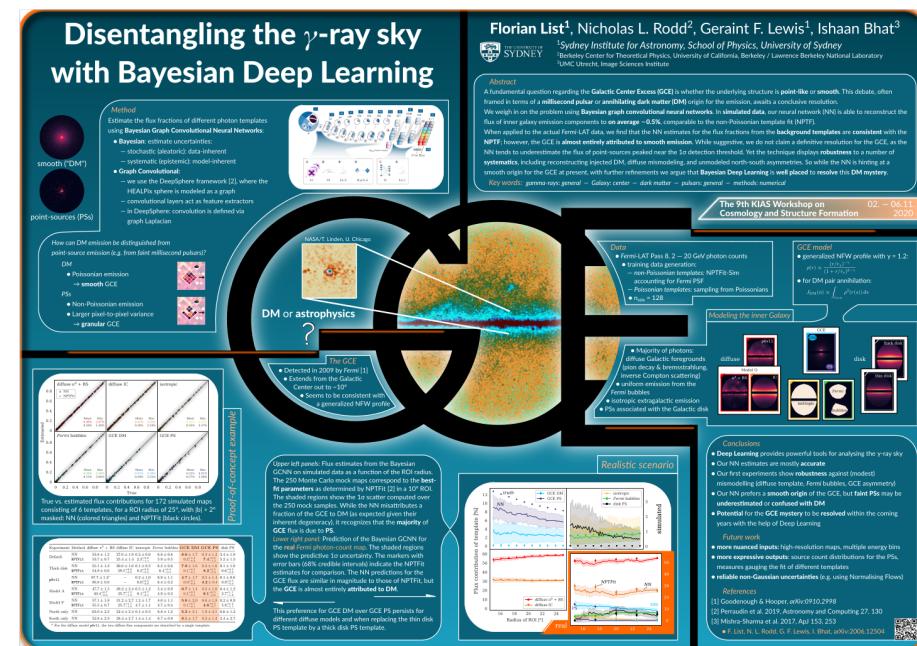




# The GCE in a new light: Disentangling the $\gamma$ -ray sky with Bayesian Deep Learning

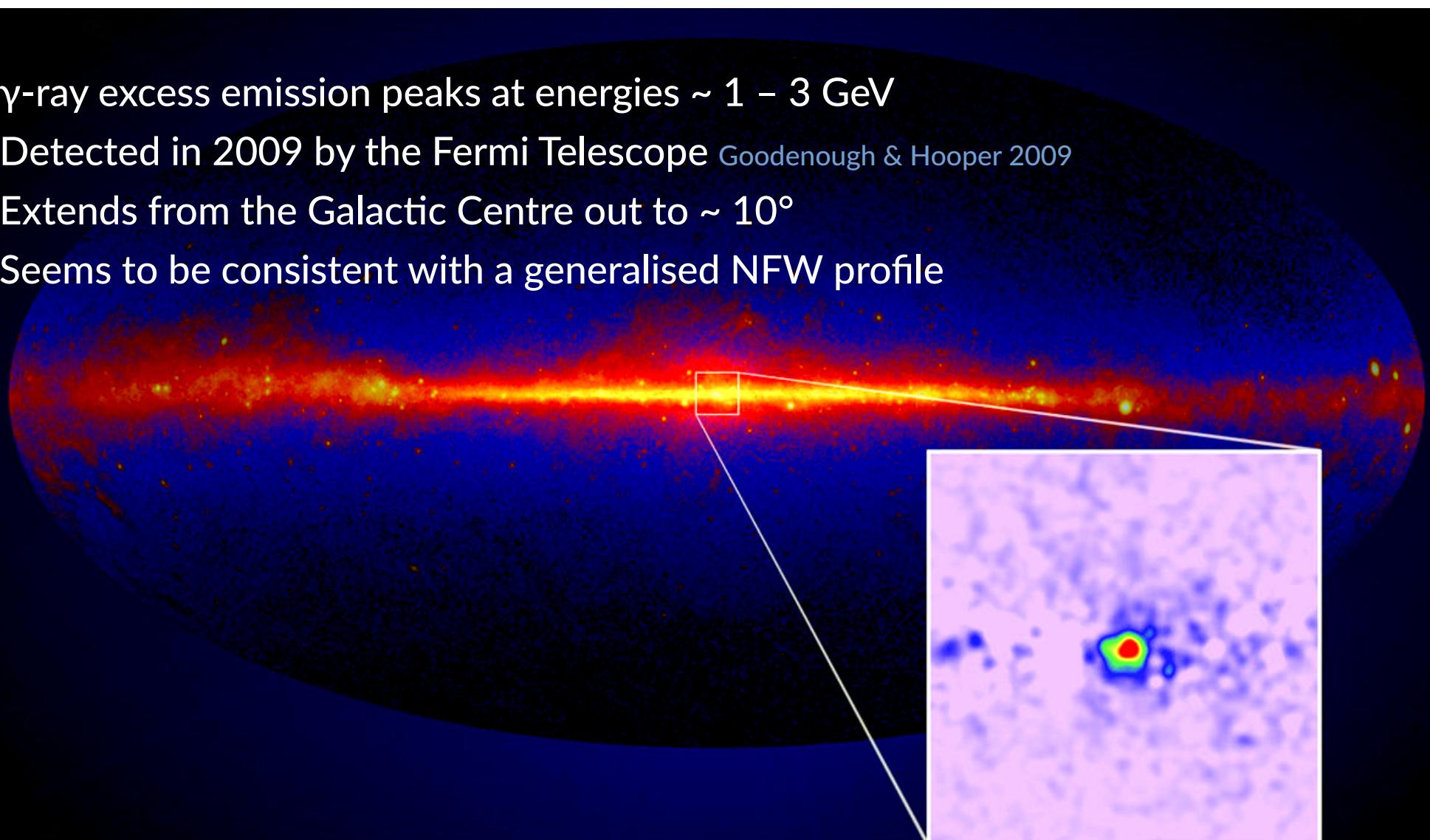
Florian List

Joint work with  
Nick Rodd,  
Geraint F. Lewis,  
Ishaan Bhat



# The $\gamma$ -ray GCE

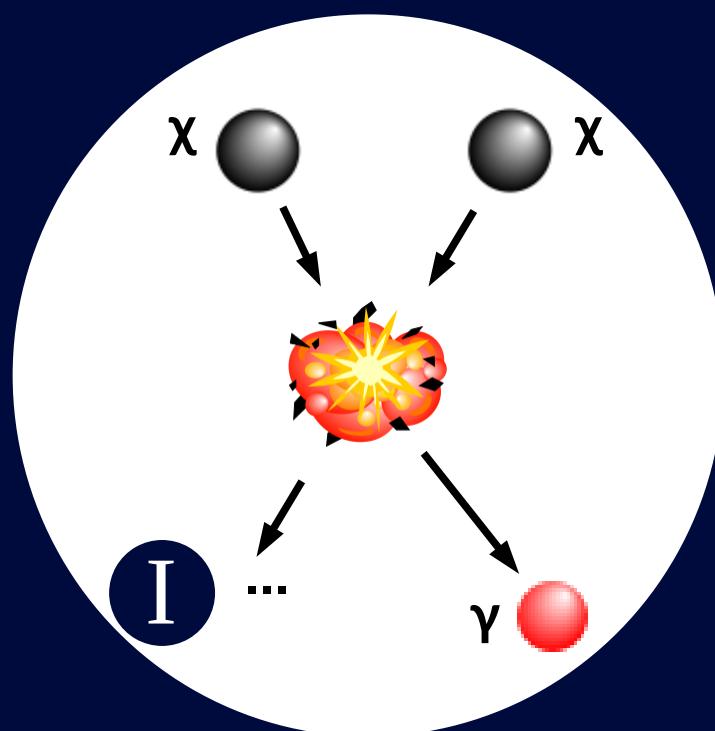
- $\gamma$ -ray excess emission peaks at energies  $\sim 1 - 3$  GeV
- Detected in 2009 by the Fermi Telescope [Goodenough & Hooper 2009](#)
- Extends from the Galactic Centre out to  $\sim 10^\circ$
- Seems to be consistent with a generalised NFW profile



# The $\gamma$ -ray GCE

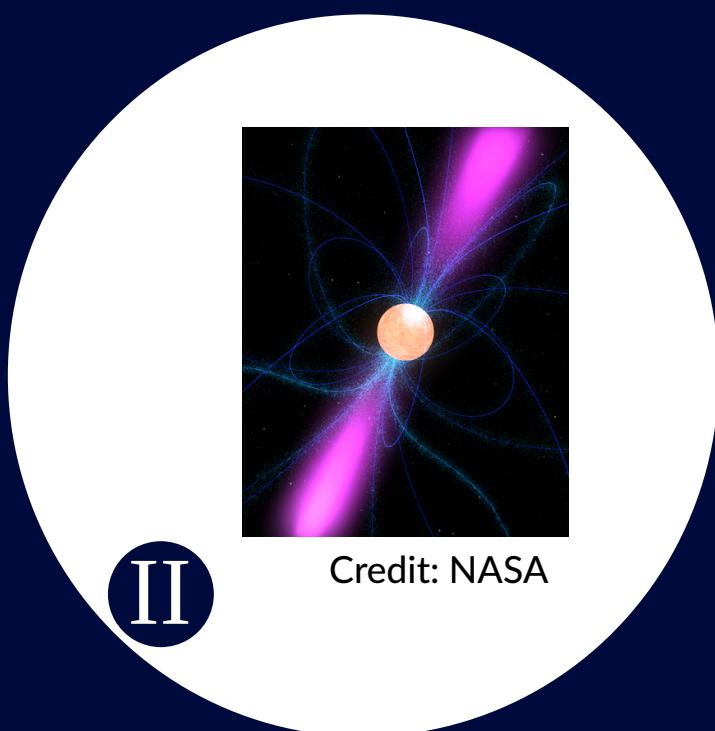
## What is the explanation?

annihilating  
DM?



I

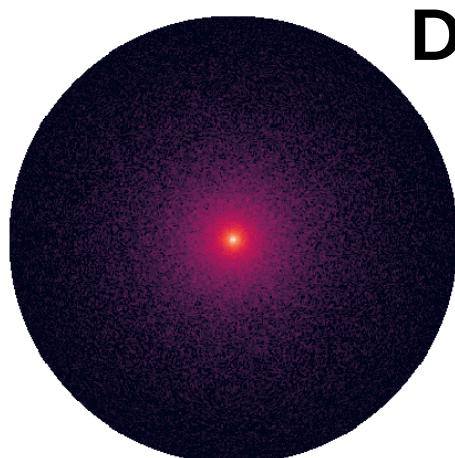
point sources  
(millisecond pulsars,  
cosmic rays, ...)?



II

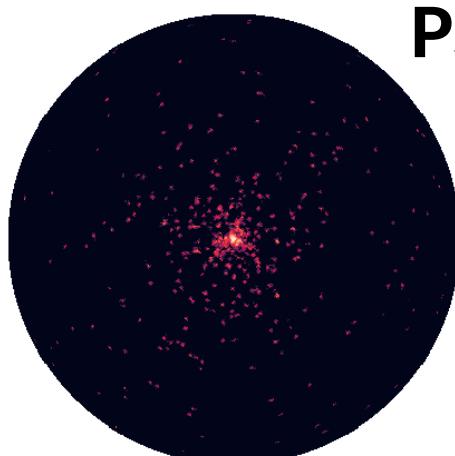
# The $\gamma$ -ray GCE

*How could annihilating DM and PSs be distinguished?*



DM

**smooth GCE,  
Poissonian emission**



PSs

**granular GCE,  
larger pixel-to-pixel variance,  
non-Poissonian statistics**

- Goodenough & Hooper (2009)  
Vitale & Morselli (2009)  
Hooper & Goodenough (2011)  
Hooper & Linden (2011)  
Boyarsky et al (2011)  
Abazajian & Kaplinghat (2012)  
Gordon & Macias (2013)  
Hooper & Slatyer (2013)  
Huang et al (2013)  
Mirabal (2013)  
Macias & Gordon (2014)  
Abazajian et al (2014, 2015)  
Zhou et al (2014)  
Daylan et al (2014)  
Petrovic et al (2015)  
Calore et al (2015)  
Cholis et al (2015)  
Selig et al (2015)  
Huang et al (2015)  
Gaggero et al (2015)  
Yuan & Ioka (2015)  
O'Leary et al (2015)  
Brandt & Kocsis (2015)  
Carlson et al (2015, 2016)  
Yand & Aharonian (2016)  
Horiuchi et al (2016)  
Lee et al (2016)  
Bartels et al (2016)  
Linden et al (2016)  
Ajello et al (2016)  
Ackermann et al (2017)  
Ploeg et al (2017)  
Macias et al (2018, 2019)  
Bartels et al (2018)  
Caron et al (2018)  
Clark et al (2018)  
Leane & Slatyer (2019, 2020a, 2020b)  
Zhong et al (2019)  
Chang et al (2020)  
Abazajian et al (2020)  
Buschmann et al (2020)

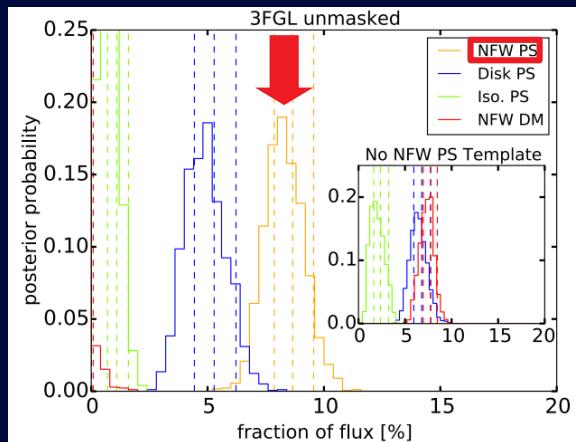
... (*not complete!*)

*x > 500 particle  
theory papers*

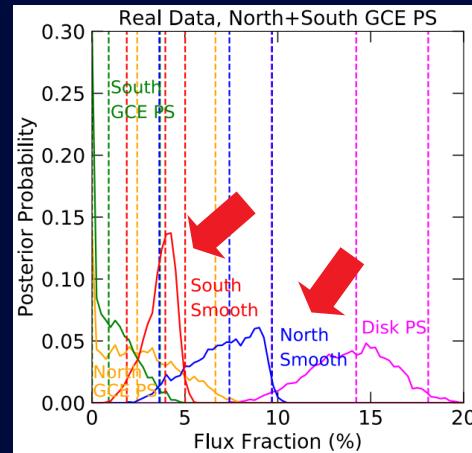
# Traditional analysis methods

Template fitting and wavelet method

- Current analysis methods are subject to **systematics**



Lee et al. 2016

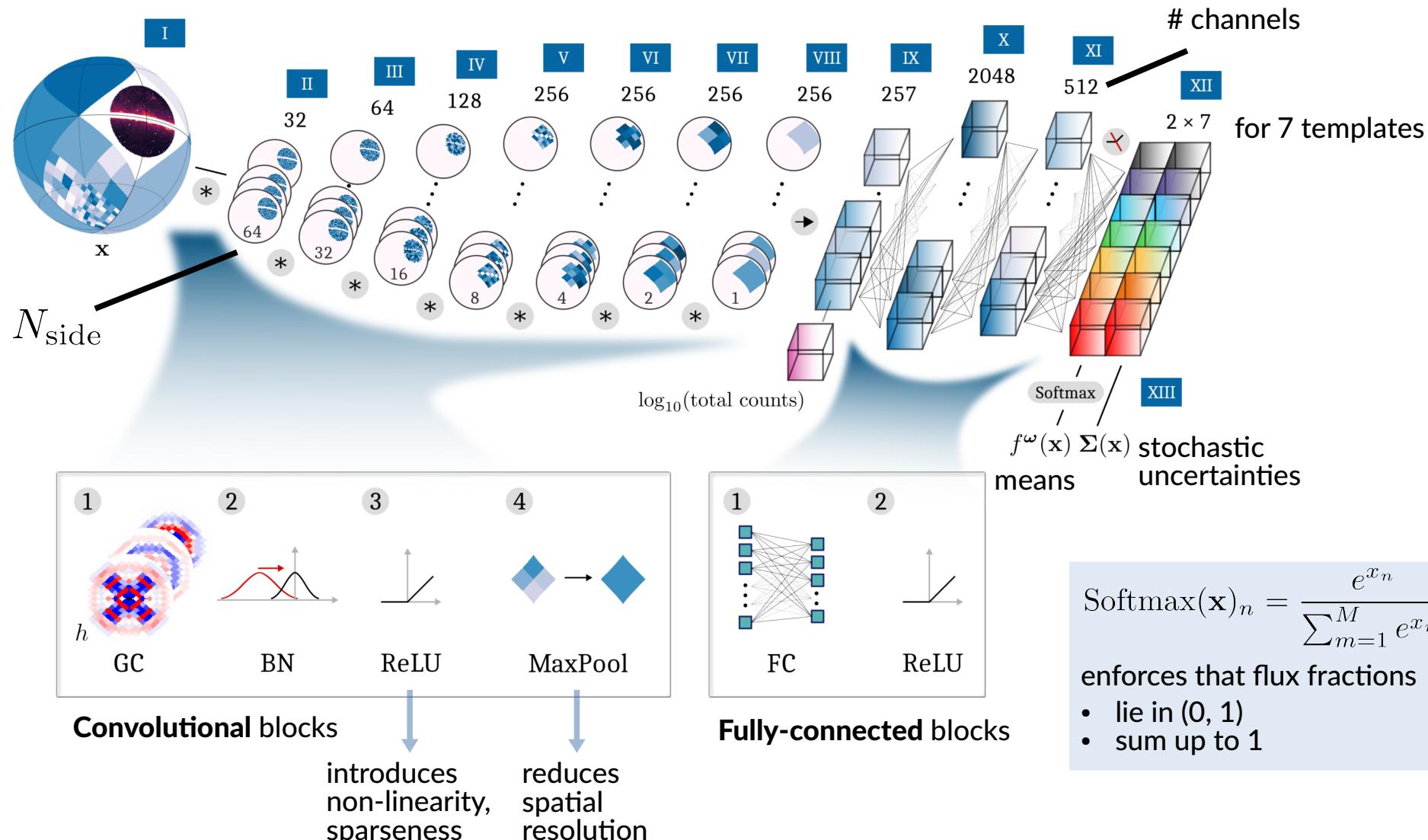


Leane & Slatyer 2020a

- See *Bartels et al. 2016, Lee et al. 2016, Leane & Slatyer 2019, Zhong et al. 2019, Chang et al. 2020, Leane & Slatyer 2020 a,b, Buschmann et al. 2020*

→ **GCE mystery still awaits its resolution!**

# → New method: Bayesian Convolutional Neural Networks

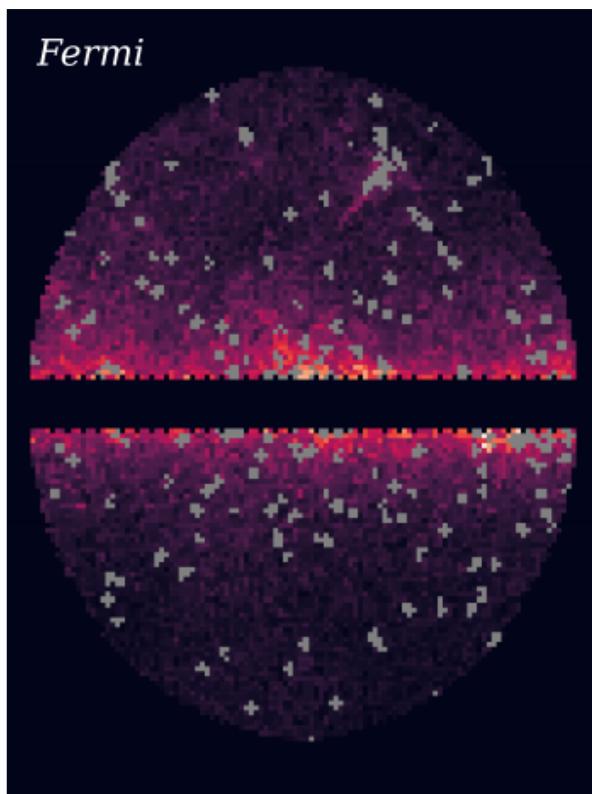


$$\text{Softmax}(\mathbf{x})_n = \frac{e^{x_n}}{\sum_{m=1}^M e^{x_m}}$$

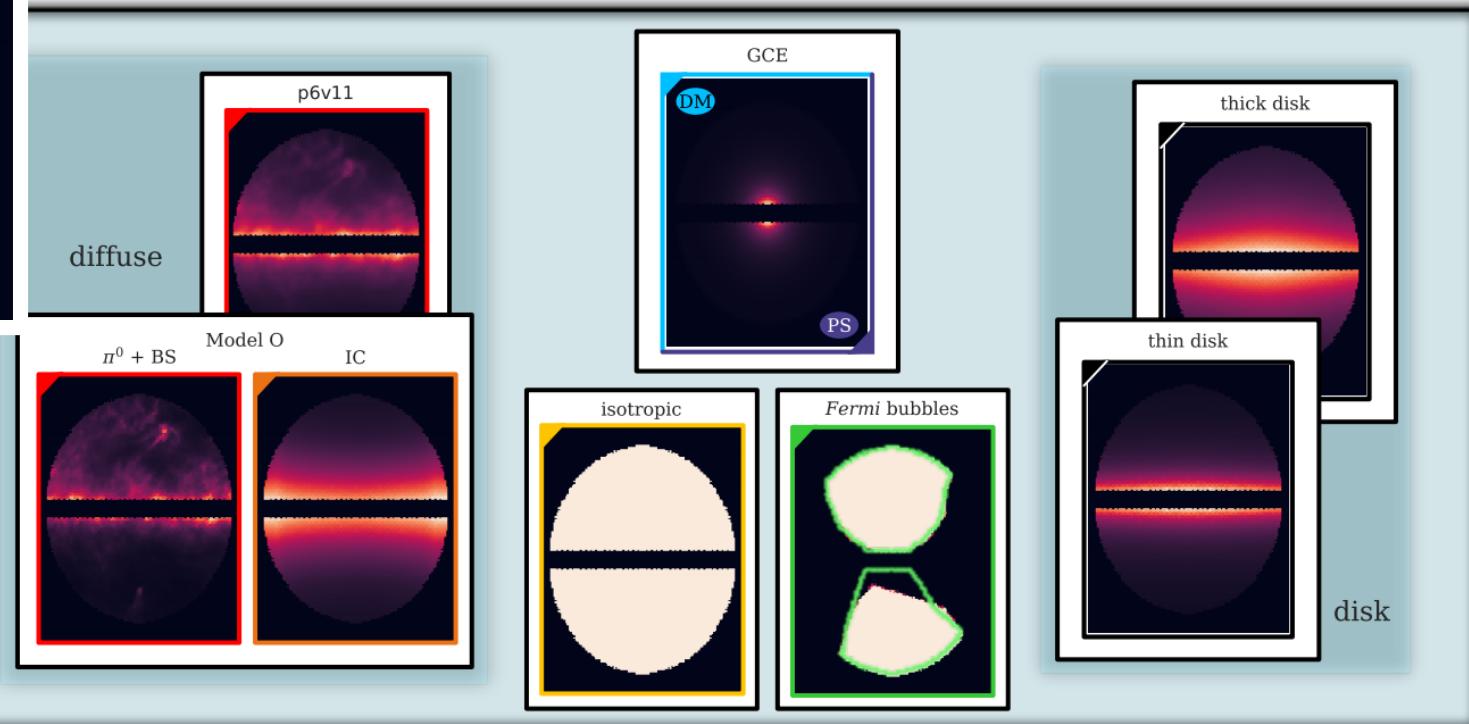
enforces that flux fractions

- lie in  $(0, 1)$
- sum up to 1

# Modelling the inner Galaxy

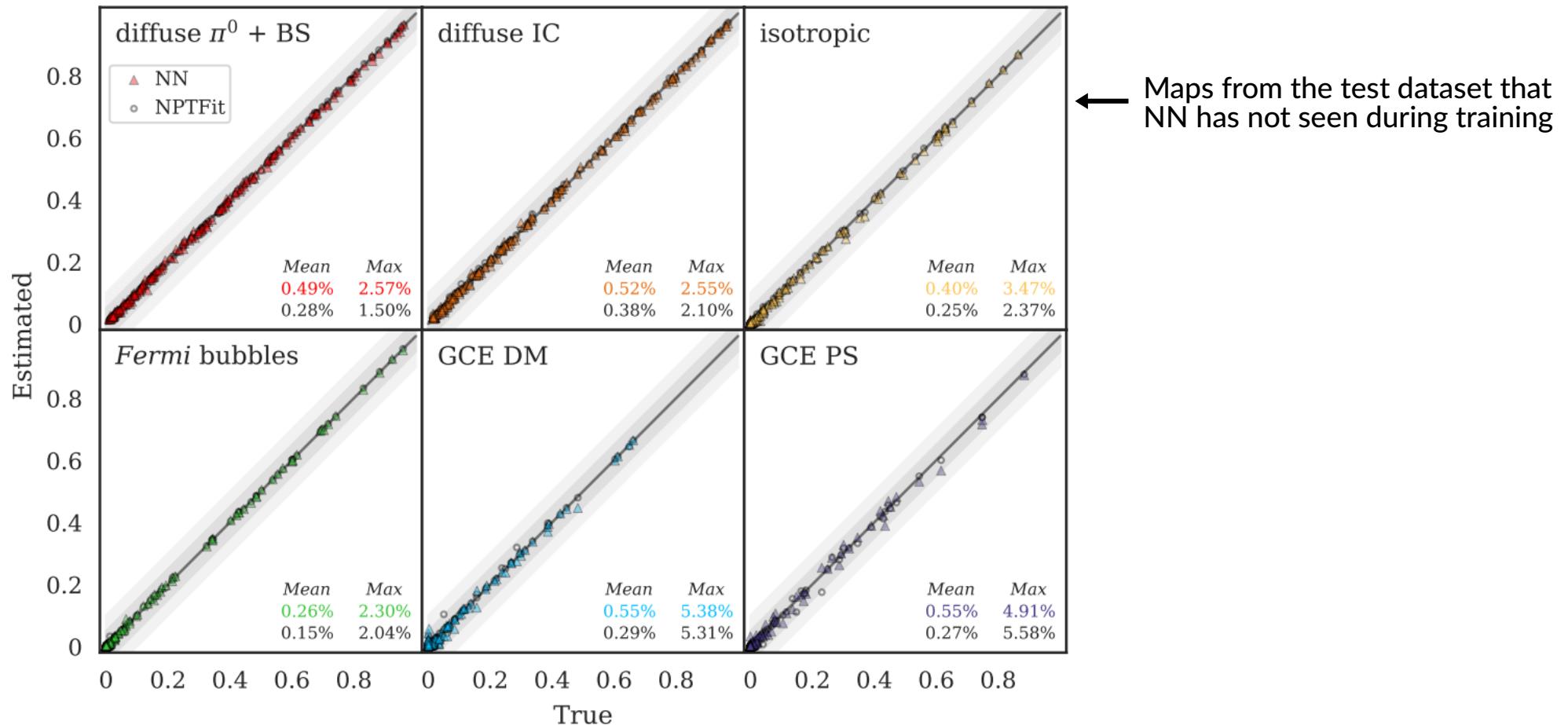


- Majority of detected photons: **diffuse Galactic foregrounds** (pion decay + bremsstrahlung, IC)
  - Uniform emission from the Fermi bubbles
  - Isotropic extragalactic emission
  - PSs associated with the Galactic Disk
- We generate the training maps using **NPTFit-Sim**



# Proof-of-concept example

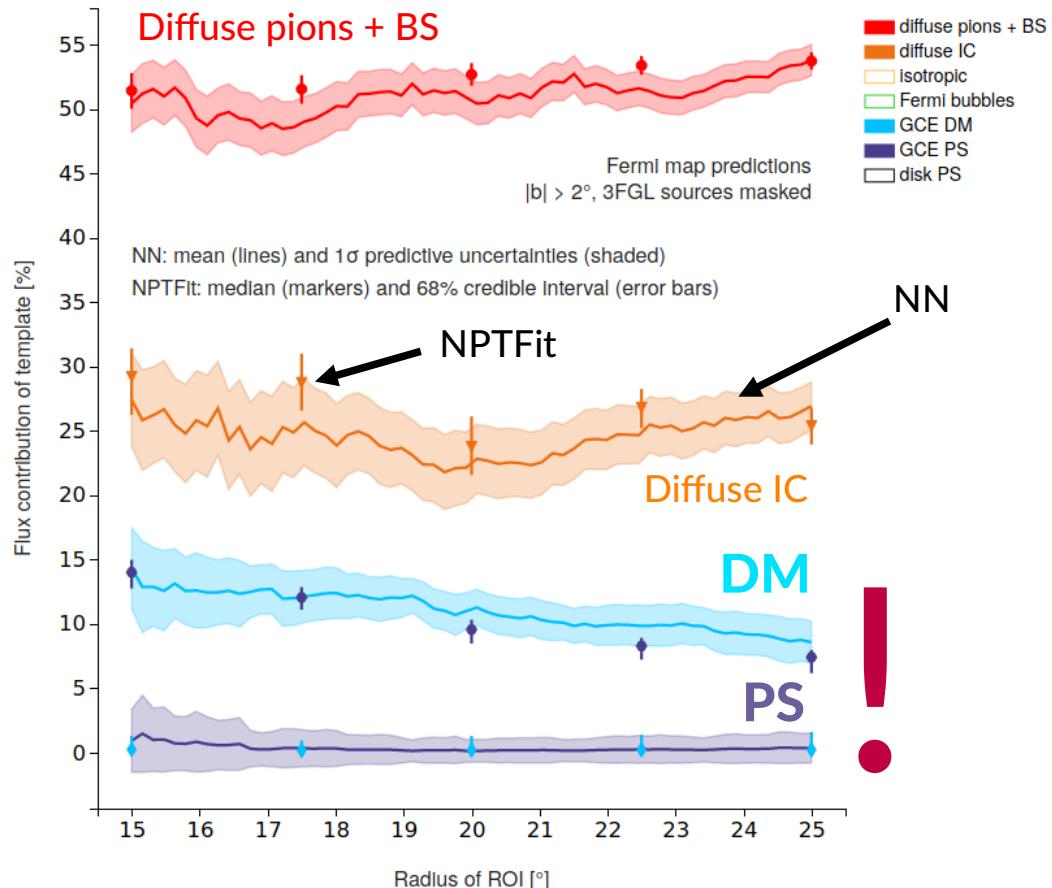
*Results on randomly generated maps*



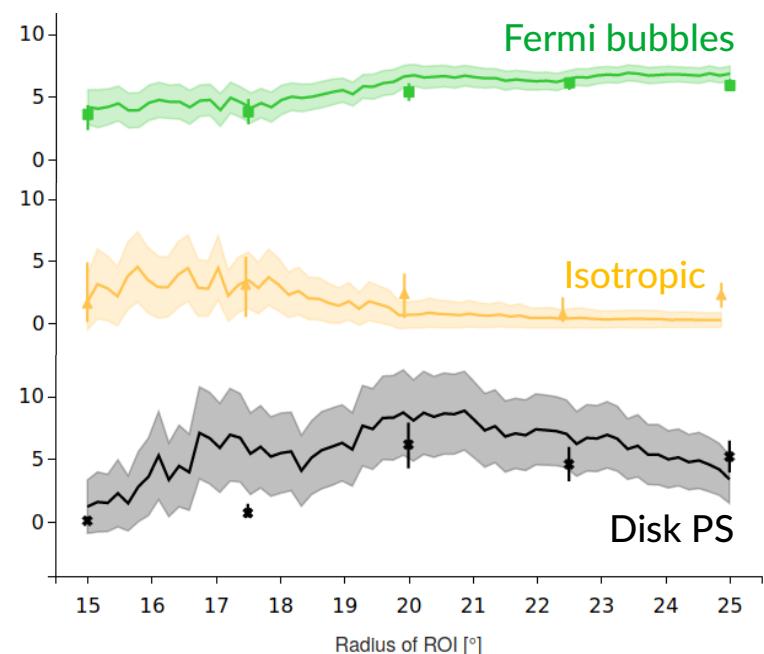
- On average, NN accuracy is a bit worse than NPTFit
- **But: mean errors** are only  $\sim 0.5\%$  (in comparison: GCE contribution is  $\sim 4 - 10\%$ )
- Maximum errors for GCE templates are very similar for NN and NPTFit

# Realistic scenario

## NN results for the *Fermi* map



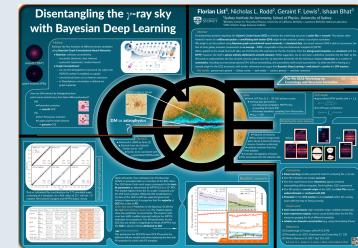
→ NN identifies a GCE that decreases monotonically with the ROI radius



→ Almost 100% of the GCE flux is attributed to DM



Poster:



# Conclusions

- Deep Learning provides powerful tools for analysing the  $\gamma$ -ray sky
- NN estimates are mostly accurate
- Our first experiments show robustness against mismodelling
- Our NN prefers a smooth origin of the GCE, but faint PSs may be underestimated / confused with DM
- Potential for the GCE mystery to be resolved within the coming years with the help of Deep Learning

