rangle_{\rho_i}^3 = -\frac{\pi^4 c(\pi^2 c^2 L^2 + 72\pi c\beta L + 864\beta^2)}{216\beta^6 L^2},$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^4 = \frac{\pi^5 c(\pi^3 c^3 L^3 + 144\pi^2 c^2 \beta L^2 + 5184\pi c\beta^2 L + 41472\beta^3)}{1296\beta^8 L^3},$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^5 = -\frac{\pi^6 c(\pi^4 c^4 L^4 + 240\pi^3 c^3 \beta L^3 + 17280\pi^2 c^2 \beta^2 L^2 + 414720\pi c\beta^3 L + 2488320\beta^4)}{7776\beta^{10} L^4},$$

$$\sum_i p_i \langle \mathcal{A} \rangle_{\rho_i} = \frac{\pi^4 c(5c + 22)}{180\beta^4}, \quad \sum_i p_i \langle T \rangle_{\rho_i} \langle \mathcal{A} \rangle_{\rho_i} = -\frac{\pi^5 c(5c + 22)(\pi c L + 48\beta)}{1080\beta^6 L},$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^2 \langle \mathcal{A} \rangle_{\rho_i} = \frac{\pi^6 c(5c + 22)(\pi^2 c^2 L^2 + 120\pi c\beta L + 2880\beta^2)}{6480\beta^8 L^2},$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^3 \langle \mathcal{A} \rangle_{\rho_i} = -\frac{\pi^7 c(5c + 22)(\pi^3 c^3 L^3 + 216\pi^2 c^2 \beta L^2 + 12960\pi c\beta^2 L + 207360\beta^3)}{38880\beta^{10} L^3},$$

$$\sum_i p_i \langle \mathcal{A} \rangle_{\rho_i}^2 = \frac{\pi^7 c(5c + 22)[7\pi c(5c + 22)L + 480(7c + 74)\beta]}{226800\beta^8 L},$$

$$\sum_i p_i \langle T \rangle_{\rho_i} \langle \mathcal{A} \rangle_{\rho_i}^2 = -\frac{\pi^8 c(5c + 22)[7\pi^2 c^2(5c + 22)L^2 + 192\pi c(35c + 262)\beta L + 40320(7c + 74)\beta^2]}{1360800\beta^{10} L^2}.$$



$$\chi_A = \frac{2\pi^3 \ell^4}{45\beta^3 L} - \frac{8\pi^4 \ell^6 (\pi c L + 12\beta)}{945c\beta^5 L^2} + \dots + O(\ell^{12}). \quad \text{NB: order } c^0 \text{ !!}$$

To the order we considered, it vanishes in the thermodynamic limit, i.e., $L \rightarrow \infty$ but with β, ℓ fixed

Low T case:

$$\sum_i p_i \langle T \rangle_{\rho_i} = \langle T \rangle_{\mathcal{T}} = \frac{\pi^2 c}{6L^2} - \frac{8\pi^2 q^2}{L^2} - \frac{12\pi^2 q^3}{L^2} - \frac{24\pi^2 q^4}{L^2} + O(q^5)$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^2 = \frac{\pi^4 c^2}{36L^4} - \frac{8\pi^4 (c - 24)q^2}{3L^4} - \frac{4\pi^4 (c - 36)q^3}{L^4} - \frac{8\pi^4 (c - 56)q^4}{L^4} + O(q^5)$$

$$\sum_i p_i \langle T \rangle_{\rho_i}^3 = \frac{\pi^6 c^3}{216L^6} - \frac{2\pi^6 (c^2 - 48c + 768)q^2}{3L^6} - \frac{\pi^6 (c^2 - 72c + 1728)q^3}{L^6} - \frac{2\pi^6 (c^2 - 112c + 3840)q^4}{L^6} + O(q^5),$$



NB: order $1/c$!!

$$\begin{aligned} \chi_A = & \left[\frac{32q^2}{15c} + \frac{24q^3}{5c} + \frac{64q^4}{5c} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^4 \\ & + \left[\frac{128(c-16)q^2}{315c^2} + \frac{32(c-24)q^3}{35c^2} + \frac{256(c-40)q^4}{105c^2} + O(q^5) \right] \left(\frac{\pi\ell}{L} \right)^6 \\ & + O(\ell^8) \end{aligned}$$

Long-interval cases

- For the long-interval case, we use the fact $S_{B,i} = S_{A,i}$, and the known result for thermal state EE:

$$S_B = \frac{c}{3} \log \left(\frac{\beta}{\pi \epsilon} \sinh \frac{\pi \ell}{\beta} \right) + \frac{\pi c L}{3\beta} - I(1 - \exp^{-\frac{2\pi \ell}{\beta}}), \quad \text{c.f. Chen \& Wu, 1412.0761, 1506.03206}$$

where $I(x)$ is the mutual information of two intervals on a complex plane with cross ratio x .

High T case:

NB: order c^0 correction!!

$$\chi_B = \frac{\pi c L}{3\beta} - \frac{2\pi^3(4\pi L - 7\beta)\ell^4}{315\beta^4 L} + \frac{32\pi^5 \ell^5}{3465\beta^5} + \frac{8\pi^4(32\pi^2 L^2 - 143\pi\beta L)\ell^6}{135135\beta^6 L^2} + \dots + O(\ell^{11}, 1/c).$$

Low T case:

$$\chi_B = S(L) - \left[\frac{32\pi\beta(\beta^2 + L^2)(4\beta^2 + L^2)}{15L^5} q^2 + O(q^3) \right] \left(\frac{\pi \ell}{L} \right)^4 + O(\ell^5).$$

Micro-canonical ensemble (MCE)

- We can also consider the Holevo information in the MCE with fixed high energy E , i.e., $p_i = \frac{\delta(E - E_i)}{\Omega(E)}$, with $\Omega(E) = \sqrt{\frac{\pi c L}{6E}} I_1\left(\sqrt{\frac{2\pi c L E}{3}}\right)$,
- We can obtain the multi-point functions/EEs in MCE by inverse Laplace transform of the ones in Canonical ensemble (CE), similar to the trick deriving the Cardy's formula for $\Omega(E)$.

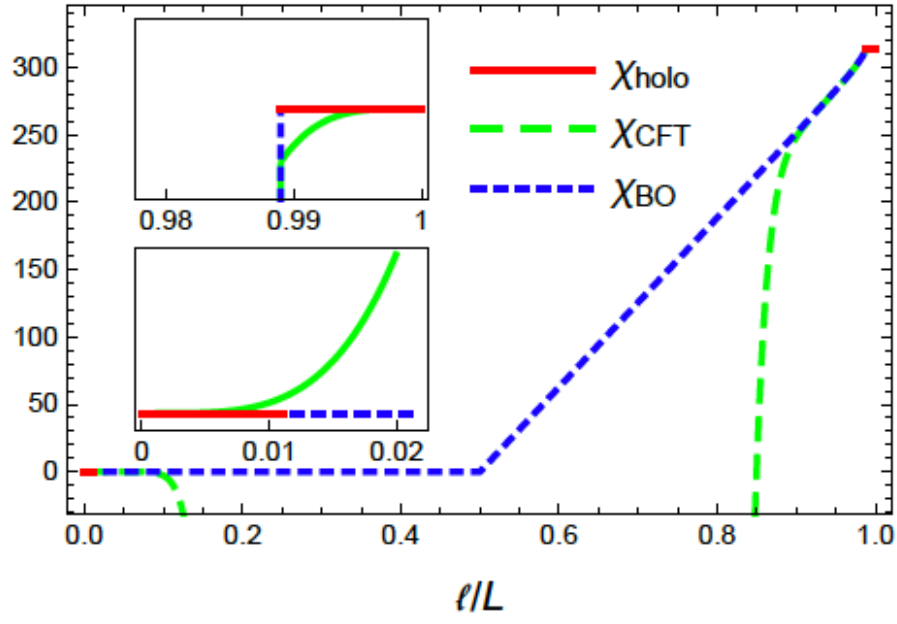
Short-interval: $\chi_A = \frac{\pi^3 \ell^4 [\pi c L (I_3 - I_1) + 24 \lambda I_2]}{540 \lambda^4 L I_1} + \dots + O(\ell^{12})$

NB: order c !!

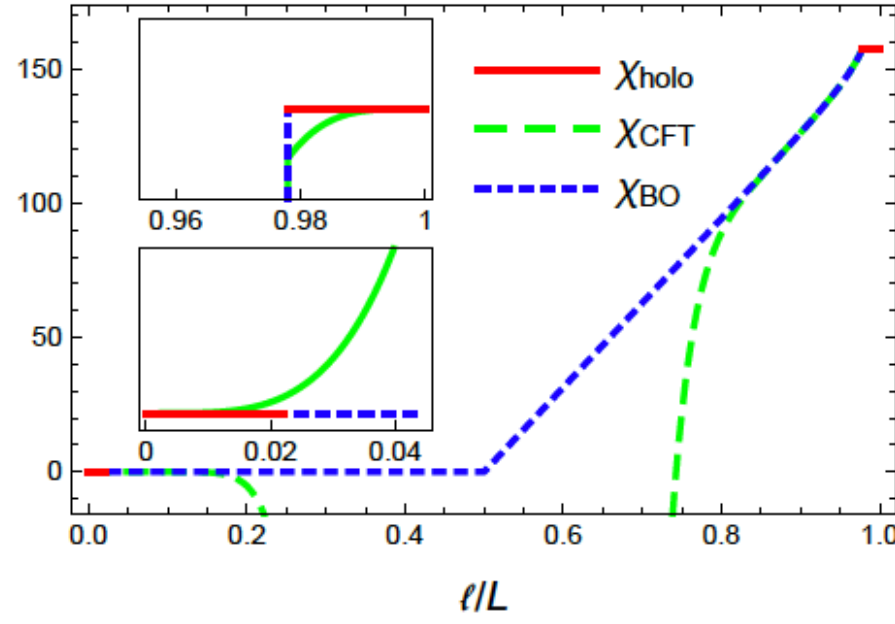
Long-interval: $\chi_B = S(L) + O(\ell^{12})$

There are also non-universal contribution from non-identity families.

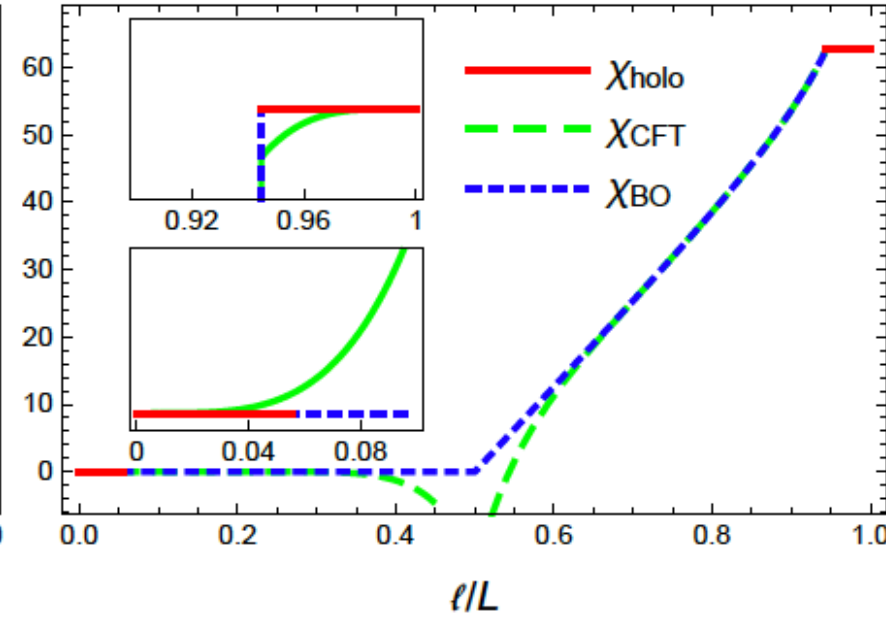
c=30 here!



(a) $\beta/L = 0.1$



(b) $\beta/L = 0.2$



(c) $\beta/L = 0.5$

$$\chi_{\text{holo}}(\ell) = \begin{cases} 0 & \ell < \frac{\beta}{2\pi} \log 2 \\ \frac{\pi c L}{3\beta} & \ell > L - \frac{\beta}{2\pi} \log 2 \end{cases}$$

$$\chi_A = \frac{2\pi^3 \ell^4}{45\beta^3 L} - \frac{8\pi^4 \ell^6 (\pi c L + 12\beta)}{945c\beta^5 L^2} + \dots + O(\ell^{12}).$$

$$\begin{aligned} \chi_B = & \frac{\pi c L}{3\beta} - \frac{2\pi^3 (4\pi L - 7\beta) \ell^4}{315\beta^4 L} + \frac{32\pi^5 \ell^5}{3465\beta^5} \\ & + \frac{8\pi^4 (32\pi^2 L^2 - 143\pi\beta L) \ell^6}{135135\beta^6 L^2} + \dots + O(\ell^{11}, 1/c). \end{aligned}$$

$$\chi_{\text{BO}}(\ell) = \begin{cases} 0 & \ell < L/2 \\ \frac{c}{3} \log \frac{\sinh \frac{\pi\ell}{\beta}}{\sinh \frac{\pi(L-\ell)}{\beta}} & L/2 < \ell < L - \frac{\beta}{2\pi} \log 2, \\ \frac{\pi c L}{3\beta} & \ell > L - \frac{\beta}{2\pi} \log 2 \end{cases}$$

Implications & Conclusions

- We see that the Holevo information got $1/c$ corrections but is consistently vanishing as interval size goes to zero.
- Recall that Holevo information=average of relative entropy over all states, including descendants.
- Since most of states are descendants (at least in the thermodynamic limit), many of which we shown are non-geometric and should contribute with higher order of c .
$$\frac{\Omega_p(E)}{\Omega(E)} \sim e^{-L \left[\sqrt{\frac{2\pi c \varepsilon}{3}} - \sqrt{\frac{2\pi(c-1)}{3} \left(\varepsilon + \frac{\pi}{6L^2} \right)} \right]} \rightarrow 0 \text{ as } L \rightarrow \infty.$$
- Thus, our results implies there is a subtle cancellation of non-geometric state's contributions in the average of relative entropy to yield order c Holevo information.
- This implies that even the non-perturbative quantum gravity effect looks wild, it is miraculously to yield averaging classical geometry of BTZ.