

Towards solving the hierarchy problem in string theory

A "small" progress report after 1st East Asia Joint WS at Hufei, 2016

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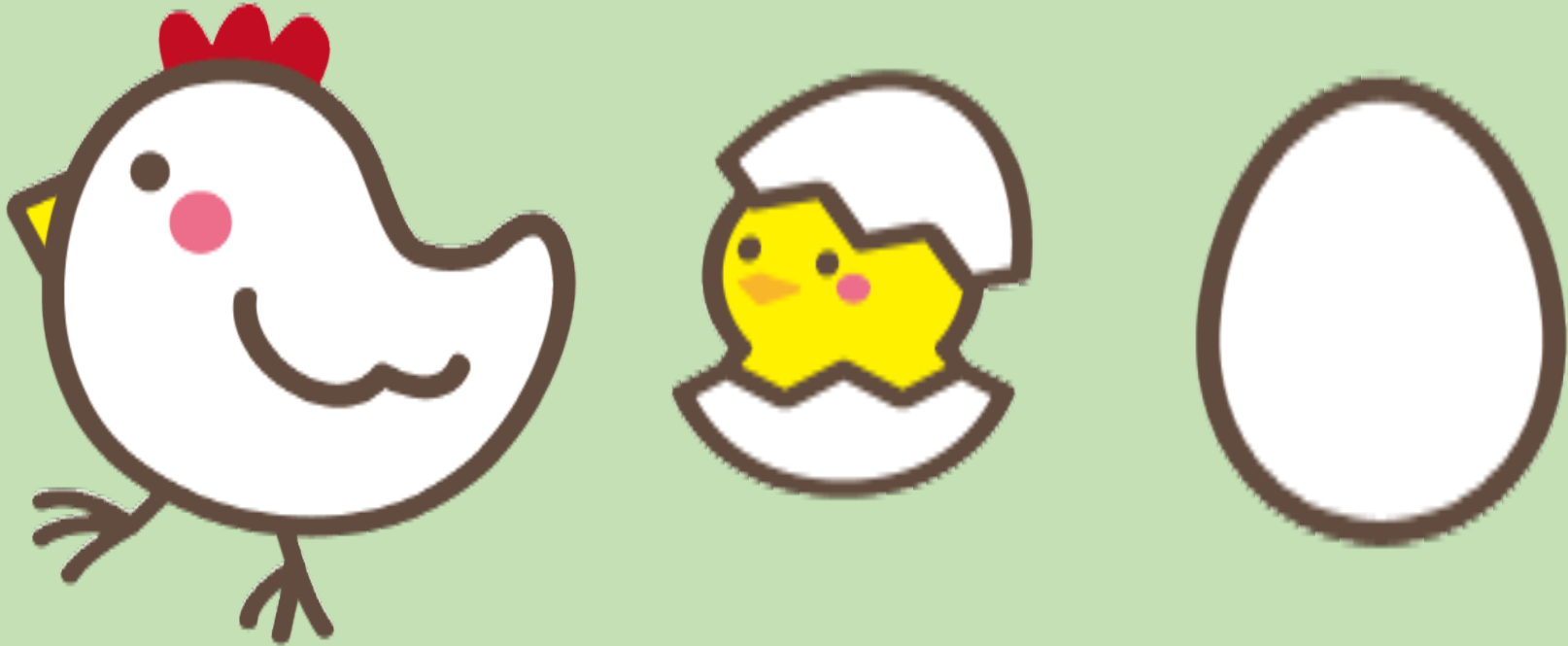
with

N. Kitazawa: 北澤敬章 (TMU) 키타자와 노리아키

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Which came first,
the chicken or the egg ?

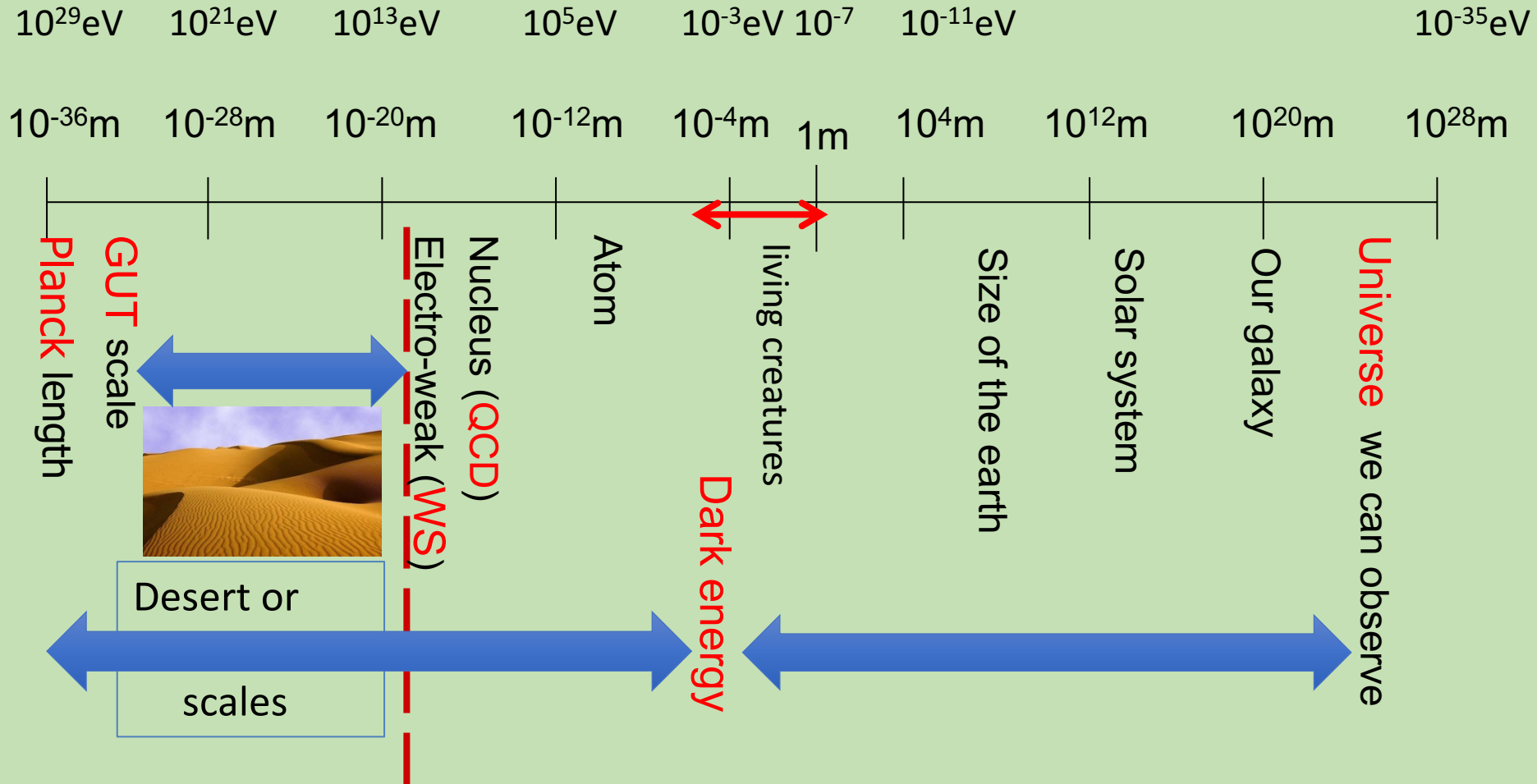


Genetic theory tells us that the egg came first
and many people may think so

Is this always the case ?

This is what I would like to discuss today.

Hierarchy of scales in nature



Naturalness problem = Hierarchy of various scales in nature

Today I will concentrate only on the EW scale.

Why **EW scale 100 GeV** is much lower than **UV scales**

GUT scale 10^{16} GeV, if exists

Planck scale 10^{18} GeV

String scale 10^{17} GeV ?

Higgs potential

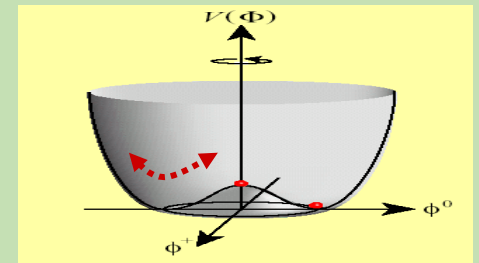
$$V = -\mu^2 |H|^2 + \lambda (|H|^2)^2$$

$$\delta V = \frac{1}{2} \int \frac{d^4 k}{(2\pi)^4} \text{Str} \log(k^2 + M(\phi^2))$$

$$= \boxed{\frac{\Lambda^2}{32\pi^2} \text{Str} M(\phi)^2} + \text{Str} \frac{M(\phi)^4}{64\pi^2} \ln \left(\frac{M^2}{\Lambda^2} - \frac{1}{2} \right)$$

quadratic divergence

$$\longrightarrow \text{Str} M(\phi)^2 = 0 \quad \text{SUSY ?}$$



Partial list of solutions to the hierarchy problem:

(1) Supersymmetry: cancellation of quadratic divergences
No SUSY particles are found, little hierarchy problem, ...

(2) Technicolor : dynamical generation of scales like QCD
Light Higgs difficult, big form factor (composite),

$$\Lambda_{TC} = M_{TC} \exp\left(-\frac{8\pi^2}{bg^2}\right)$$

Higgs is a light pseudo-NG ? $SO(4)/SO(3)$ etc.

(3) Multiverse / Anthropic ?

(4) Classical conformal : Coleman-Weinberg radiative breaking

$$V(\phi) \sim \frac{\lambda}{4} \phi^4 \left(\ln \frac{\phi^2}{M^2} - \frac{1}{2} \right) + V_0$$

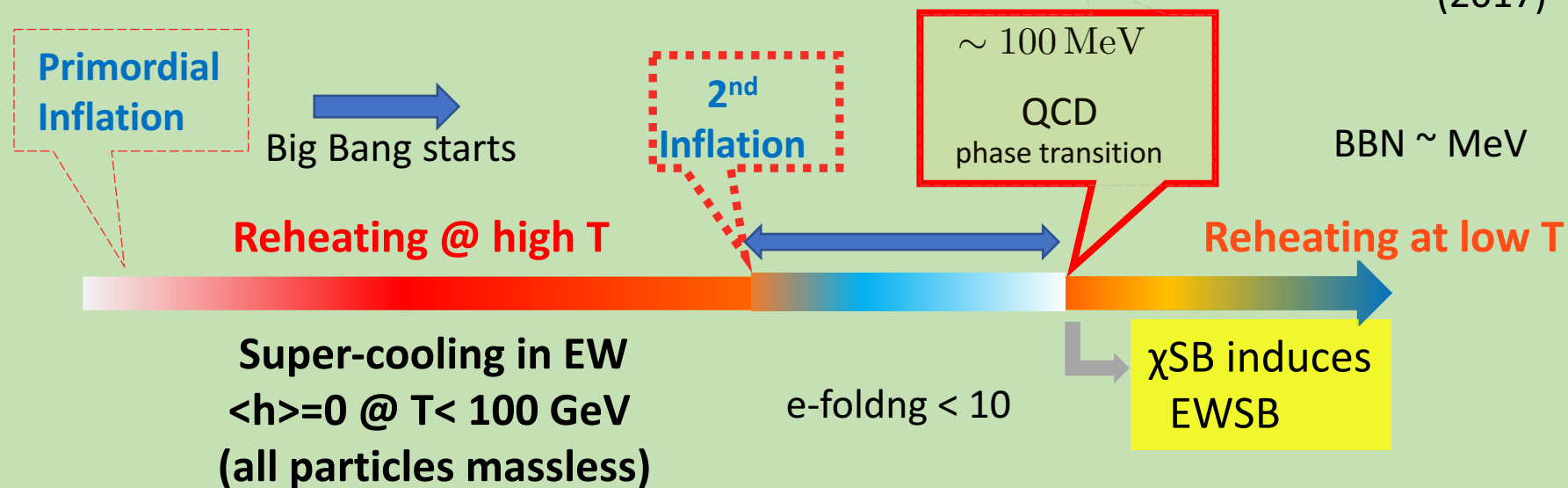
My favorite model

coupling to gravity \rightarrow quantum scale invariance ?

1-page summary of recent progress in the classical conformal pheno.

Super-cooled universe with the second (low scale) inflation in the classically conformal model

Serpico, Shimada, SI
(2017)



DM produced \rightarrow dilute \rightarrow "super-cool DM"
(an appropriate number of e-folding)

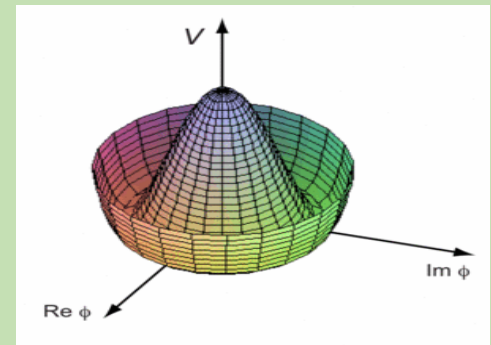
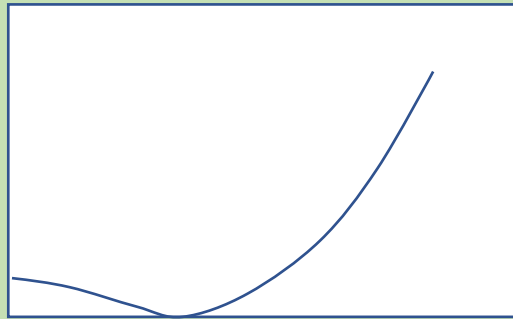
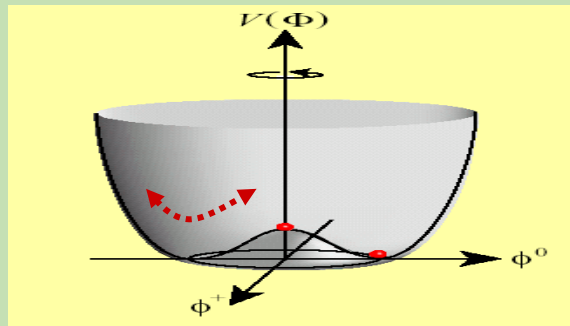
Thermal inflation starts at TeV scale and ends at 100 MeV !

gravitational wave, PBH, Baryogenesis

Anyway

There are many proposals to solve the hierarchy problem,
but there is one common basic assumption

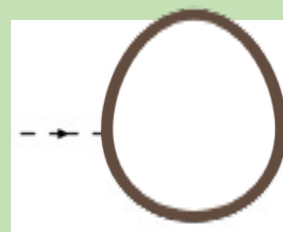
"Calculate the Higgs potential first !"



And **then obtain the solution** = minimum of the potential.

one solution to one Higgs potential

→ Also we are faced with the naturalness problem

A Feynman diagram representing a tadpole loop, consisting of a circle with two external lines. The diagram is shown with dashed lines on the left and right, and a solid line on the top and bottom.
$$= \frac{3y_t^2}{8\pi^2} \Lambda^2$$

Which came first,
the chicken or the egg ?



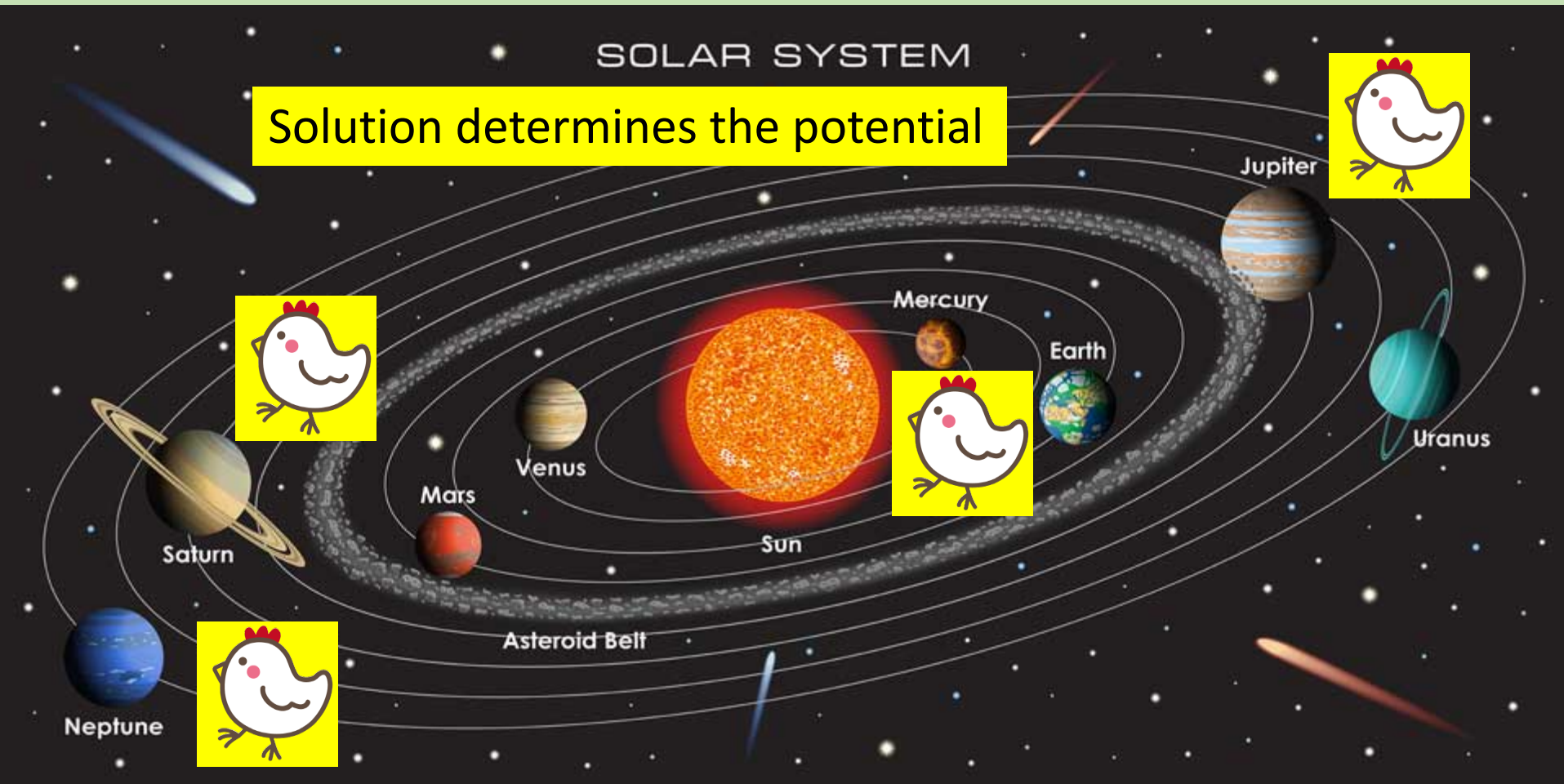
Usually calculate the potential first, then obtain a solution.



Is it possible to obtain a solution first , then calculate potential?



A classical example of the **Chicken first approach**

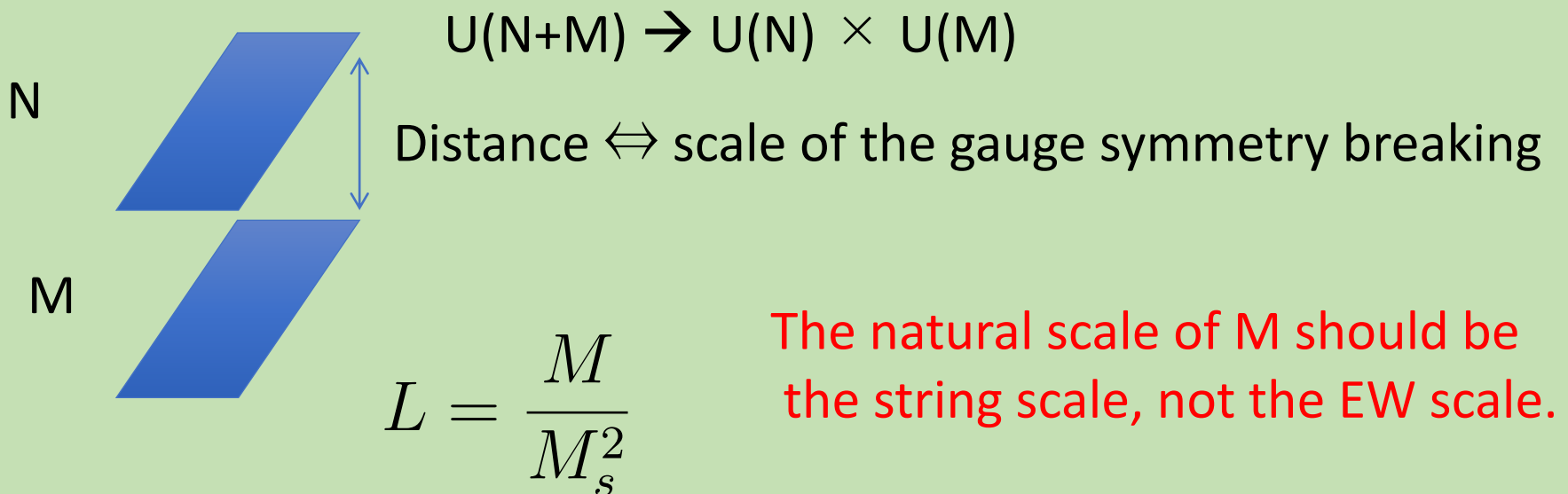


There are many orbits around the Sun: the Earth, the Jupiter, Mercury .. each of the orbit is at the bottom of the corresponding potential. But the underlying dynamics is the same.

A similar mechanism
to dynamically generate the EW scale.

I talked about the basic idea
at 1st East Asia Joint WS @ Huhei

In the D-brane model building,
the distance between branes (=moduli)
gives a vev of the scalar field.



The natural scale of M should be
the string scale, not the EW scale.

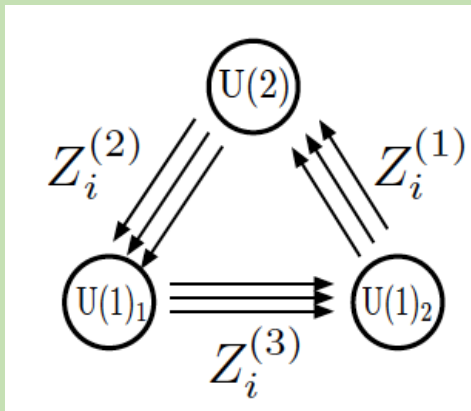
Hierarchy problem in string theory
= Difficulty to generate the EW scale in string theory

Ex : D3s+anti-D7 on Z_3 orbifold $R^4 \times (T^2 \times T^2 \times T^2)/Z_3$

Put 4 D3-branes on a fixed point of T^6/Z_3

Assignment of Z_3 charge for D3s

$$\gamma_3 = \text{diag}(\mathbf{1}_2, \alpha, \alpha^2)$$



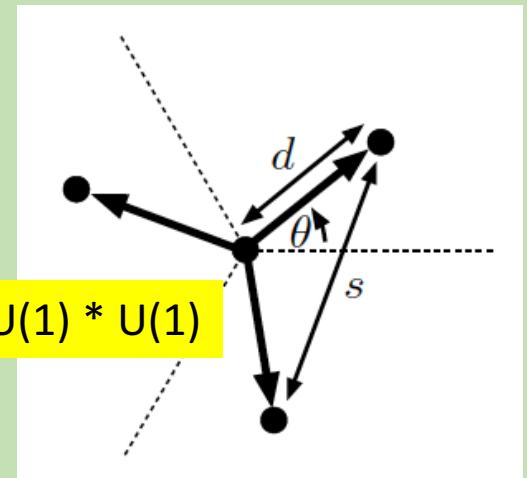
Quiver gauge theory

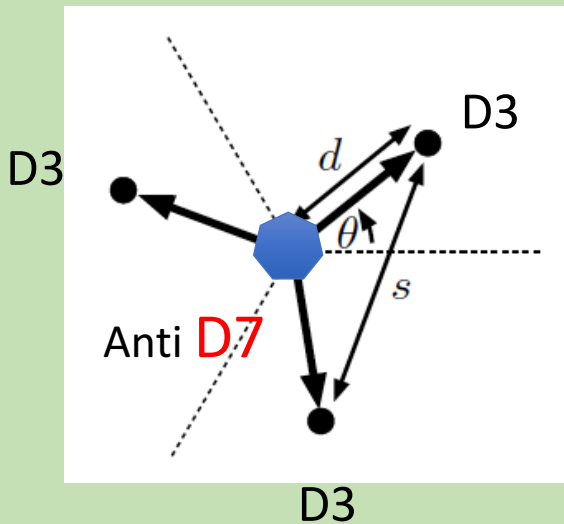
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$$U(2) * U(1) * U(1)$$

vev

$$U(1) * U(1)$$





Attractive force between D3s and anti-D7
due to open string 1-loop amplitudes

$$\mu^2 \sum_a |Z_3^{(a)}|^2$$

$$\mu^2 = \frac{1}{C^2} \frac{g^2}{16\pi^2} M_s^2$$

1-loop suppressed

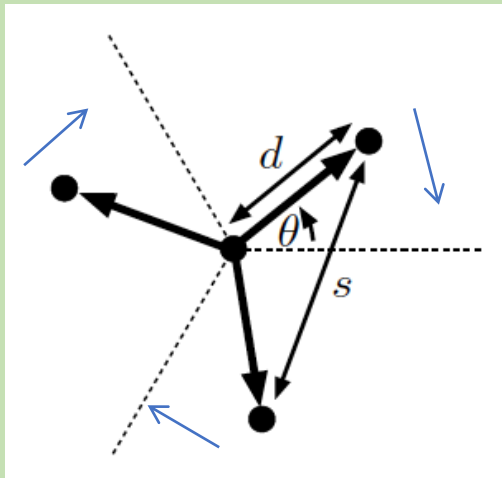
Repulsive centrifugal force
by revolution of D3s



Solution:
Hierarchy of EW scale

$$M \ll M_s$$

N.Kitazawa SI
PTEP,2015



High angular frequency $\omega = \mu \sim \frac{g}{4\pi} M_s$

Low velocity $v = \omega d \sim \frac{v_0}{M_s} \ll 1$

Merry-go-round scenario (named by Kimyeong)

It is possible to make a classically stable state with a short distance $r \ll l_{\text{string}}$.

But the large angular frequency
→ two problems

$$\omega = \mu \sim \frac{g}{4\pi} M_s$$

- Dispersion relation of Higgs violates Lorentz symmetry (Coriolis force)

$$\omega^2 = \begin{cases} 2(p^2 + m^2) \\ \frac{\lambda v^2}{2\omega_0^2} p^2 \end{cases}$$

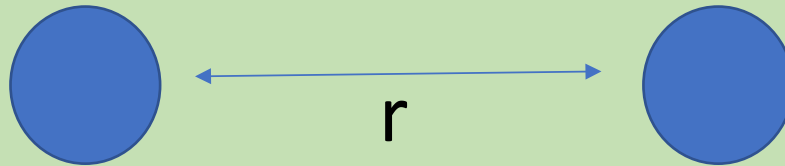
- closed string emission → unstable

To avoid these problems,
 it is necessary to make a bound (or resonant)
 state with $\omega \ll m_{str}$.

→ We need **weak attractive force**: $V \ll (m_{string})^3 r^2$

A simple way to **avoid** $\omega = m_{str}$ is to consider

Flat moduli such as $Dp - Dp$,



BPS = **no interaction at rest**
 but

v-dependent attractive force is generated
 when they are moving with a constant velocity

In the closed string region,

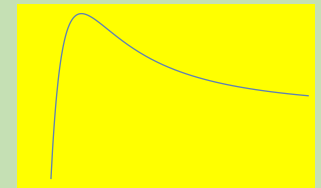
$$r \gg l_{str} \quad \frac{-Mv^4}{r^{7-p}} + \frac{J^2}{Mr^2}$$



No minimum exists.

Potential barrier at $r < l_{str}$

closed string picture



In the closed-string dominated region ($r \gg l_{\text{str}}$),
the repulsive force surpasses the attractive force.

However, "Beauty is Attractive"

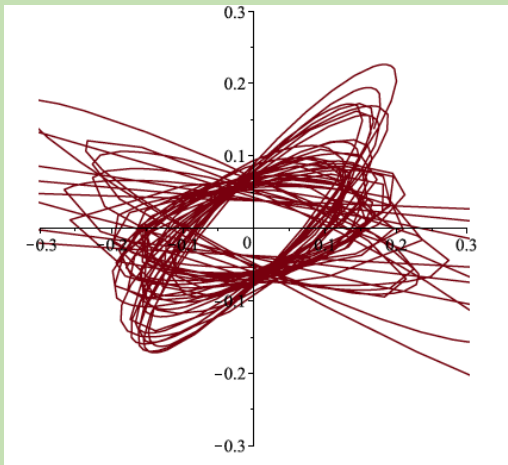
Kofman Linde Liu Malony
MaClister Silverstein 2004



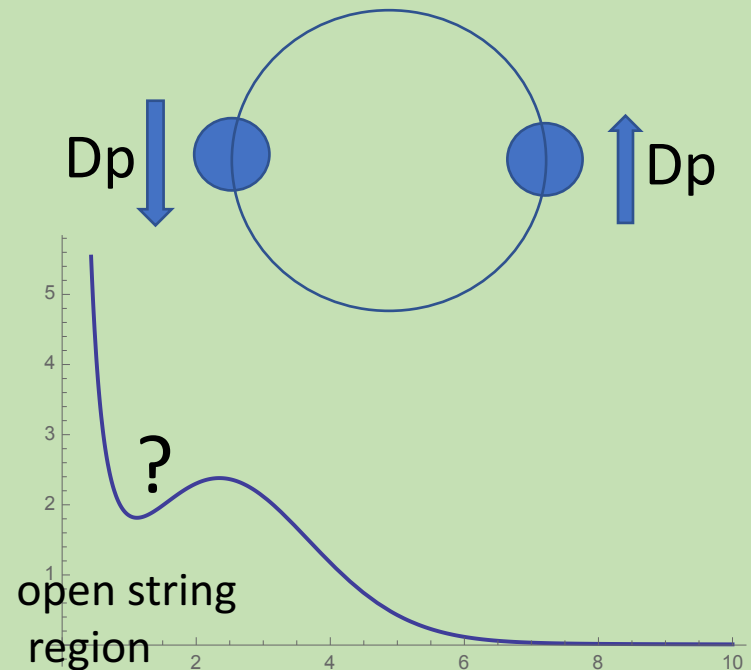
Mass changes rapidly. \rightarrow non-adiabatic process
Open string particles are produced (= preheating).

$$n_{\chi} = \frac{(gv)^{3/2}}{(2\pi)^3} e^{-\pi g\mu^2/v}.$$

- \rightarrow lose energy and the trajectory shrinks.
- \rightarrow and when the trajectory becomes circular, no more particle production occurs.



Is it possible to make
 a "bound state" with $r \ll l_{\text{str}}$
 between revolving D-branes?



Many works on D0-branes (N=16 SU(N) supersymmetric QM)

Witten index (Piljin Yi) $\text{Tr}(-1)^F = 1$ ---- threshold bound state
 massless graviton

Quantum bound state (Kabat Pouliot) ---- resonant state $E > 0$

These states are near the ground state.

What we want is highly excited, but almost stable resonant states.
 (like a solar system, not like a hydrogen atom)

Poor man's Calculation of attractive potential between revolving Dp, in particular p=0.

T. Suyama,
H. Ohta, SI
to appear soon

$$\begin{aligned}x_1(t) &= r \cos \omega t, \\y_1(t) &= r \sin \omega t,\end{aligned}$$



$$\begin{aligned}x_2(t) &= -r \cos \omega t, \\y_2(t) &= -r \sin \omega t.\end{aligned}$$

It is straightforward, but
not so trivial
because of the **time-dependent**
boundary conditions:

In the rotational frame, the boundary condition becomes simple,
but the system is interacting:

$$\begin{aligned}S = & -\frac{1}{4\pi\alpha'} \int d^2\sigma \left[-\partial_\alpha \tilde{T} \partial^\alpha \tilde{T} + \partial_\alpha \tilde{X} \partial^\alpha \tilde{X} + \partial_\alpha \tilde{Y} \partial^\alpha \tilde{Y} + \partial_\alpha \tilde{X}^i \partial^\alpha \tilde{X}_i \right. \\& \left. + 2\omega \partial_\alpha \tilde{T} (\tilde{X} \partial^\alpha \tilde{Y} - \tilde{Y} \partial^\alpha \tilde{X}) + \omega^2 (\tilde{X}^2 + \tilde{Y}^2) \partial_\alpha \tilde{T} \partial^\alpha \tilde{T} \right].\end{aligned}$$

Furthermore, by taking variation with respect to the T fields,

$$\delta S \Big|_{bdy} = -\frac{r^2}{2\pi\alpha'} \delta \tilde{T} \left[-\partial_\sigma \tilde{T} + v(\tilde{X} \partial_\sigma \tilde{Y} - \tilde{Y} \partial_\sigma \tilde{X}) + v^2(\tilde{X}^2 + \tilde{Y}^2) \partial_\sigma \tilde{T} \right] \Big|_{bdy}$$

Thus T – field must satisfy \rightarrow at the boundaries:

$$\begin{aligned} (1 - v^2) \partial_\sigma \tilde{T} - v \partial_\sigma \tilde{Y} &= 0. & \sigma = 0 \\ (1 - v^2) \partial_\sigma \tilde{T} + v \partial_\sigma \tilde{Y} &= 0. & \sigma = \pi \end{aligned}$$

These conditions can be simplified by introducing a new variable

$$T := \tilde{T} - \frac{v}{1 - v^2} x(\sigma) \tilde{Y}$$

With these (and a few more) changes of world sheet fields, the action of open strings stretched between revolving D0s becomes

Open string world sheet action between revolving D0s

$$\begin{aligned}
 S = & -\frac{r^2}{4\pi\alpha'} \int d^2\sigma \left[-\dot{X}^2 - \dot{Y}^2 - (\dot{X}^i)^2 + \left(X' - \frac{2}{\pi}\right)^2 + (Y')^2 + (X^{i'})^2 \right. \\
 & + \left[1 - v^2 \left((X + x(\sigma))^2 + Y^2 \right) \right] \left[\left(\dot{T} + \frac{vx(\sigma)}{1-v^2} \dot{Y} \right)^2 - \left(T' + \frac{v}{1-v^2} (x(\sigma)Y' - \frac{2}{\pi}Y) \right)^2 \right] \\
 & - 2v \left(\dot{T} + \frac{vx(\sigma)}{1-v^2} \dot{Y} \right) \left((X + x(\sigma))\dot{Y} - Y\dot{X} \right) \\
 & + 2v \left(T' + \frac{v}{1-v^2} (x(\sigma)Y' - \frac{2}{\pi}Y) \right) \left((X + x(\sigma))Y' - Y \left(X' - \frac{2}{\pi}\right) \right) \Big].
 \end{aligned}$$

boundary conditions

$$X|_{\sigma=0,\pi} = 0, \quad Y|_{\sigma=0,\pi} = 0, \quad \partial_\sigma T|_{\sigma=0,\pi} = 0, \quad X_i|_{\sigma=0,\pi} = 0$$

One-loop amplitude of open string between revolving D0s
perturbation with respect to the angular velocity ω
up to ω^2

$$\begin{aligned} Z &= \int_0^\infty \frac{dt}{2t} \text{Tr} \left[e^{-2\pi t (H_{rot} - \frac{1}{8})} \right] (\eta(it))^{-21} \\ &= \int_0^\infty \frac{dt}{2t} \text{Tr} \left[e^{-2\pi t (H_{rot} - 1)} \right] \prod_{m=1}^{\infty} (1 - e^{-2\pi m t})^{-21} \end{aligned}$$

The assumptions of our calculation are

$$r \ll l_{string}$$



low lying open string
states dominate

$$v = r\omega \ll 1$$



perturbation is valid

Effective potential induced by **massive states**

: Double expansion with respect to w^2 and r^2 .

$$\begin{aligned}\mathcal{V}_2 &\sim m_{str} \left(c_1 + c_2 \left(\frac{w}{m_{str}} \right)^2 + \dots \right) \left(\frac{r}{l_{str}} \right)^2 \\ &+ m_{str} \left(c_3 + c_4 \left(\frac{w}{m_{str}} \right)^2 + \dots \right) \left(\frac{r}{l_{str}} \right)^4 + \dots\end{aligned}$$

Bosonic case: the first excited massive state

$$\begin{aligned}\mathcal{V}_2(r, \omega) &= \frac{1296 + 106\alpha'\omega^2}{8\sqrt{\alpha'}} + \frac{1296 + (-730 + 432\pi^2 + 1296\pi^2\epsilon_0)\alpha'\omega^2}{16\sqrt{\alpha'}} \left(\frac{r^2}{\pi^2\alpha'} \right) \\ &- \frac{1296 + (-2430 - 864\pi^2 + 2592\pi^2\epsilon_0)\alpha'\omega^2}{64\sqrt{\alpha'}} \left(\frac{r^2}{\pi^2\alpha'} \right)^2 + \mathcal{O}(\omega^4, r^6).\end{aligned}$$

In superstring

$$c_1 = c_3 = 0$$

$$c_2 = c_4 = 0 ?$$

→ We are checking how it deviates from constant velocity case

Effective potential induced by **massless states** stretched bet. D0s

$$m_{str} \sqrt{\left(\frac{r}{l_{str}}\right)^2 + C_1 \left(\frac{\omega}{m_{str}}\right)^2 + C_2 \left(\frac{r}{l_{str}}\right)^2 \left(\frac{\omega}{m_{str}}\right)^2}$$

mass² acquired by the effect of revolution

c.f. Field theory calculation

$$\begin{aligned} -\text{Tr} \log(\Delta + m^2)^{-1/2} &= \int_{\epsilon}^{\infty} \frac{dt}{2t} \int \frac{d^{p+1}k}{(2\pi)^{p+1}} e^{-(k^2 + m^2)t} \\ &\sim \int_{\epsilon}^{\infty} \frac{dt}{2t} t^{-\frac{p+1}{2}} e^{-m^2 t} \sim \begin{cases} (m^2)^{\frac{p+1}{2}} & p = \text{odd} \\ (m^2)^{\frac{p+1}{2}} \log m^2 & p = \text{even}. \end{cases} \end{aligned}$$

Note that C_1 is negative for vectors (positive for KK scalar) which indicates that the system is unstable for large ω .

Closed string emission with spins is suppressed for $\omega \ll m_{\text{string}}$

angular momentum is conserved \rightarrow r and ω are related:

$$J = Mr^2\omega = \frac{V}{gl_{\text{str}}^p} \left(\frac{r}{l_{\text{str}}} \right)^2 \left(\frac{\omega}{m_{\text{str}}} \right)$$

The total potential per unit volume for revolving Dp-brane is

$$\frac{m_{\text{str}}}{2g} \left(\frac{r}{l_{\text{str}}} \right)^2 \left(\frac{\omega}{m_{\text{str}}} \right)^2 + \text{the following attractive potential}$$

Dp-Dp potential (p=odd)

$$\sum_i (-1)^{F_i} n_i \left(\left(\frac{r}{l_{\text{str}}} \right)^2 + C_i \left(\frac{\omega}{m_{\text{str}}} \right)^2 \right)^{\frac{p+1}{2}} \log \left(\left(\frac{r}{l_{\text{str}}} \right)^2 + C_i \left(\frac{\omega}{m_{\text{str}}} \right)^2 \right)$$

$$\begin{aligned} \mathcal{V}_2 &\sim m_{\text{str}} \left(c_1 + c_2 \left(\frac{\omega}{m_{\text{str}}} \right)^2 + \dots \right) \left(\frac{r}{l_{\text{str}}} \right)^2 \\ &+ m_{\text{str}} \left(c_3 + c_4 \left(\frac{\omega}{m_{\text{str}}} \right)^2 + \dots \right) \left(\frac{r}{l_{\text{str}}} \right)^4 + \dots \end{aligned}$$

The potential has a minimum around

$$\frac{r}{l_{str}} \sim \frac{\omega}{m_{str}} \ll 1$$

consistent with the assumptions
of our calculations

Total angular momentum

$$J = \frac{V}{gl_{str}^p} \left(\frac{r}{l_{str}} \right)^2 \left(\frac{\omega}{m_{str}} \right)$$

For $p=0$, $J \ll 1$

→ only s-wave (ground state) is allowed.

Thus only threshold bound state exists.

For $p>0$ J can be larger than 1.

A resonant state with higher J may exist.

Summary "The chicken-first approach" to hierarchy problem

Stationary solutions of D-branes

If resonant states exist,

it must be a hierarchical solution

$$\frac{r}{l_{str}} \sim \frac{\omega}{m_{str}} \ll 1$$

which means $E_{EW} \ll m_{string}$.

We have calculated the effective potential
corresponding to the solution.

For $p > 0$, there may exist a "classical" bound state satisfying

$$\omega \ll m_{string} \quad \text{Lorentz violation, instability by radiation} \quad \rightarrow \quad \text{OK}$$

Future issues:

- (1) Smarter calculation by D-brane EFT in t-dep background
- (2) construct phenomenologically realistic models
- (3) SUSY breaking: TeV SUSY? solution to little hierarchy?

관심을 가져 주셔서 감사합니다