## Generating Catalogues of Lensed Galaxies to Test Weak Lensing Pipelines

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Weak Lensing Formalism

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$$\vec{\beta} - \left(\vec{\theta} - \alpha(\vec{\theta})\right) = 0$$

Image credit Bartelmann and Schneider, (1999)

#### Lensing Parameters



 $\beta_{i} = A_{ij}\theta_{j} + \frac{1}{2}D_{ijk}\theta_{j}\theta_{k}$   $A_{ij} = \begin{pmatrix} 1 - \kappa - \gamma_{1} & -\gamma_{2} \\ -\gamma_{2} & 1 - \kappa + \gamma_{1} \end{pmatrix}$   $D_{ij1} = -\frac{1}{2}\begin{pmatrix} 3F_{1} + G_{1} & F_{2} + G_{2} \\ F_{2} + G_{2} & F_{1} - G_{1} \end{pmatrix}$   $D_{ij2} = -\frac{1}{2}\begin{pmatrix} F_{2} + G_{2} & F_{1} - G_{1} \\ F_{1} - G_{1} & 3F_{2} - G_{2} \end{pmatrix}$ 

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Data produced from current work

![](_page_3_Figure_0.jpeg)

Develop realistic galaxy images, incorporating observational features found in image data.

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Produce a realistic cluster gravitational lensing potential.

Use lensing potential to transform galaxies from source to lensing plane.

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### The Multidark N-Body Simulation

Data Field	Description
MainHaloID	Unique cluster identifier
HaloMass	Mass of dark matter halo, in solar masses
GalaxyType	Distinguishes between primary halo and subhalos
NFWConcentration	NFW concentration of halo
Х	X position of halo (in Mpc)
У	Y position of halo (in Mpc)
Z	Z position of halo (in Mpc)

- An N-body simulation run in 2013.
- ► Particle number: 3840<sup>3</sup>
- ▶ Box size: 1  $h^{-1}$  Gpc
- ▶ Mass resolution:  $1.5 \times 10^9 M_{\odot}$
- 10 million halos downloaded from the SAG catalogue.

(Knebe et al, 2016)

![](_page_5_Figure_0.jpeg)

## Following One Galaxy

Images produced from current work

#### Lenser Tests

- Lenser was used to process 100 generated images (around an SIS profile), extracting their flexion.
- I compare inputted flexions to the flexion measured by Lenser.
- Residuals have RMS values of
  - $F_1: 0.0136 \ arcseconds^{-1}$
  - F<sub>2</sub>: 0.0077  $arcseconds^{-1}$

![](_page_6_Figure_6.jpeg)

#### Future Work

- This tool generates catalogues to model specific lensing scenarios. This will provide a valuable resource to lensing pipelines which extract information from lensed galaxies.
- There are more tests to be run with Lenser, with larger catalogues, more complex lensing profiles, and a larger distribution of input flexions.
- There are additional potential selection criteria for MultiDark clusters that will allow specific physical systems to be modeled.

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![](_page_8_Picture_0.jpeg)

# Thank You!

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Image credit: NASA / ESA / Hubble

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