

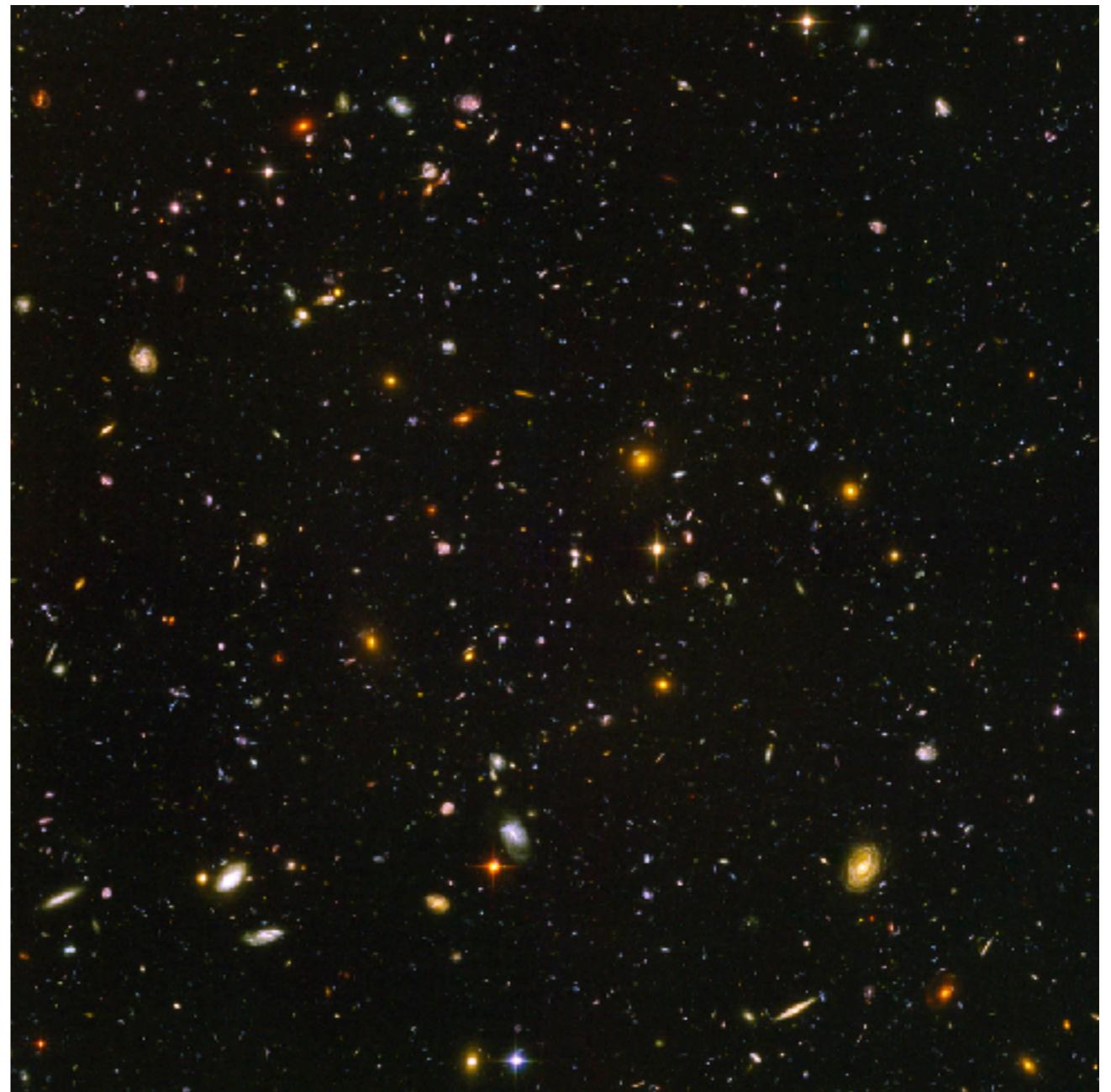
# Disentangling the role of halo mass, density and cosmic web filaments in shaping galaxy growth at high redshift

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# What shapes the diversity of galaxies?

*Halo mass, local density, ...*



*... large-scale environment: clusters, anisotropic structures, ...*

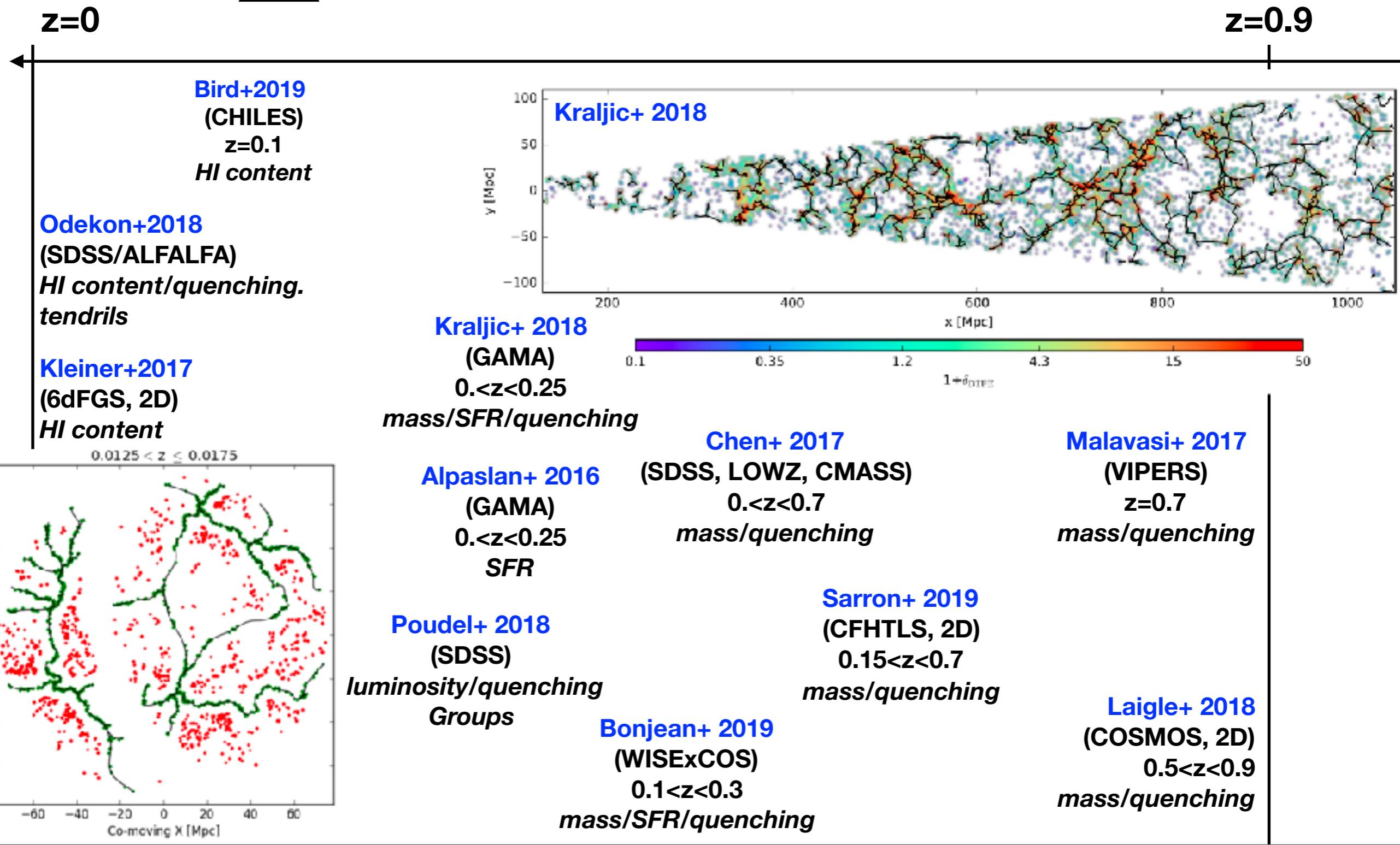
# Properties of galaxies in cosmic filaments from z=0 to z=1

## A review

(c) C. Laigle



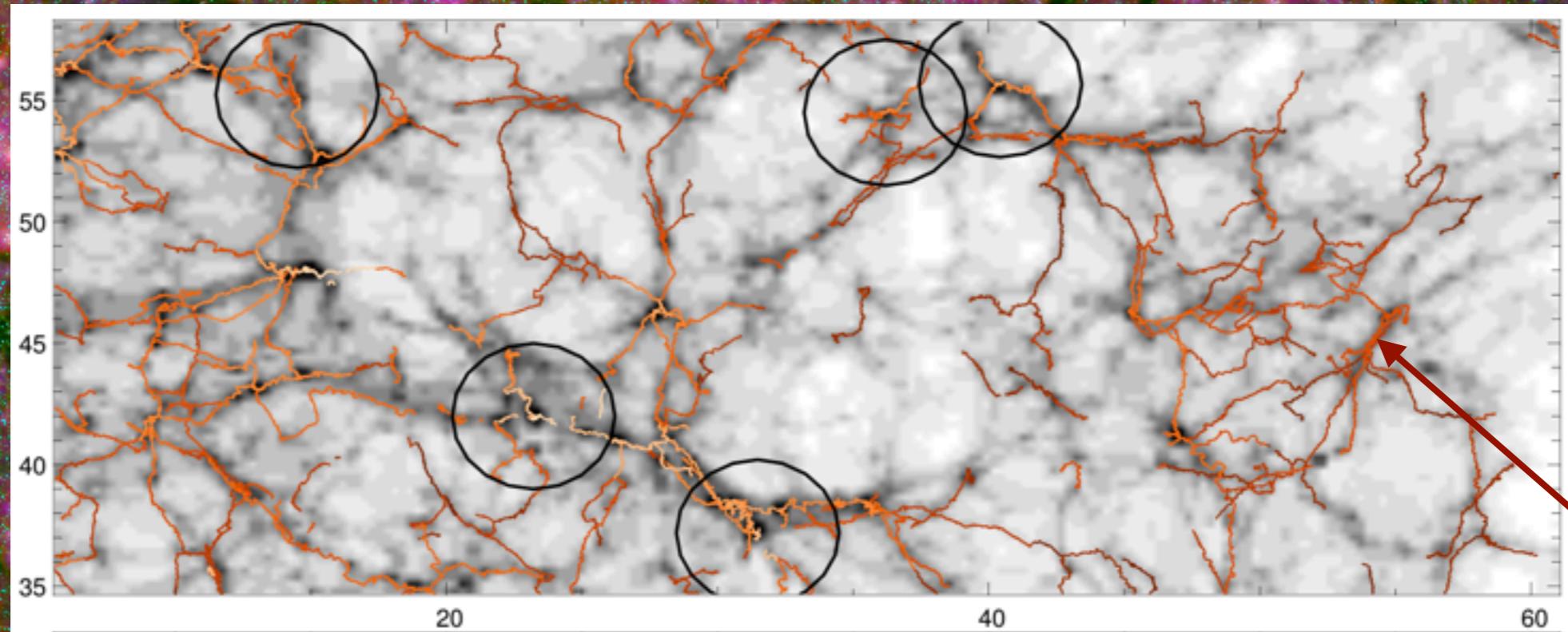
Different scales/filaments extractors/mass ranges ...



# Horizon-AGN

Redshift 2 ~ cosmic noon

100Mpc/h

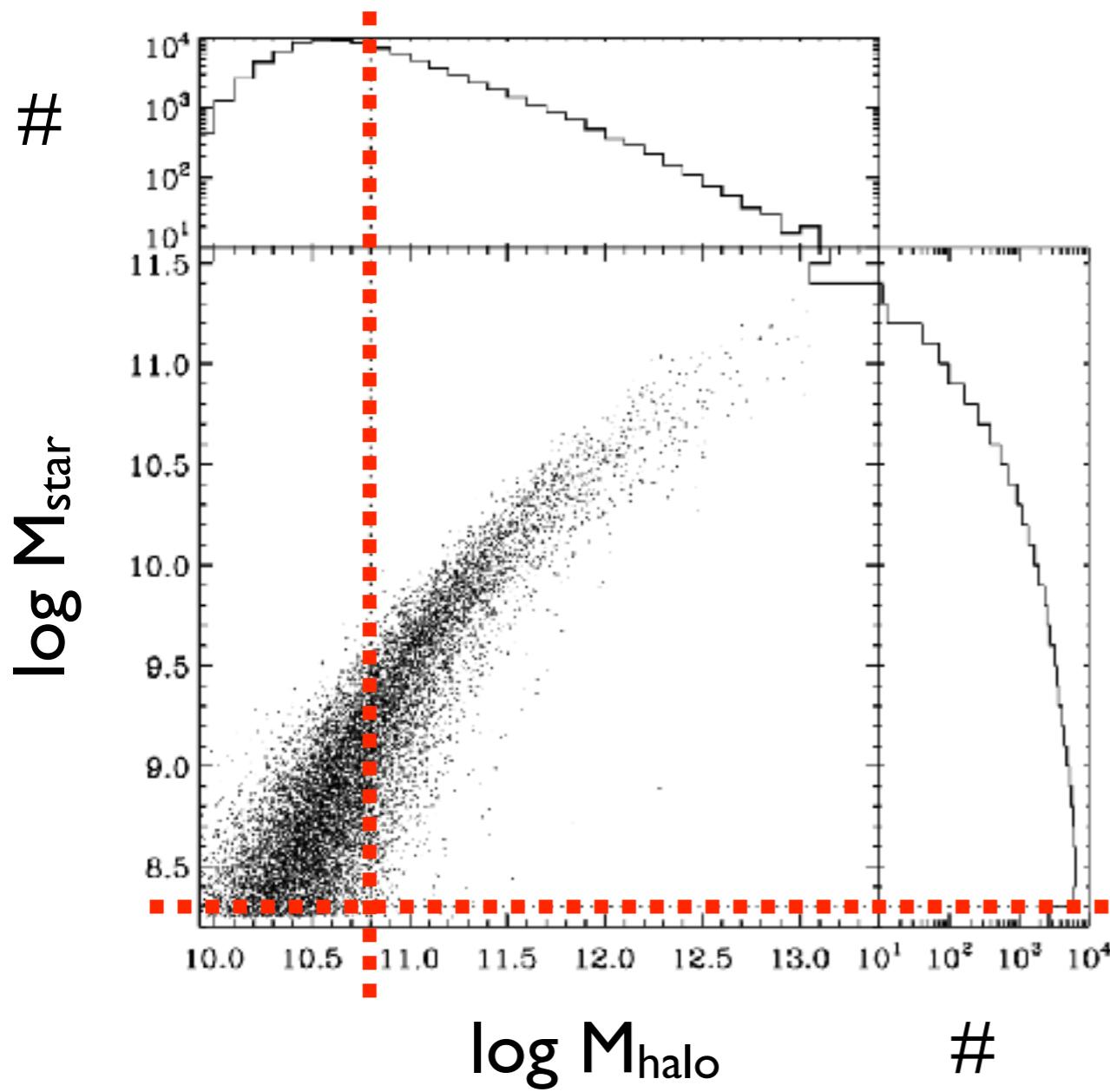


Dubois et al. (2014)

Filaments  
extracted  
using  
Disperse  
(Soubie  
2011)

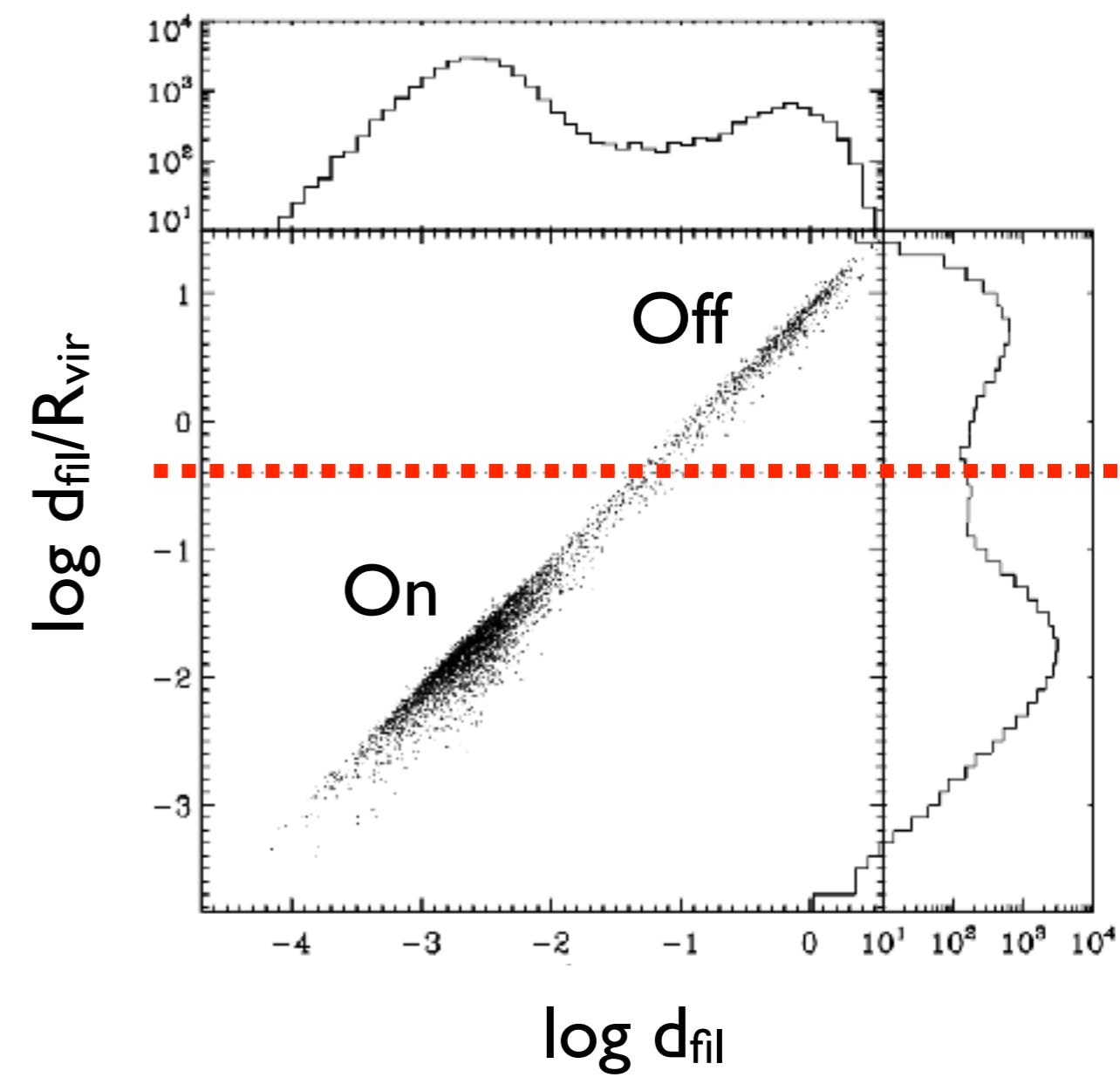
# Data

Mass-complete sample  
of central galaxies



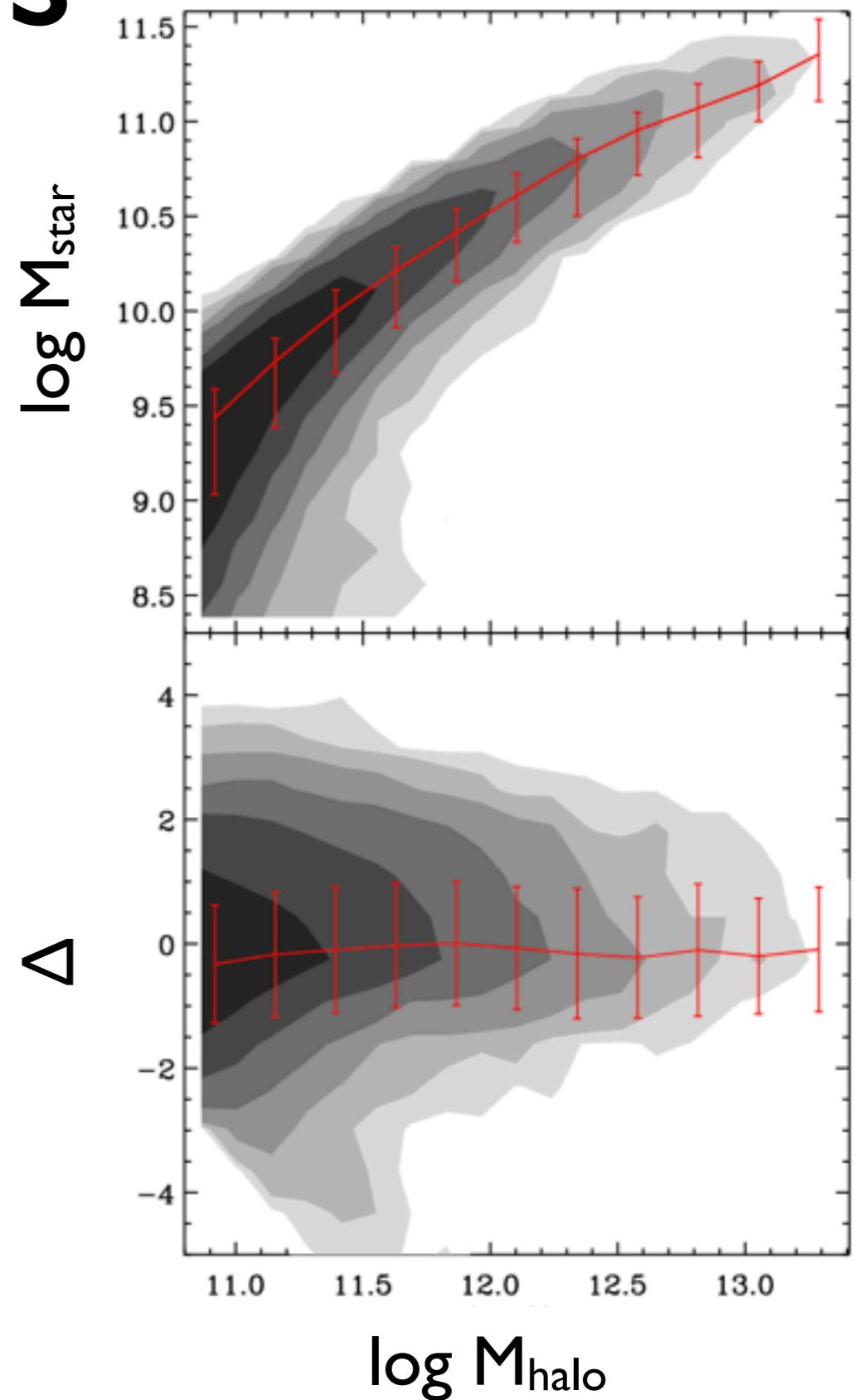
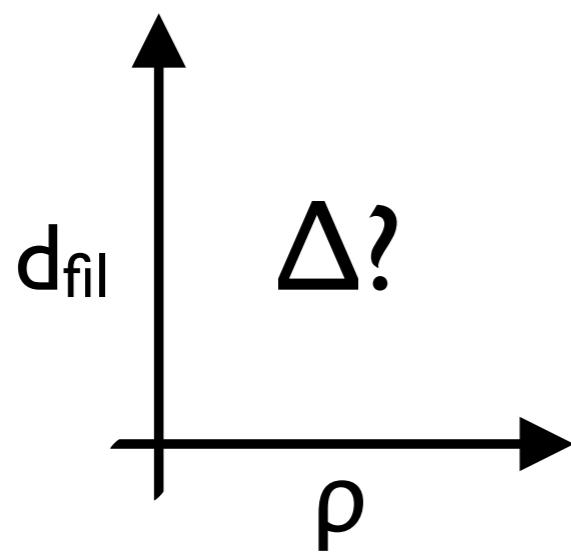
\*cluster galaxies excluded

On- & off-filament  
subsamples

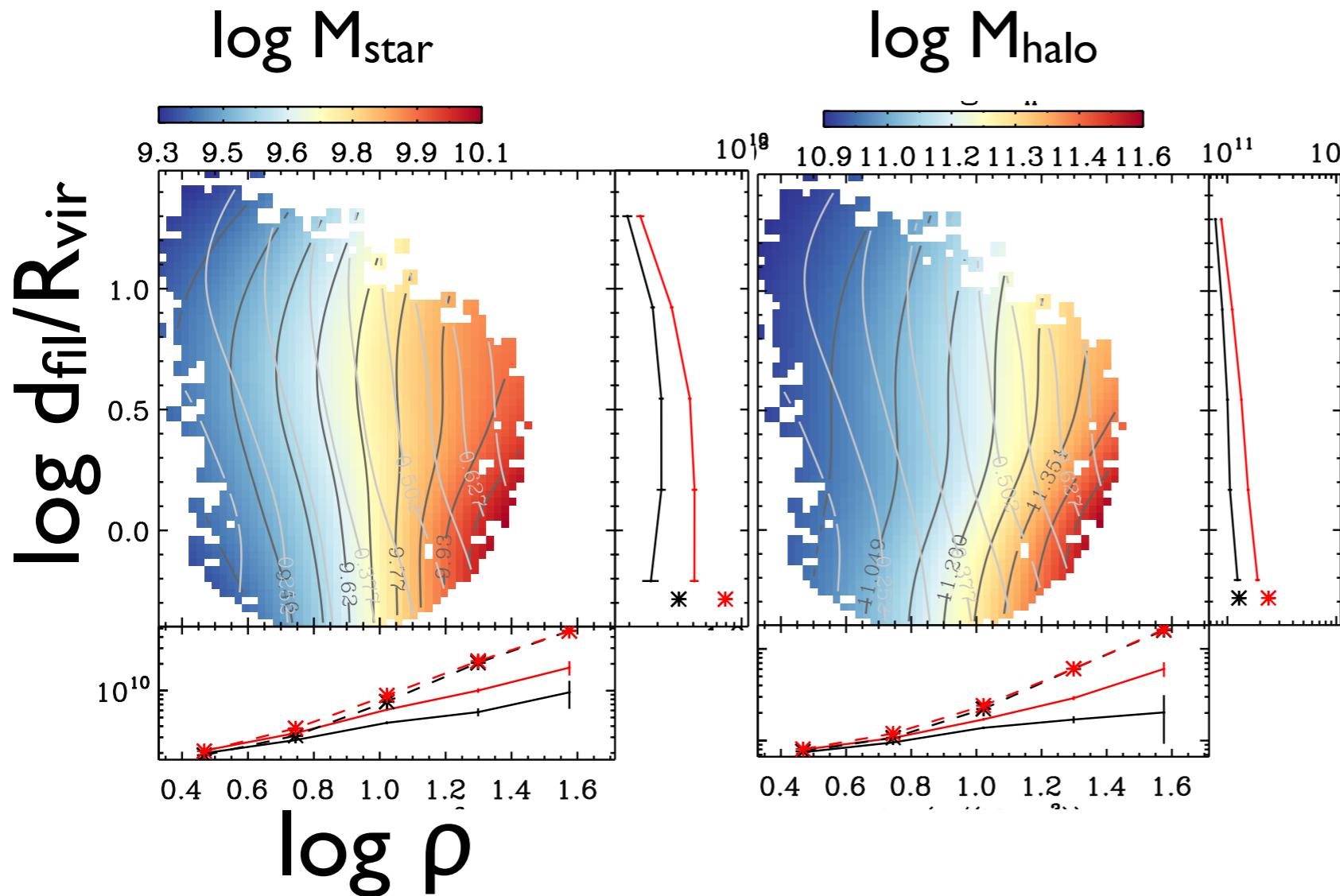


# Analysis

- Galaxy property  
 $= f( M_{\text{halo}}, \rho, d_{\text{fil}} )$
- Residual ( $\Delta$ )  
= offset from the mean trend  
at a given halo mass  
/scatter in the halo mass bin



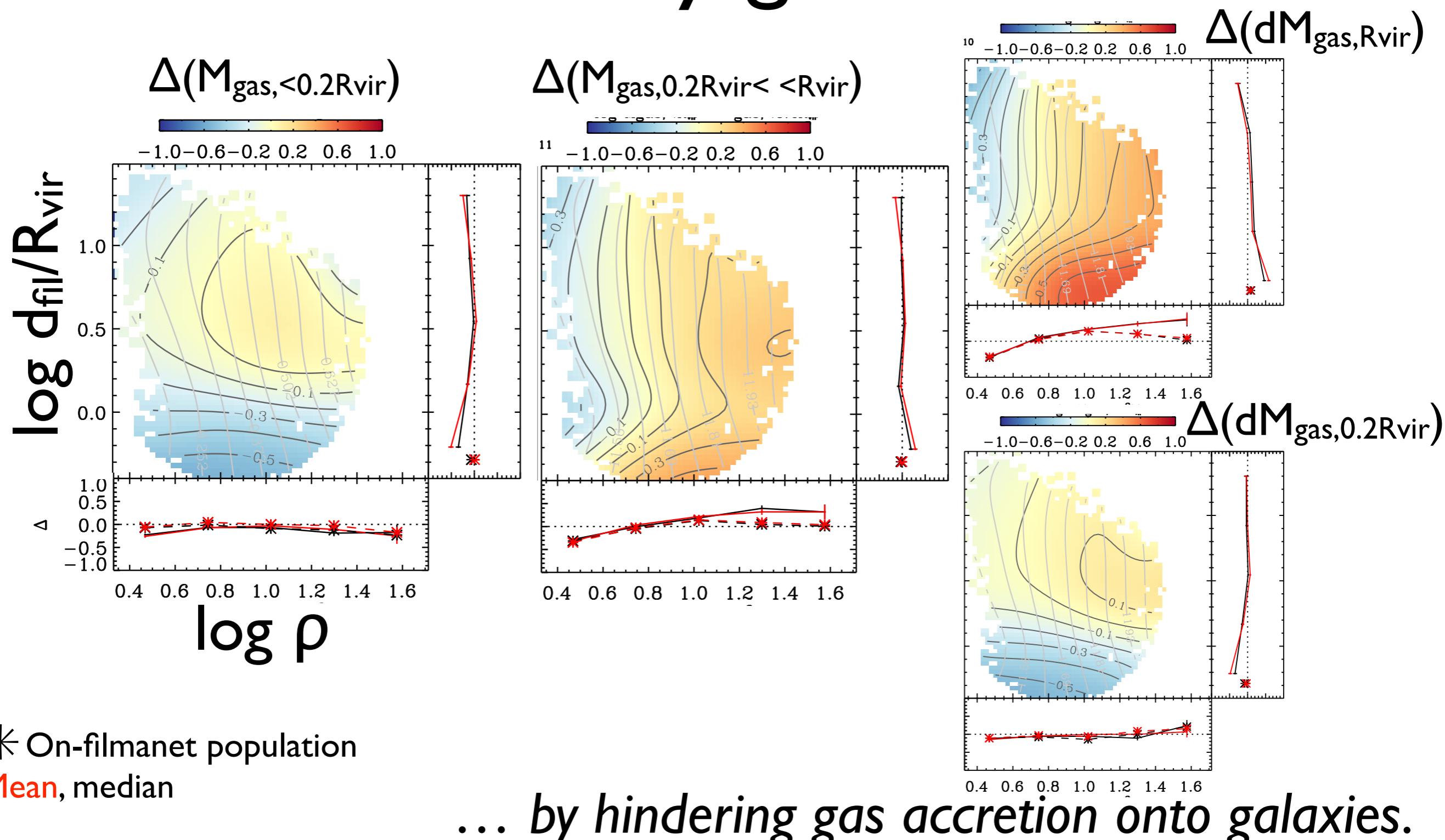
# What is the role of filaments in shaping galaxy growth?



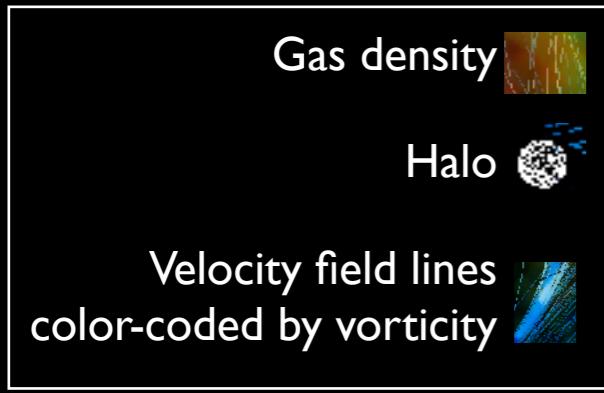
... Proximity to filaments enhances SF up to a certain distance,  
 \* On-filament population but suppresses SF until galaxies enter filaments.

Mean, median

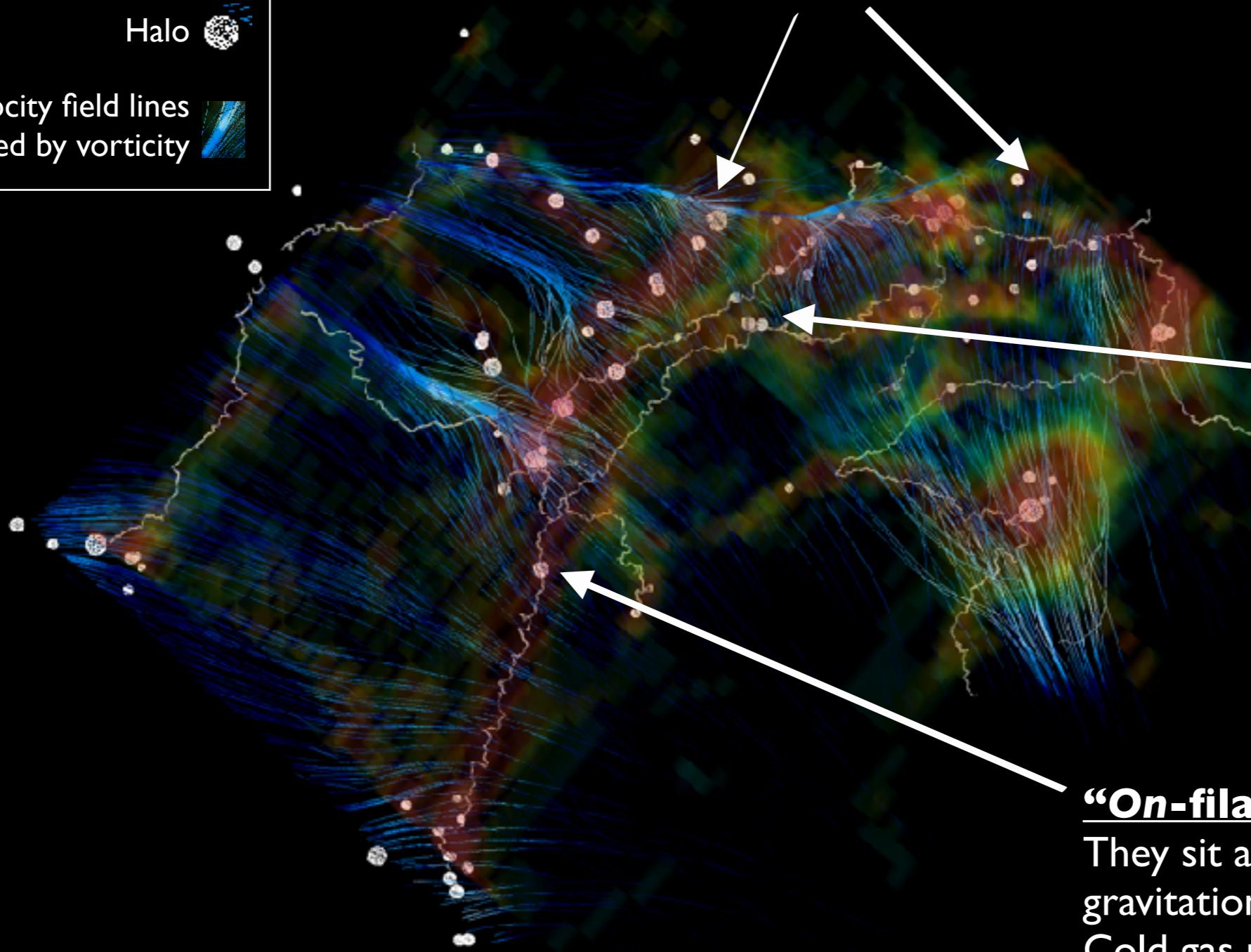
# Filaments suppress star formation of nearby galaxies



# Summary



**“Off-filament” halos far from filaments:**  
Star-forming cold gas efficiently injected into galaxies via laminar “tendrils”



**“Off-filament” halos close to filaments:**  
Although gas is on overall denser, accretion onto the galaxy is truncated (increasing anisotropic shear/flow redirection/gas shock heating)

**“On-filament” halos:**  
They sit at the bottom of gravitational potential well. Cold gas might be more easily funneled again through galaxies.