

Supernova Cosmology, Luminosity Evolution, & Dark Energy

The key assumption made in the discovery of
Dark Energy is in serious error

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Team YONSEI

(YOnsei N earby S upernovae E volution I nvestigation: since 2010)

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Spectroscopy
of host-galaxies



Young-Lo Kim
(Yonsei/CNRS Lyon)

Light-curve analysis
of SNe Ia



Chul Chung
(Yonsei)

Population synthesis
models (YEPS)

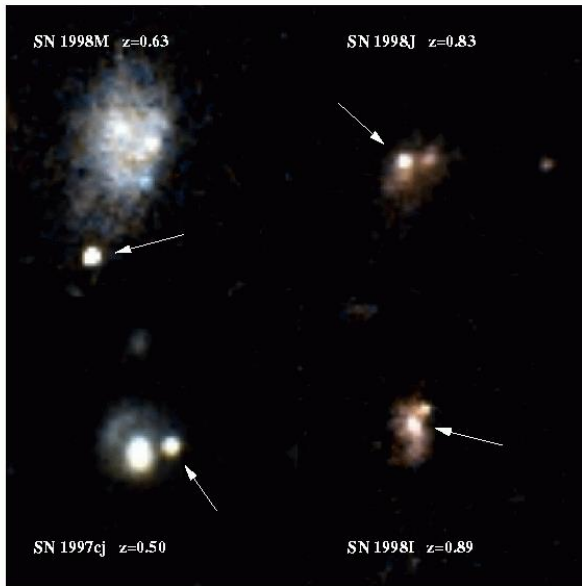
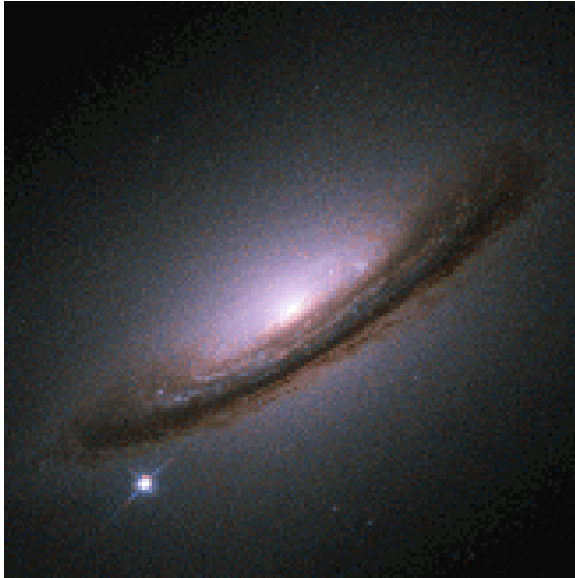


Dongwook Lim
(Yonsei/Heidelberg)

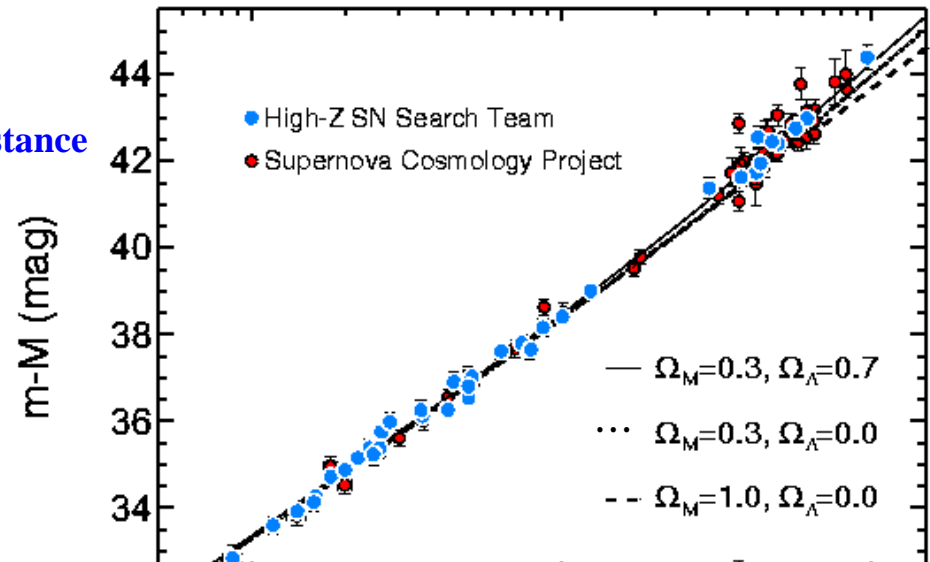
Chang Hee Ree
(KASI)

Spectroscopy,
Photometric analysis

Most direct evidence for Dark Energy is still from SN cosmology



Distance

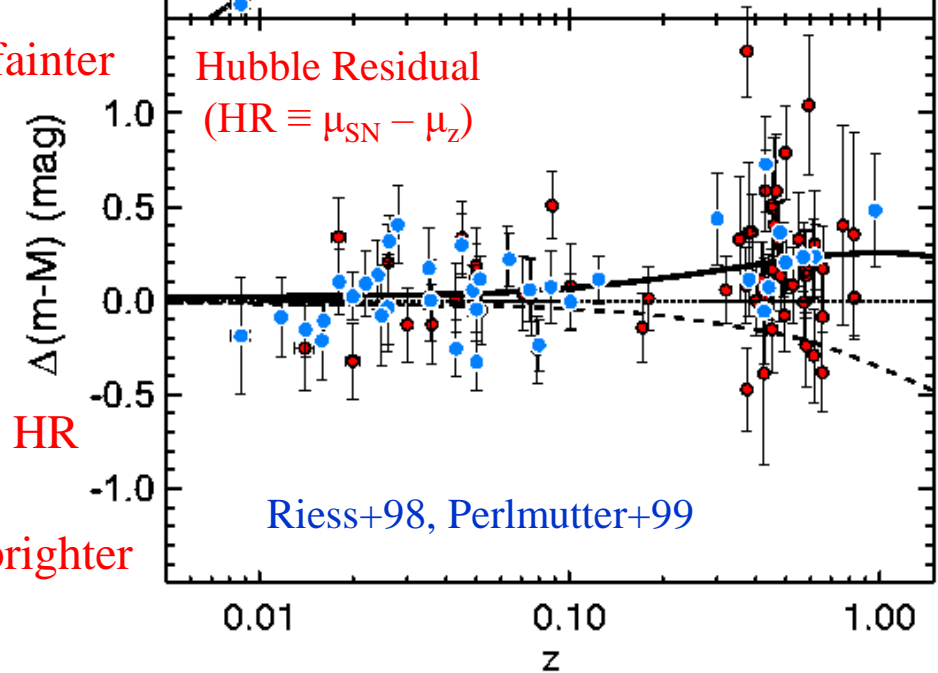


fainter

$\Delta(m-M)$ (mag)

brighter

Hubble Residual
($HR \equiv \mu_{SN} - \mu_z$)



Riess+98, Perlmutter+99

Redshift \rightarrow

Project YONSEI:

Yonsei Nearby Supernovae Evolution Investigation (Since 2010)

High S/N (~ 175) spectra for ~ 70 nearby early-type host galaxies

→ 34 nights with LCO 2.5m & MMT 6.5m long-slit spectroscopy

→ Only ETGs because of **high-precision age dating** & low extinction

→ For the first time, **directly** measure population age & metallicity

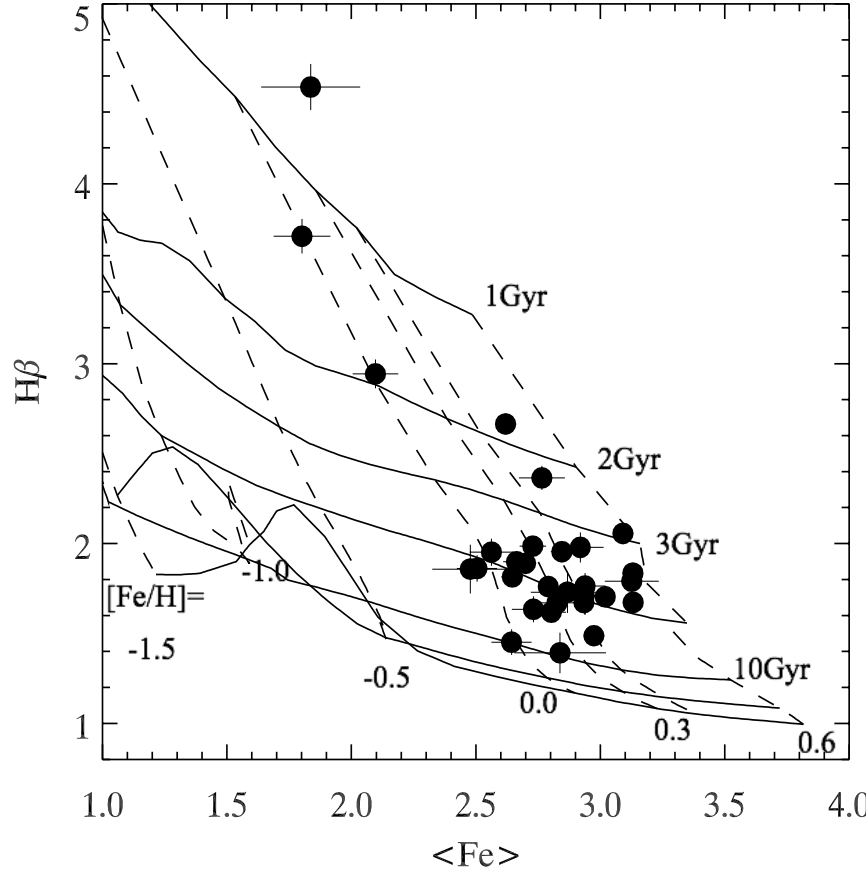
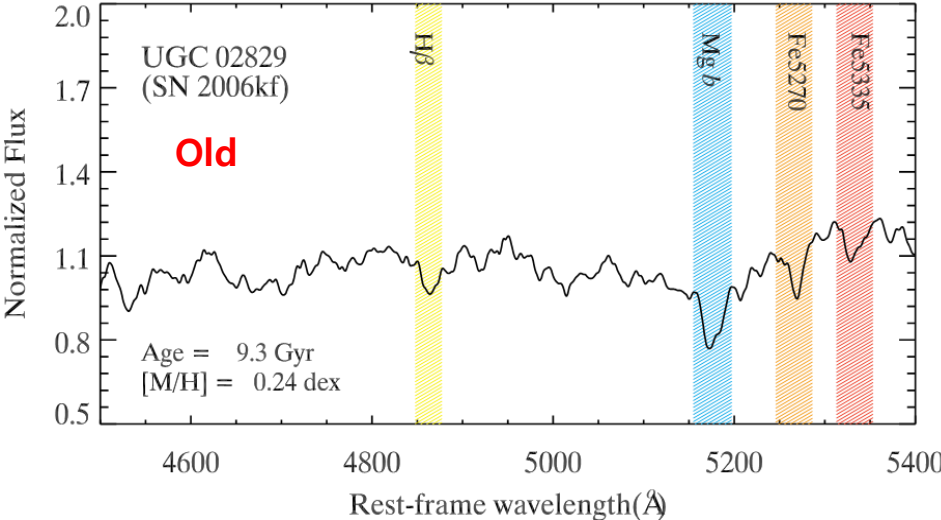
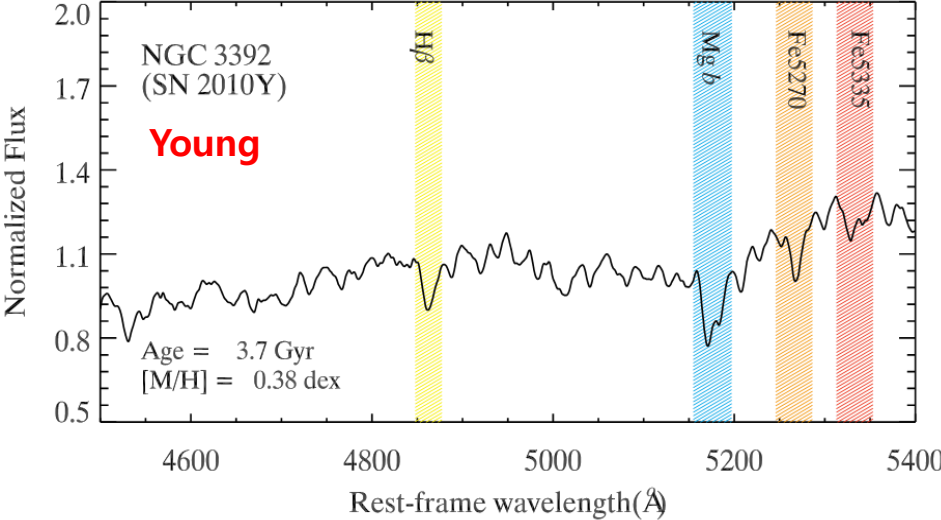
from absorption lines (H_{β} , Fe, Mgb) & well-established

population synthesis models (Yonsei, Thomas, Schiavon)

→ SALT2/SNANA SN LC analysis & standardization (Y.-L. Kim+2018)

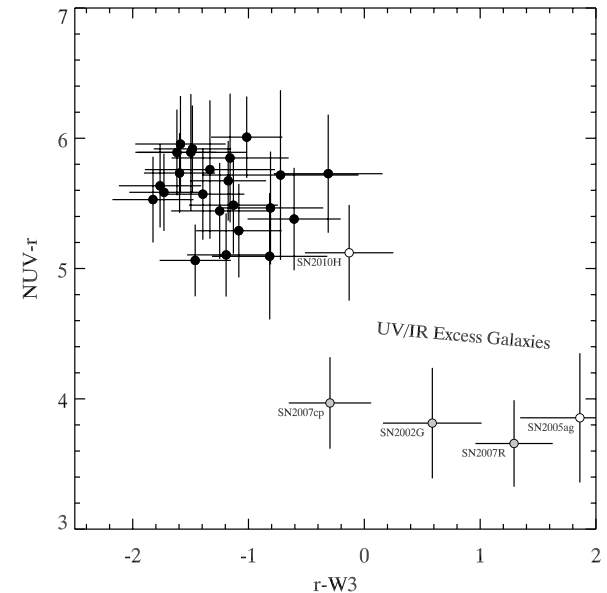
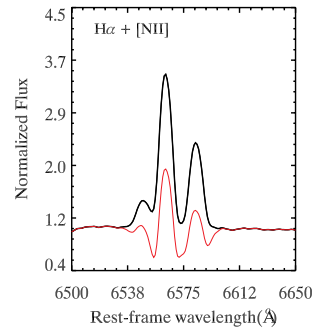
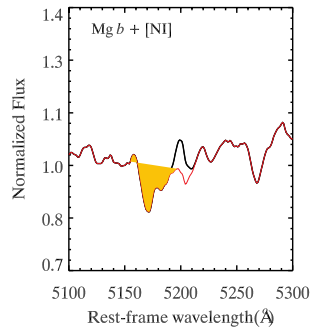
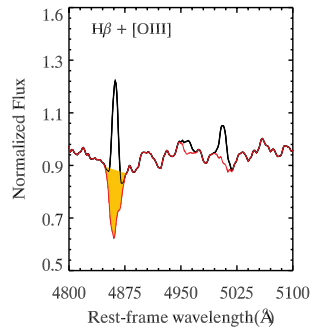
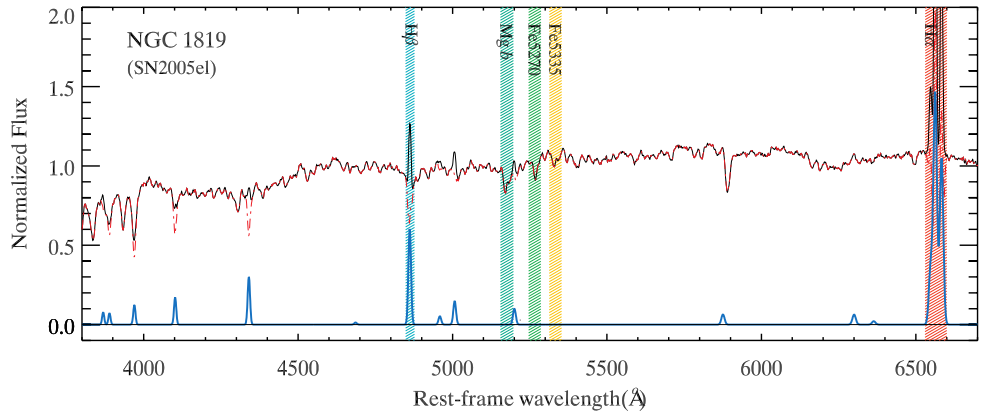
Most direct & stringent test ever made for the luminosity evolution!

High Precision Measurement of Host Galaxy Age & Metallicity

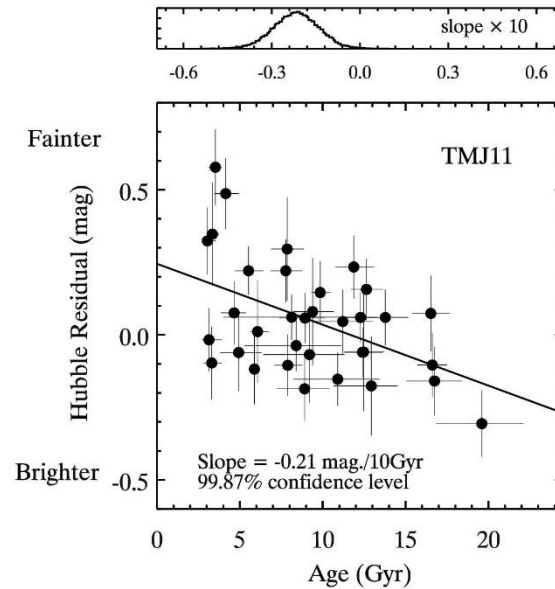
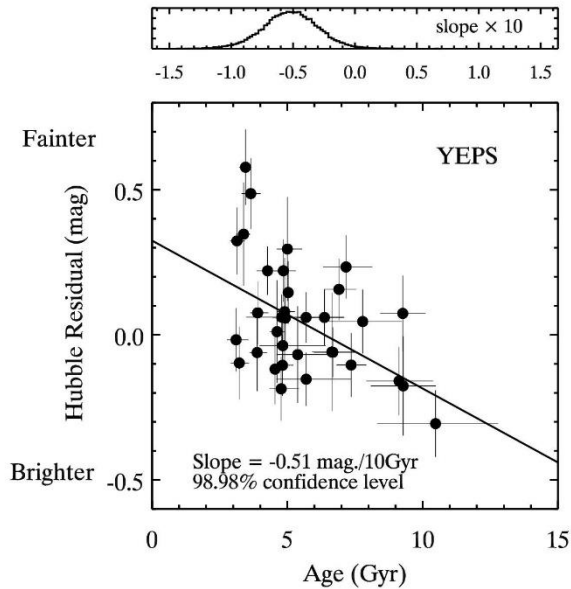


Population Synthesis Models:
Chung+13; Thomas+11; Schiavon 07

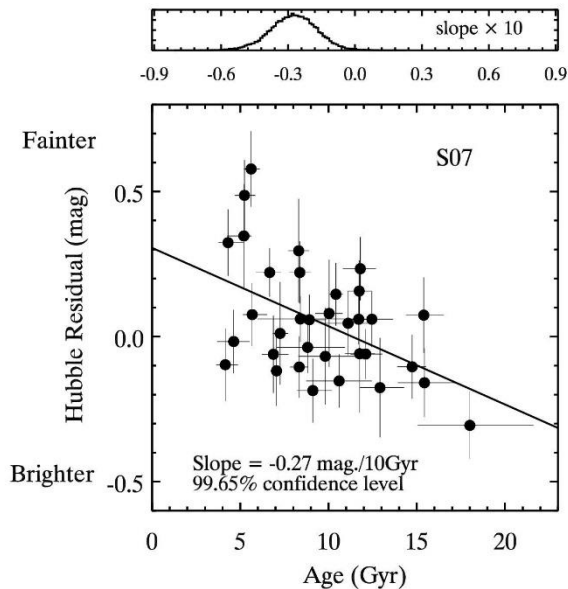
Abnormal (recent star-forming) ETGs were removed based on emission lines and/or UV/IR excess



Correlation between SN Luminosity & Population Age



Non-genuine ETGs with recent SF excluded as ages are seriously underestimated or highly uncertain



Correlation is significant at ~99.5% confidence
→ SNe in younger hosts are fainter (after standardization)

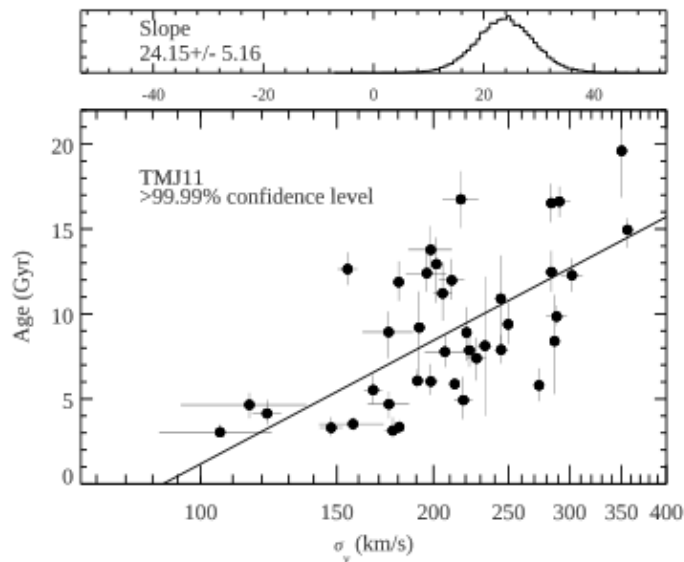
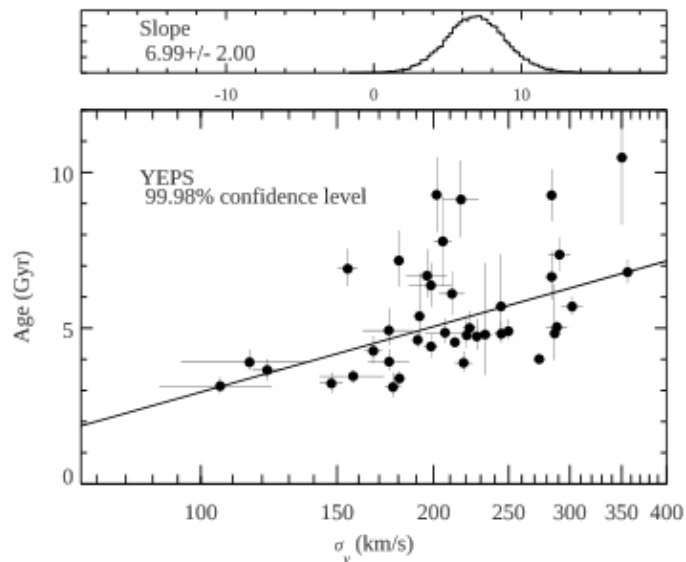
→ Evidence for luminosity evolution!

→ ~0.27 mag fainter for $\Delta t \sim 5.3$ Gyr

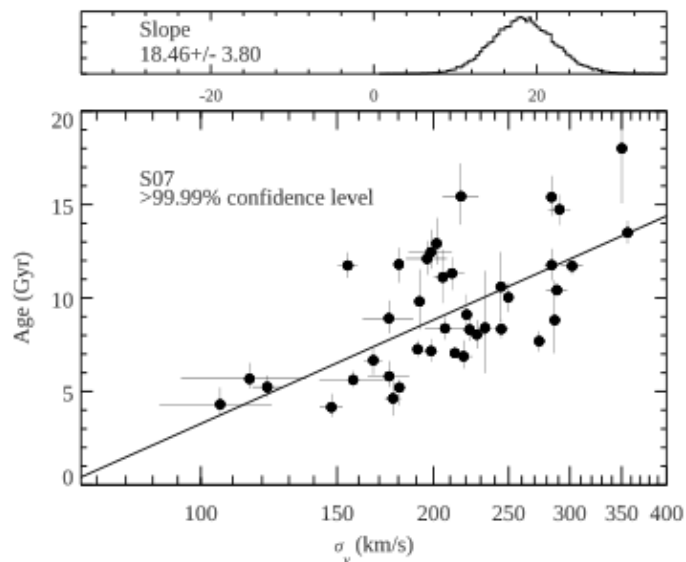
Our result is insensitive to the choice of population synthesis model

Correlation between Host Mass (σ_v) and Population Age: "Downsizing"

(Kang, Lee+2016, 2020)



Similar results from LTGs
(van de Sanda+2018)



MCMC analysis

→ Correlation is significant at 99.99% confidence level!

→ No apparent correlation with metallicity

→ *Host mass – SN luminosity correlation is due to population age difference!*

Other well-established correlations between SN luminosity & host property (related with age)

Host Property	Reference	Original Correlation	Direction	Converted to Age difference
Morphology	Hicken et al. (2009)	$\Delta HR / \Delta \text{morph.}$	Fainter in	$\sim 0.19 \text{ mag} / 5.3 \text{ Gyr}$
		$\approx 0.14 \text{ mag} / (\text{Scd/Irr-E/S0})$	Later type galaxy	Fainter in Younger galaxy
Mass	Sullivan et al. (2010)	$\Delta HR / \Delta \text{mass}$	Fainter in	$\sim 0.21 \text{ mag} / 5.3 \text{ Gyr}$
		$\approx 0.08 \text{ mag} / (\Delta \log M_{\star} \sim 1)$	Less massive galaxy	Fainter in Younger galaxy
Local SFR	Rigault et al. (2018)	$\Delta HR / \Delta \text{local SFR}$	Fainter in	$\sim 0.35 \text{ mag} / 5.3 \text{ Gyr}$
		$\approx 0.16 \text{ mag} / (\Delta \log(\text{sSFR}_{>-10.8} - \text{sSFR}_{<-10.8}, \text{yr}^{-1} \text{ kpc}^{-2}))$	Higher SFR environments	Fainter in Younger galaxy
Population Age	This work	$\Delta HR / \Delta \text{pop. age}$	Fainter in	$\sim 0.27 \text{ mag} / 5.3 \text{ Gyr}$
		$\approx 0.051 \text{ mag} / \text{Gyr (YEPS)}$	Younger galaxy	Fainter in Younger galaxy

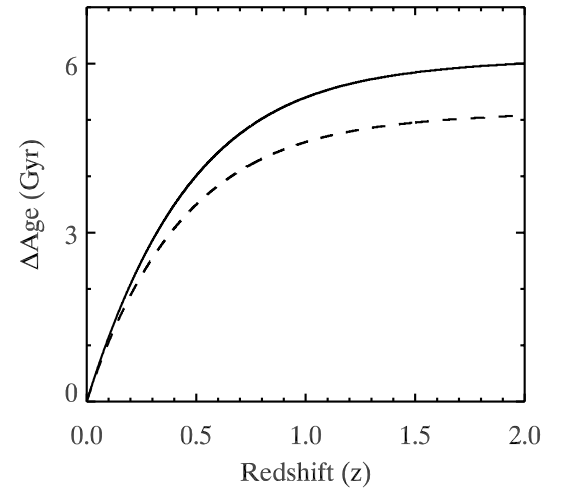
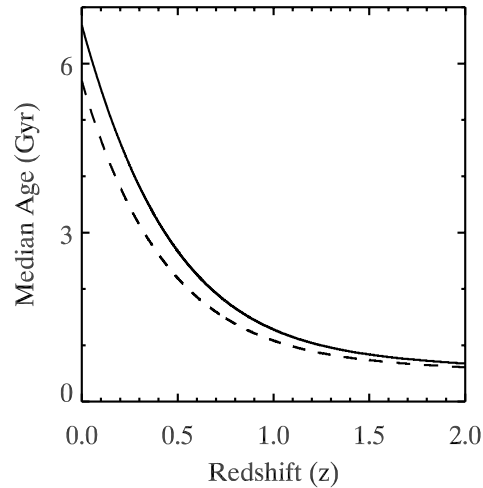
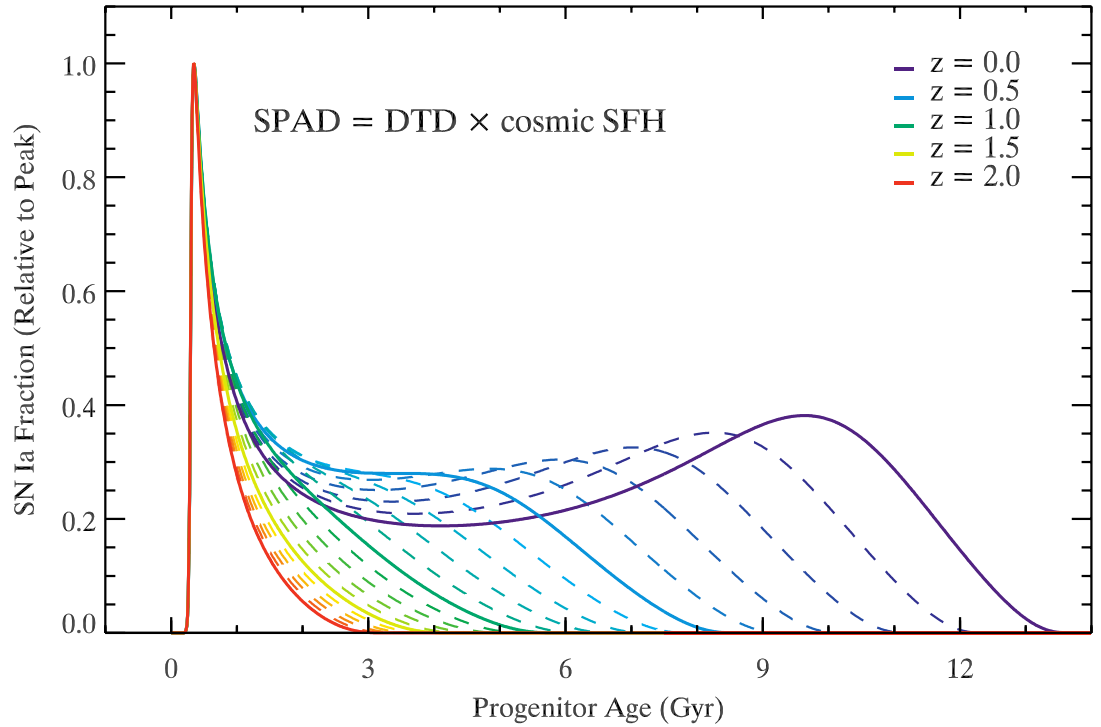
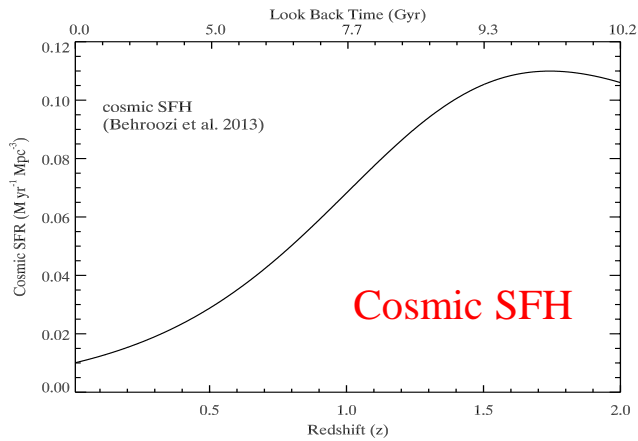
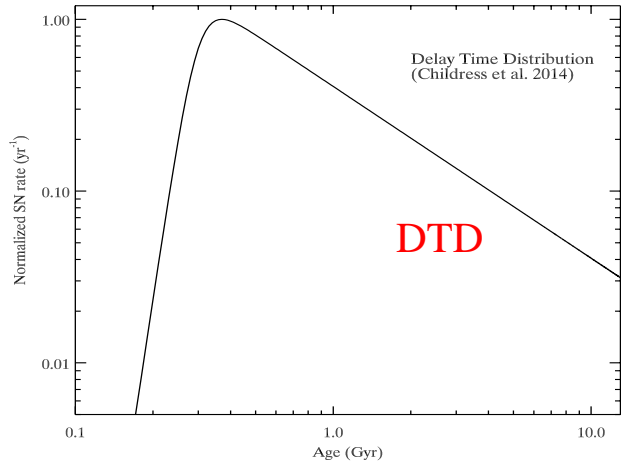
When they are converted to age difference

based on our result, Scott et al. 2017, & Galbany et al. 2014

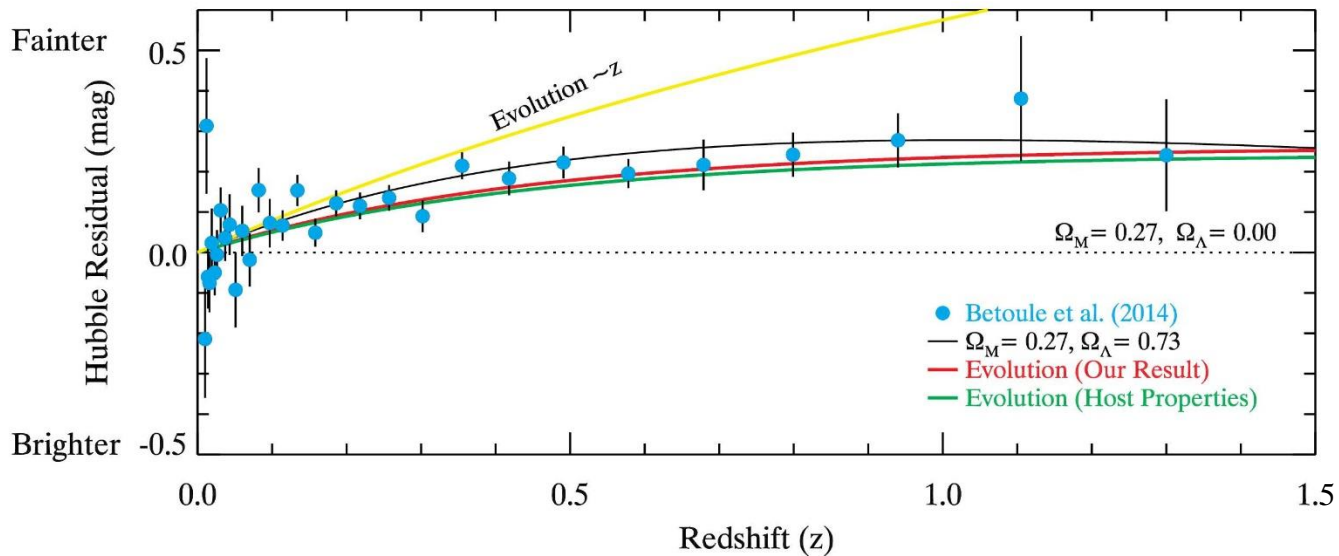
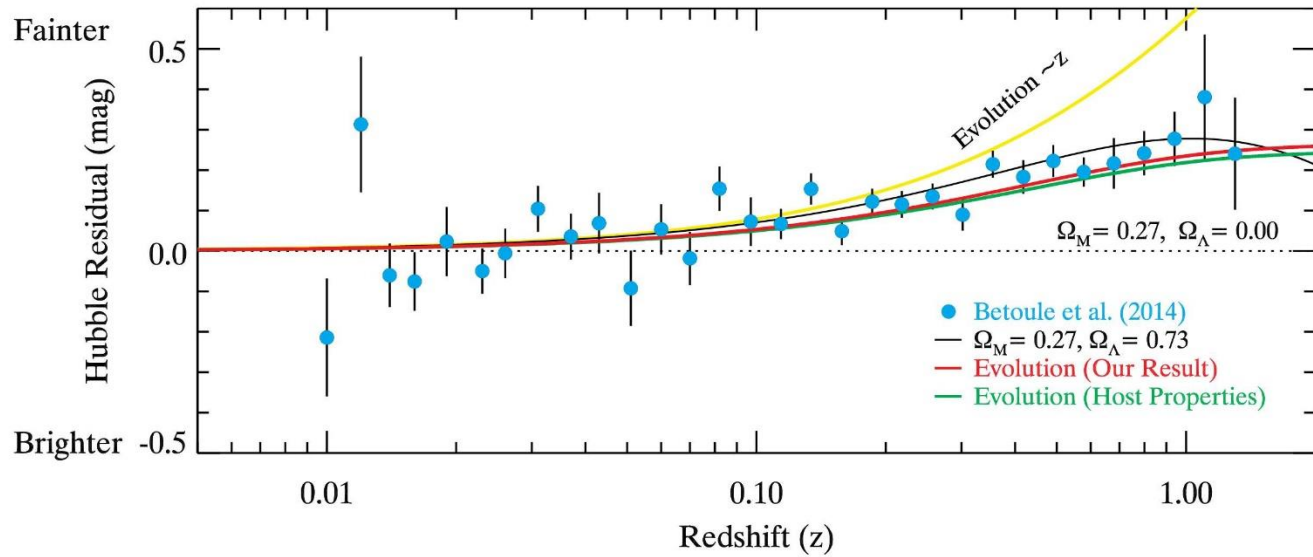
\rightarrow *They are all pointing to the same direction!*

\rightarrow **SNe Ia in younger galaxies (i.e., high-z) are fainter!**

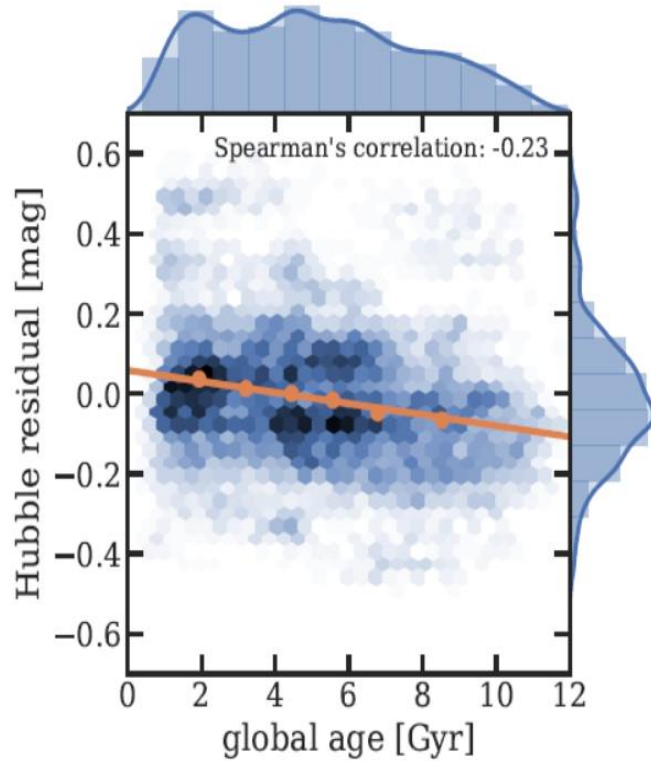
Supernova Progenitor Age Distribution (SPAD): following Childress et al. (2014)



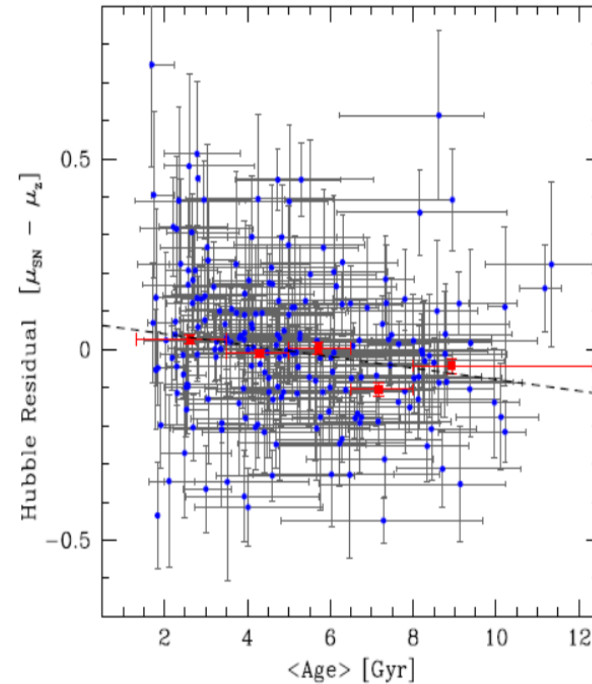
Luminosity evolution can mimic Hubble residuals used in the discovery of dark energy



Secondary Age-dating based on photometry (SED) is also pointing to the same direction (for all galaxy types)



Rose, Garnavich+2019



Gupta+2011

See also Pan, Sullivan+2014 & Neill+2009 for similar correlations

→ 0.08 – 0.11 mag/5.3 Gyr or "step"? → can mimic ~40 % of HR
→ *Still pose a serious problem to the assumption of no L evolution!*

Conclusion

- This is the most direct & stringent test ever made for the luminosity evolution of SNe Ia!
- Our result (**based on high-precision ages**) and other correlations (based on 2^{ndary} age-dating or age related properties) all indicate that SNe Ia in younger population environments (i.e., high-z) are fainter (after standardization)!
- *Taken at face values*, the luminosity evolution appears to be significant enough to question the very existence of Dark Energy!

“Extraordinary claims require extraordinary evidence”

Carl Sagan

→ *Is the evidence for Dark Energy secure from SN cosmology?*

*To put this result on a firmer refined basis,
further observations/investigations are definitely required*

1. for more ETG host sample (SDSS targets with Gemini/MMT)
2. for the local property at the site of SN (IFU or long-slit)
3. Improve SED based age-dating for all morphological types of host galaxies (GALEX UV to IR)
4. Origin of "over-correction" at younger environment?
5. Better statistical analysis (with $\sigma_{\text{int}} = 0.0$) for SN cosmology
6. BAO, CMB ?...

How about other cosmological probes?

Concordance model from SNe Ia, CMB, & BAO??

- 1. CMB from Planck:** no longer supports concordance model? (Di Valentino, Melchiorri, Silk+2019)
 - 2. BAO & other low- z probes:** shown to be consistent with a non-accelerating universe (Tutusaus, Blanchard+2017)
 - 3. SNe Ia:** luminosity evolution mimics dark energy (Kang, Lee+2020)
- The Crisis in Cosmology?

