



고등과학원(KIAS) & SDSS 한국과학자 그룹(KSG)

# SDSS-KSG 하계 워크숍

▶ 일시 : 2005년 8월 22일 ~ 24일

▶ 장소 : 안면도



# SPECTR

(Spectrophotometer for Transmission Spectroscopy of Exoplanet Systems)

## observation of planetary transit systems

MYEONG-GU PARK (KYUNGPOOK NAT. U.)

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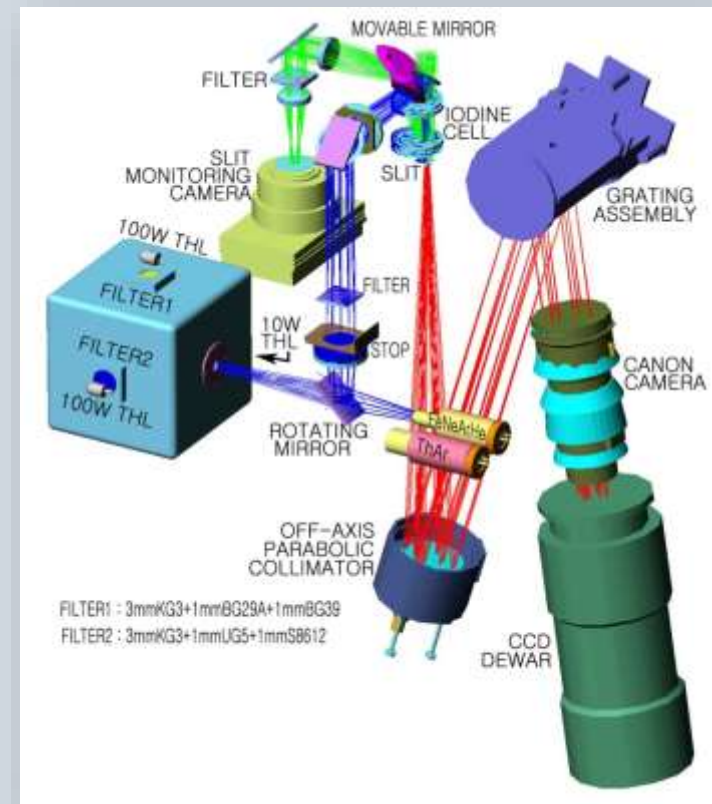
KNU: TAE-YANG BANG,

KASI: KANG-MIN KIM (PROJECT ENGINEER), YEON-HO CHOI,  
CHAN PARK, INWOO HAN, BYEONG-CHEOL LEE, UEEJEONG  
JEONG, BE-HO JANG, JEONG GYUN JANG, SANG-MIN LEE

KONGJU: JAE-RIM KOO

# Exoplanet Studies with BOES

- Bohyunsan Observatory Eschelle Spectrograph + 1.8m
- Precision RV observations of giants and bright stars
- K-EXO Collaboration
  - KASI: Byeong-Cheol Lee, Kang-Min Kim (BOES), Inwoo Han (RVI2CELL) +
  - KNU: MGP +
- Achievement
  - **35** exoplanets out of 700 stars over 16 years
    - Mostly evolved giant stars
  - Stellar oscillation, rotation, and surface activity



# From discovery to the characterization of planets

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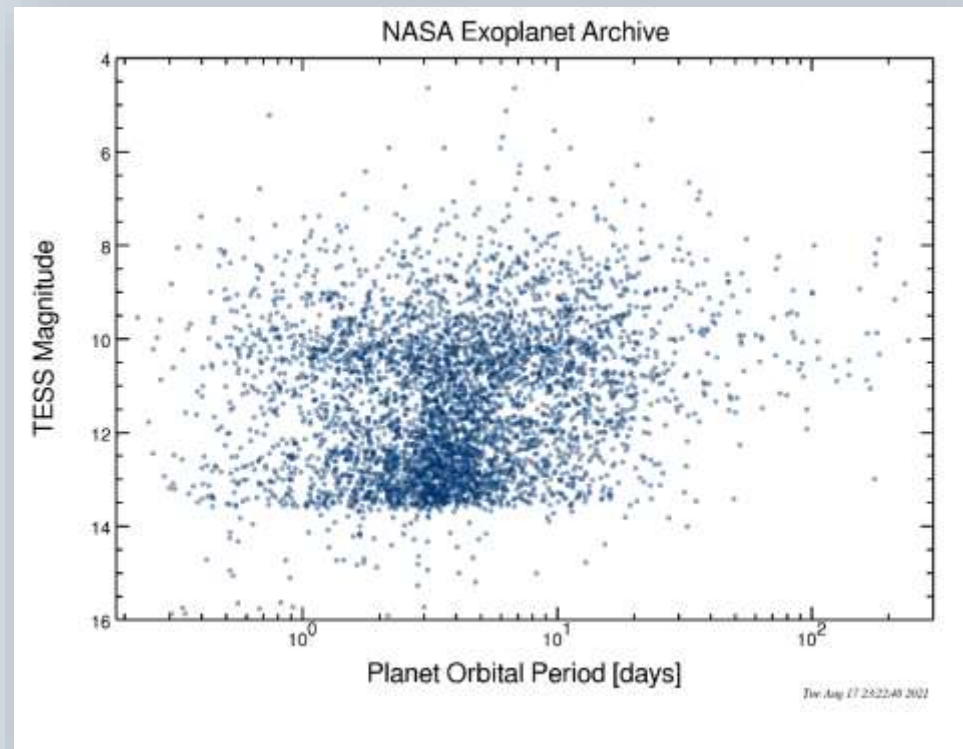
Density: rocky or gaseous

- Radius – Transit observation
- Mass – RV observation

Exoplanet atmosphere

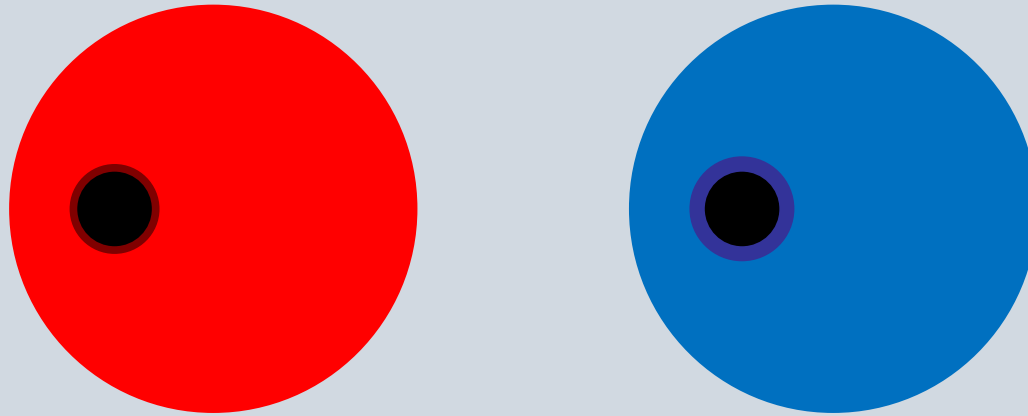
- Composition of exoplanet atmosphere
- Biomarker from earth-like planets within habitable zone
  - **Key science of GMT+G-CLEF**

Bright-enough targets available, e.g, from TESS



# Transmission Spectroscopy (투과분광)

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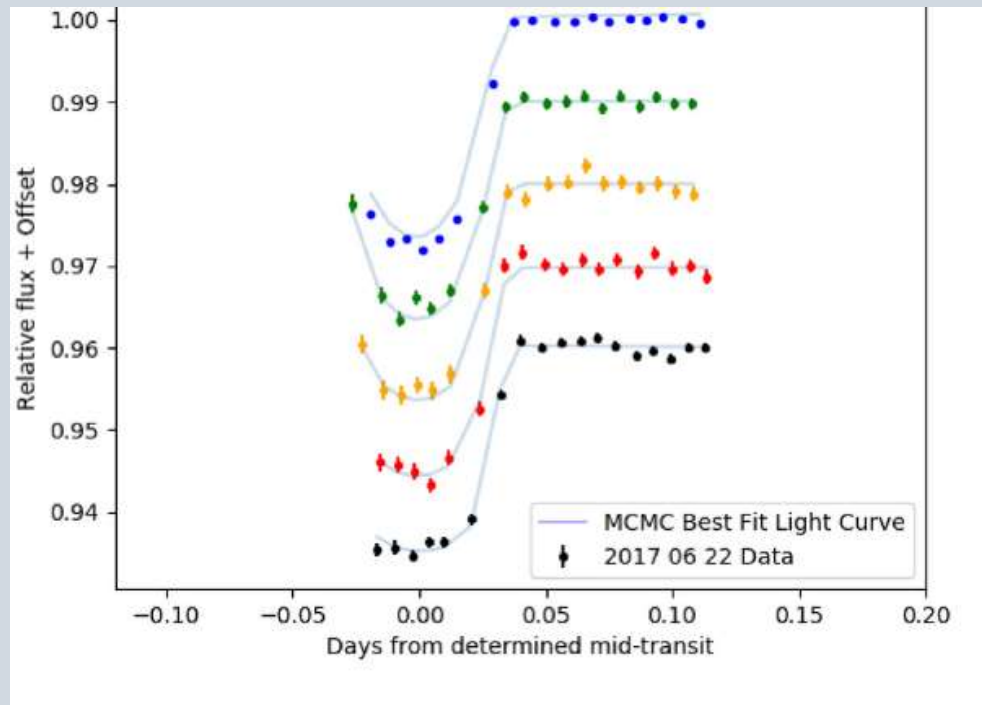


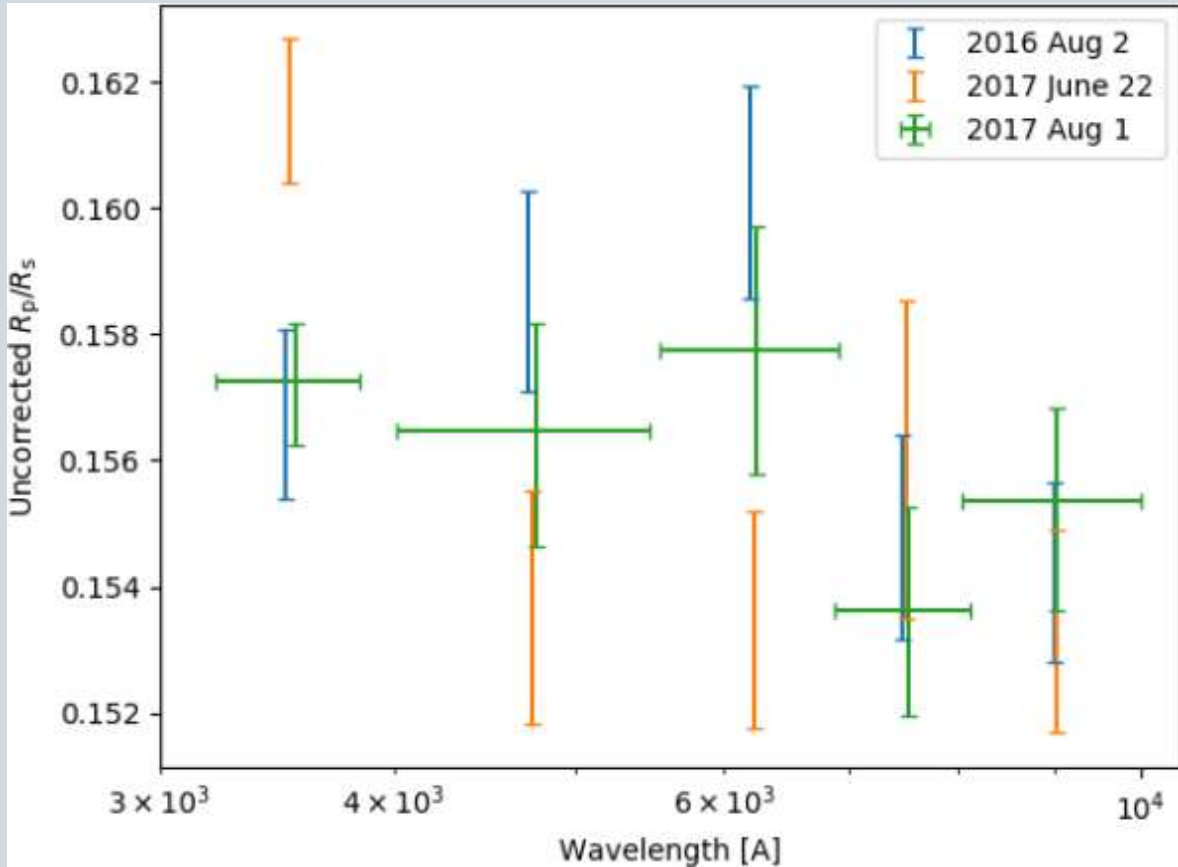
- Brightness change during the transit depends on the wavelength.
- Transmitted spectrum reflects the composition of the exoplanet atmosphere.
- Precise spectrophotometry required.

# Multi-Band Transmission Photometry

Kasper et al. (2019)

- A transmission spectrum of HD 189733b from multiple broad-band filter observations
- Wyoming Infrared Observatory
- 2.3 m telescope
- Sloan  $u'$ ,  $g'$ ,  $r'$ ,  $I'$ ,  $z'$
- 5 comparison stars





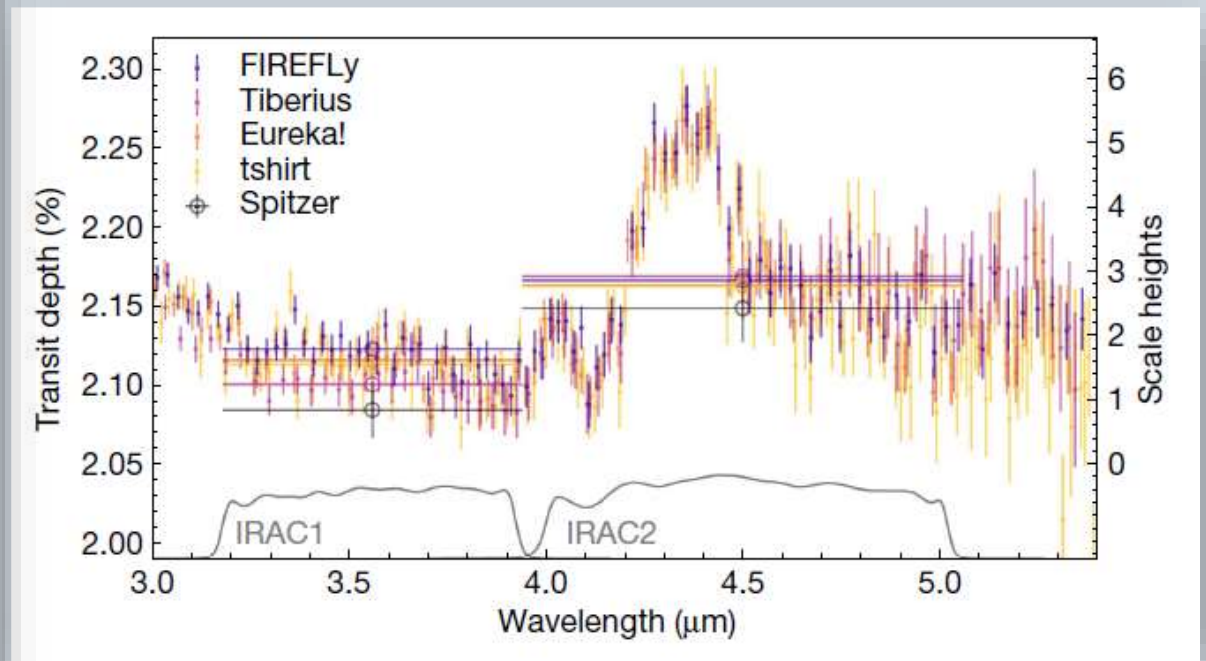
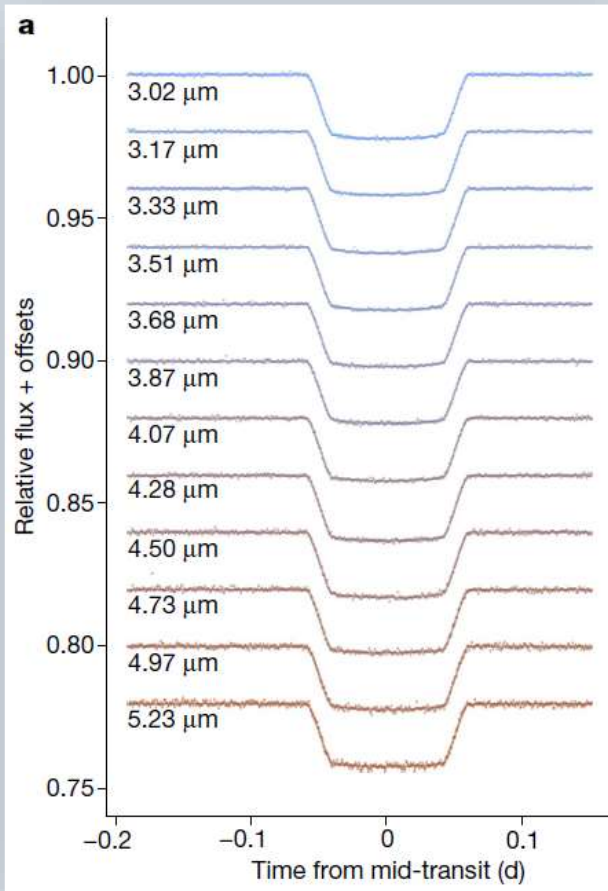
Uncorrected for star spots

Kasper+ 2019

$$\delta(R_p/R_*) \sim 0.002$$



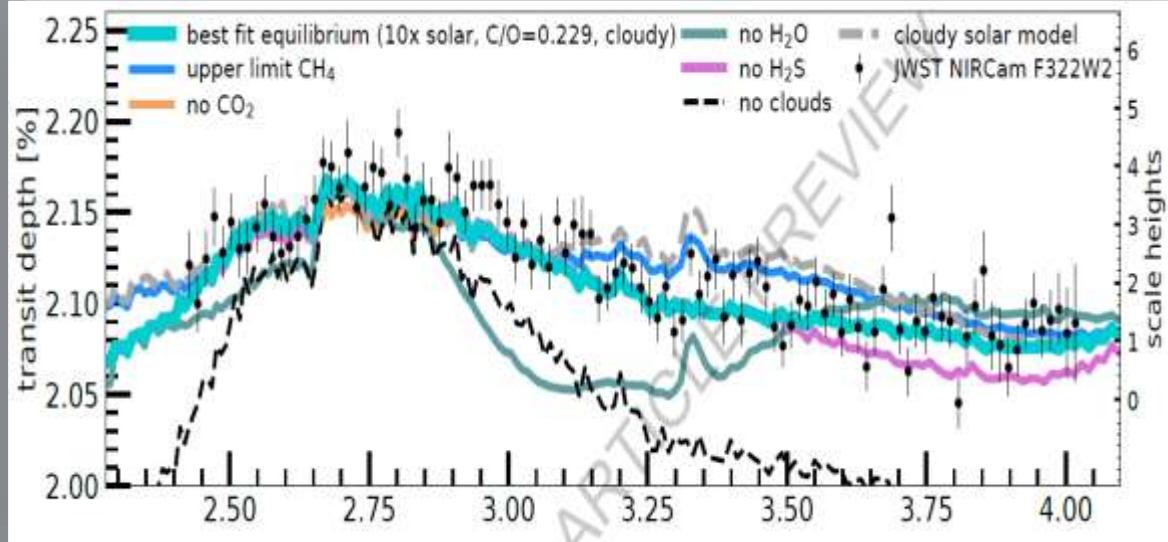
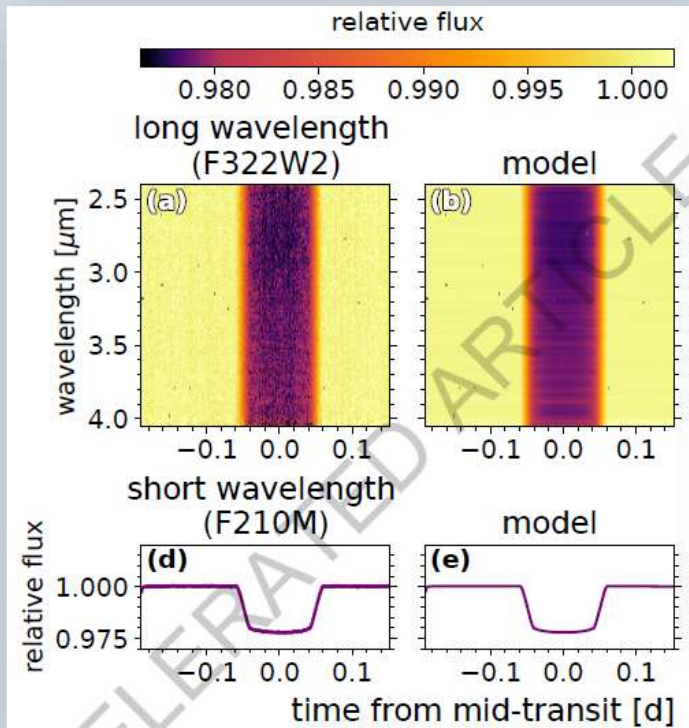
# JWST NIRSpec on WASP-39b



JWST Transiting Exoplanet Community  
Early Release Science Team 2023 Nature



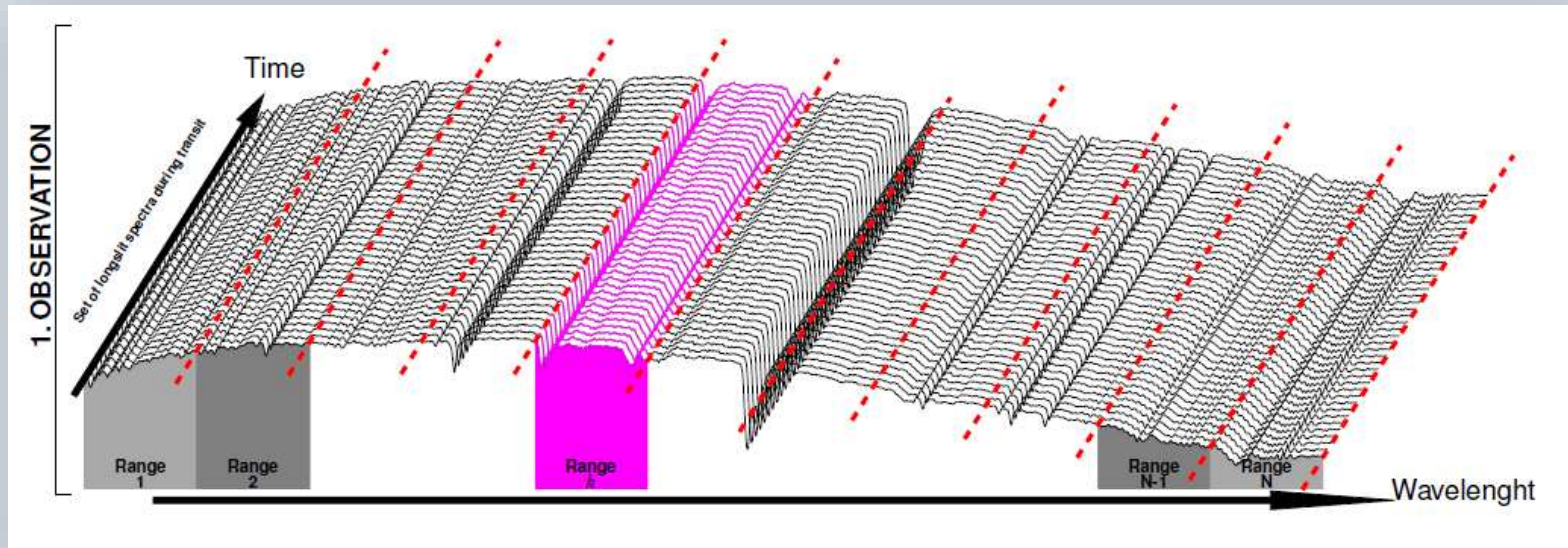
# JWST NIRCcam on WASP-39b



Ahrer+ 2023 Nature

# TR Spectroscopy with 1~2 m Telescopes

Time series photometry with low dispersion spectral observation during and off the transit

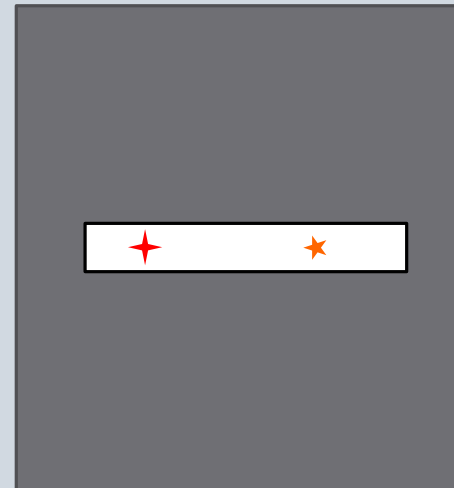


G. Valyavin

# Design Requirements

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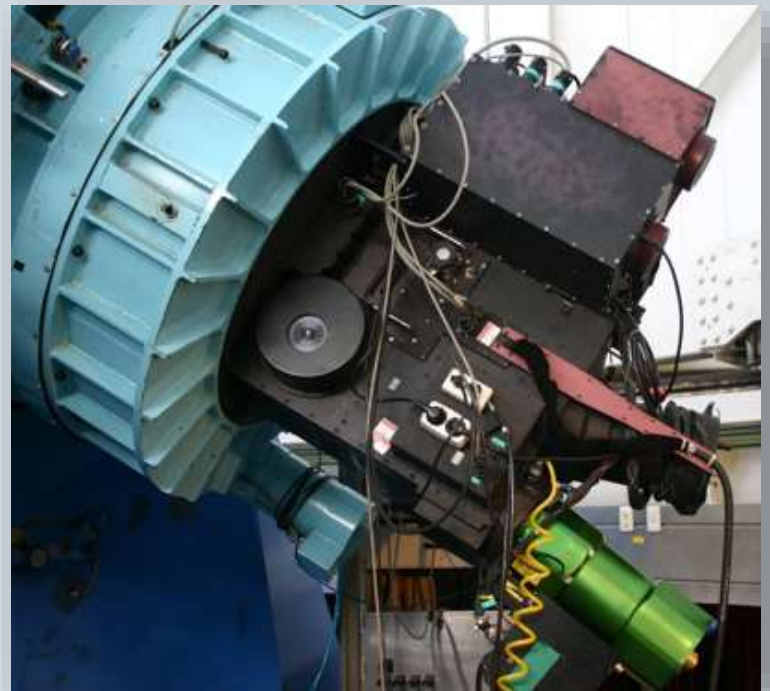
- Wavelength coverage: as wide as possible
- Very wide slit width to minimize the seeing induced variation
- Low to medium spectral resolution
- **Very long slit length for simultaneous observation of target and comparison star**
- Direct imaging capabilities



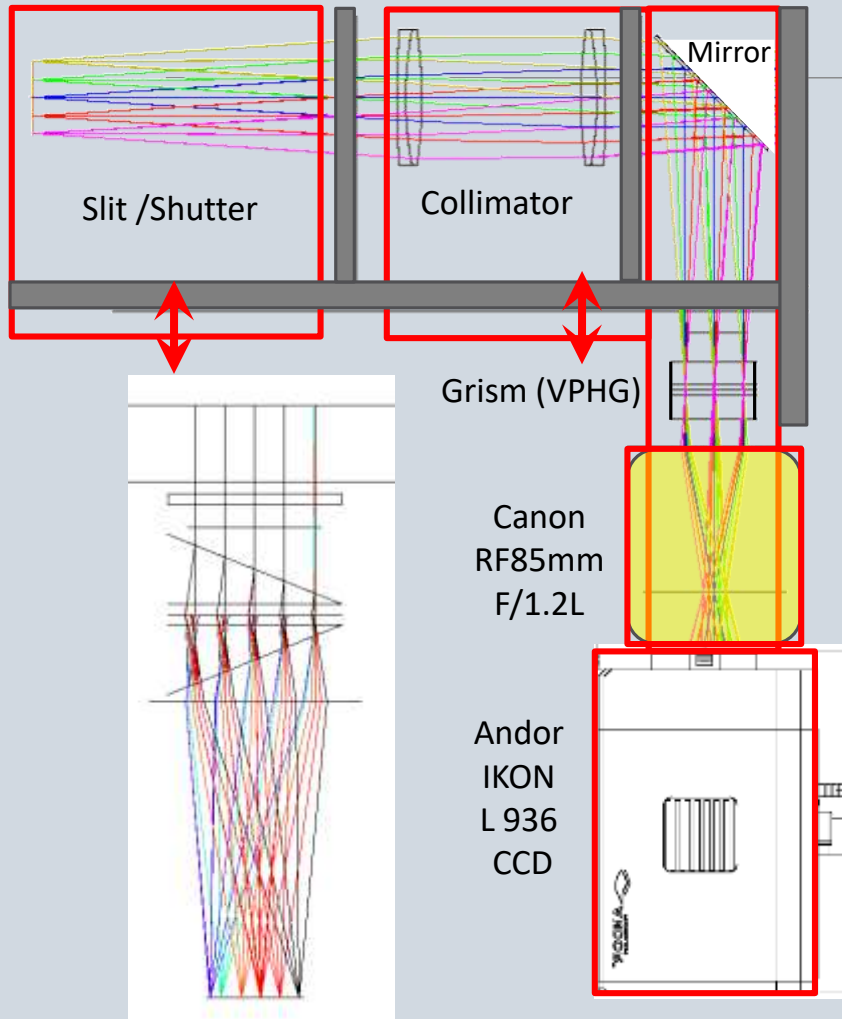
# Design Limitations

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- Limited space behind the 1.8m telescope
- Auxiliary systems as in CIM needed
  - Calibration lamps
  - Slit monitoring camera
  - Auto-guiding system
- Mounting/demounting CIM too costly



# Optics



Wavelength : **380-685**nm

Length: 401.7mm + 533.7mm

FOV : 10 x 4.6 arcmin.

Collimator : **f280mm**, Optosigma DBL80-500 x 2ea

Imaging Lens : Canon RF **85mm** F/1.2L

CCD : Andor IKON L936 CCD( $13.5\mu\text{m} \times 2048^2$ )

$\Phi$ (pupil) : **35** mm

Grism : Wasatch **900** g/mm, 50 x 50 mm, Blaz. $\lambda$ @526.1nm

Apex Angle of Wedge Prisms : 25.3 deg.

Cal Optics, guiding, slit monitoring in CIM

Plate Scale :  $21.3 \mu\text{m}/\text{arcsec}$  (4px/2.5", 0.62"/px)

Resolution: **R > 400** (RS ~ 2000 ['])

No Rotational (tilting) VPHG function.

\* Storage Container : Test & Alignment

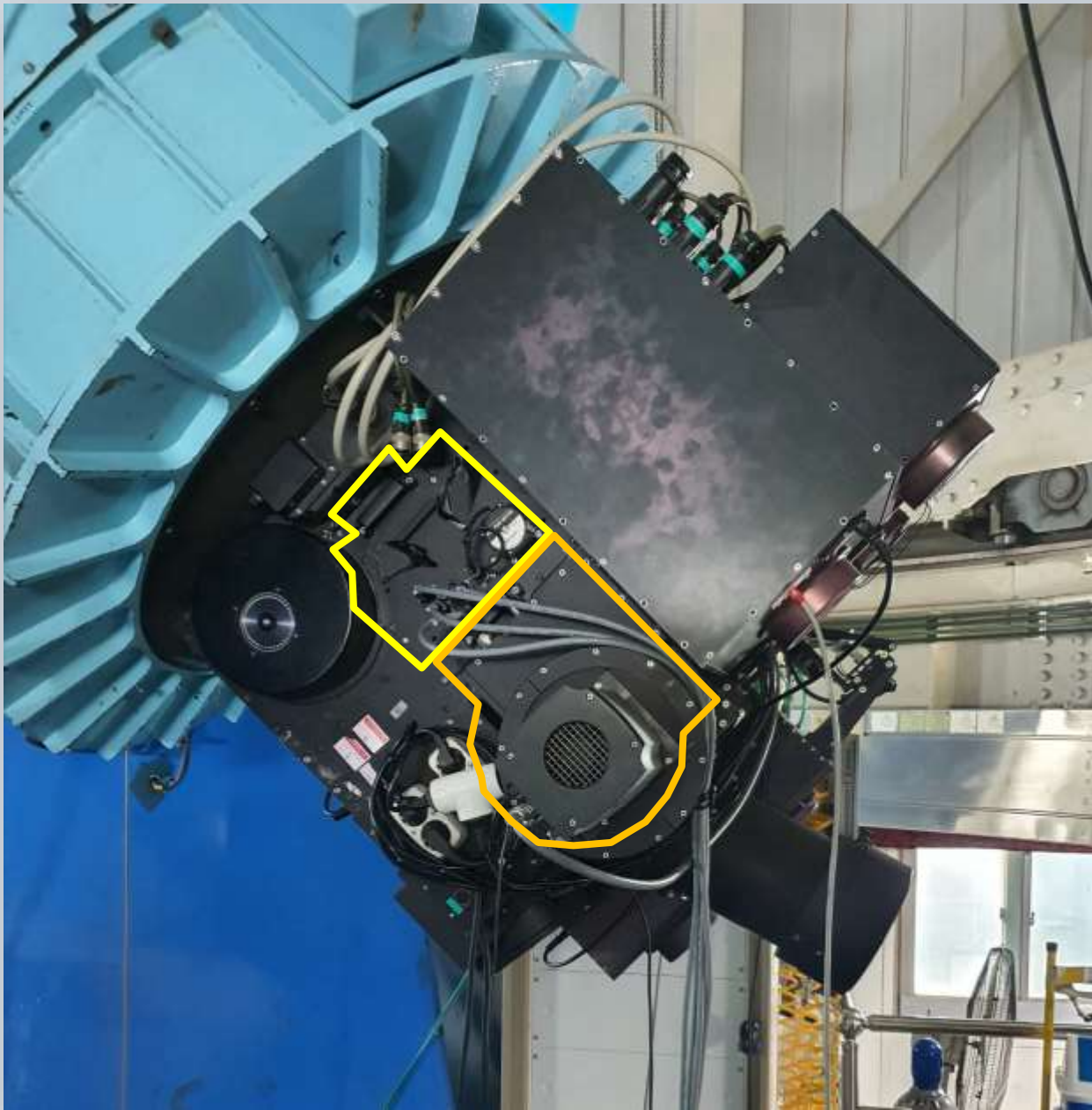
# Two Cartridge System

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- Slit assembly
- Collimator assembly
- Mirror assembly
- Grism assembly
- Imaging lens assembly
- CCD assembly









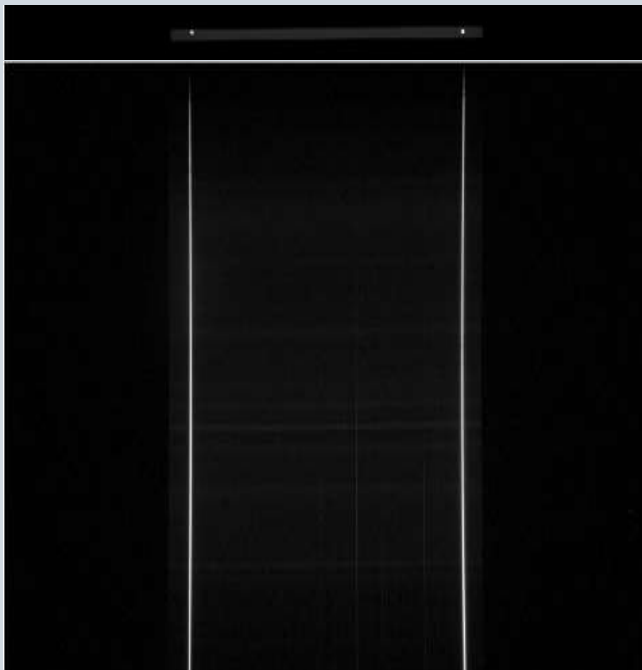
# System Adjustment & Test Observation

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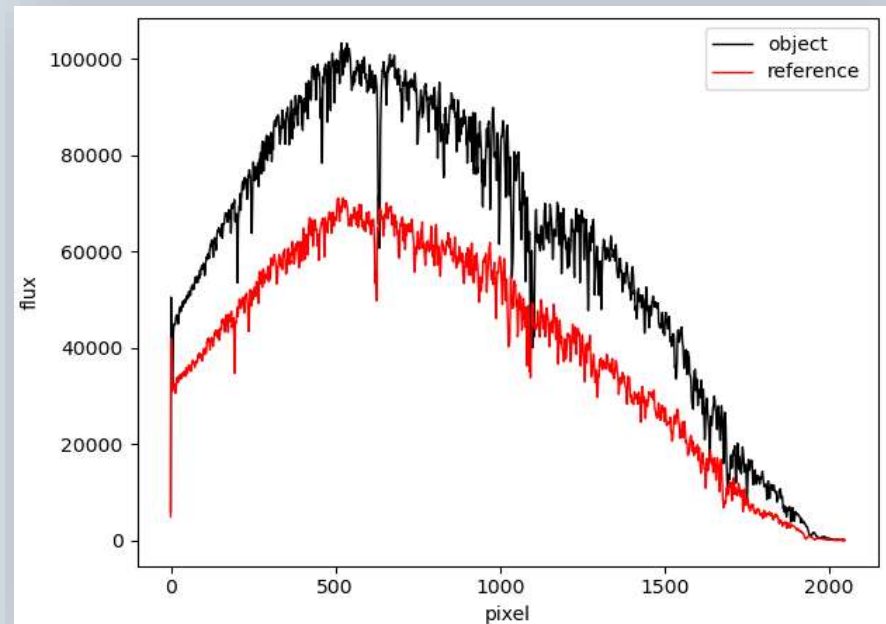
- 20 days: September & October of 2021
- System check & adjustment
- Test observation
  - 6 TESS transit systems

# HD 189733

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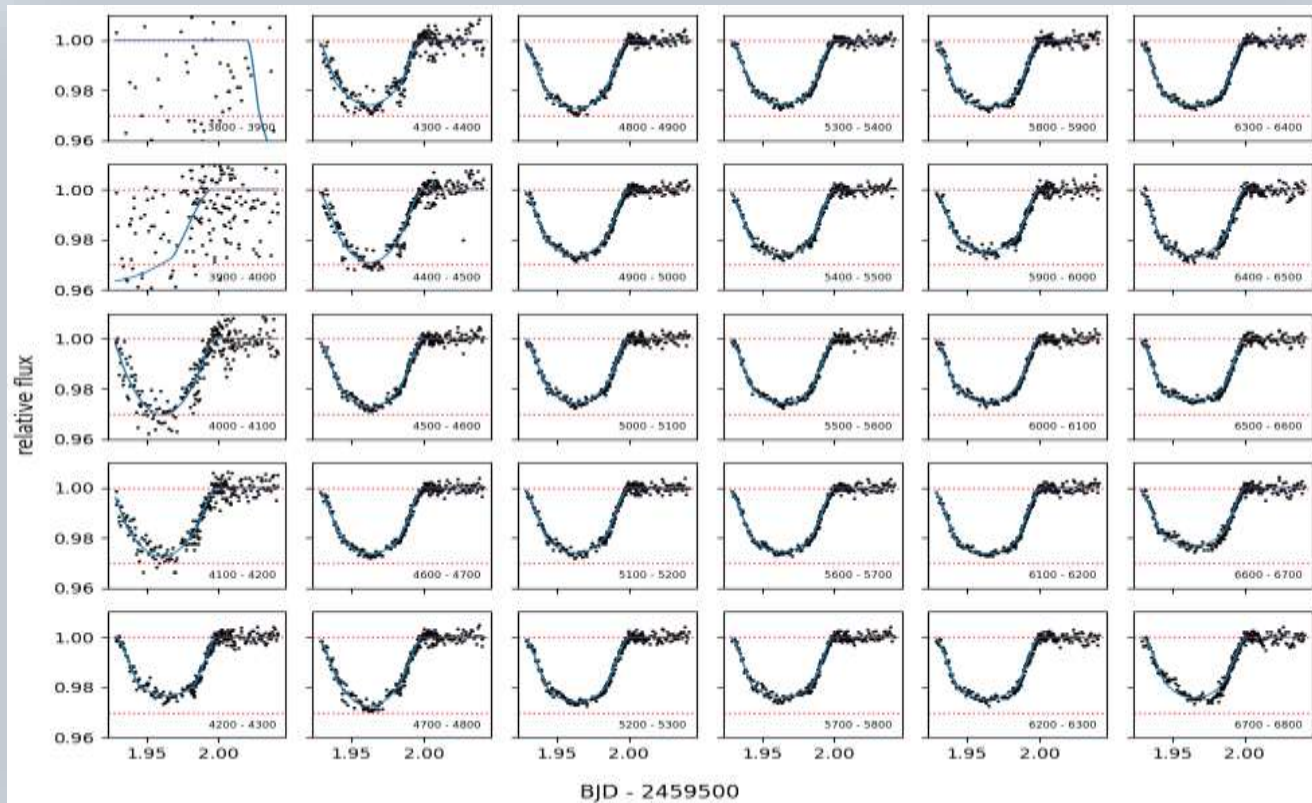


8.7'

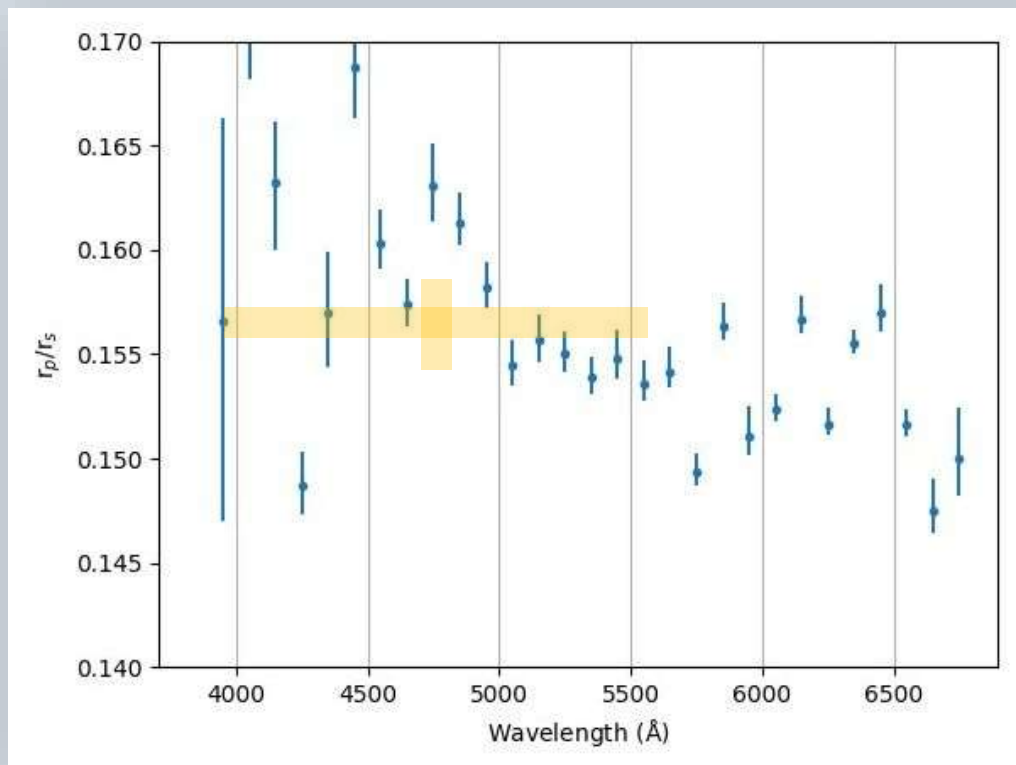


V = 7.7, 8.1

- Transit fits (PyTransit) for each 100 Å band



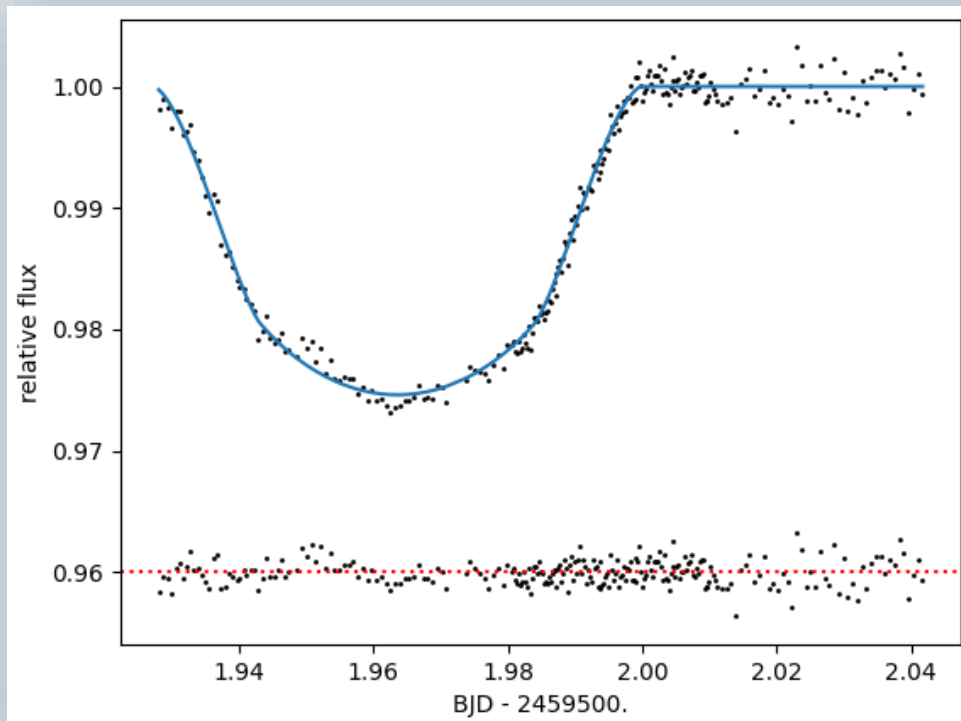
- Planet radius spectrum



$\delta(R_p/R_*) \sim 0.001$

Kasper+ 2019

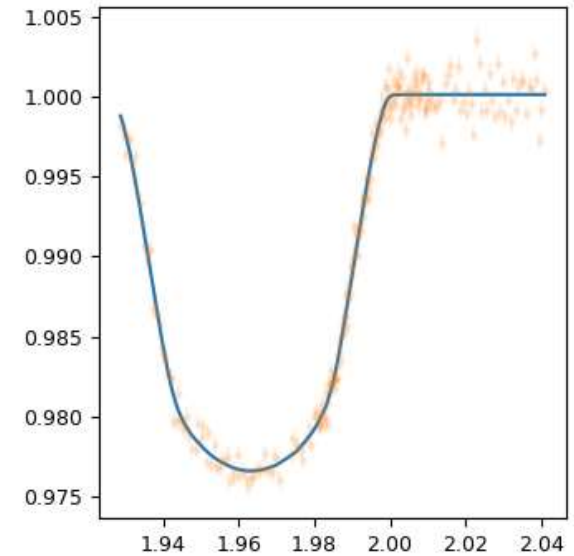
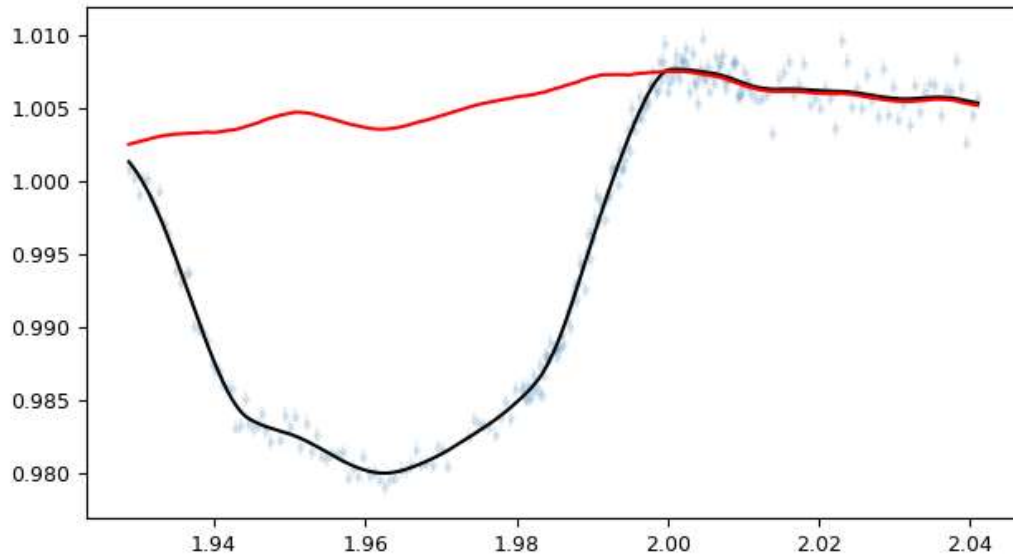
- Photometric accuracy



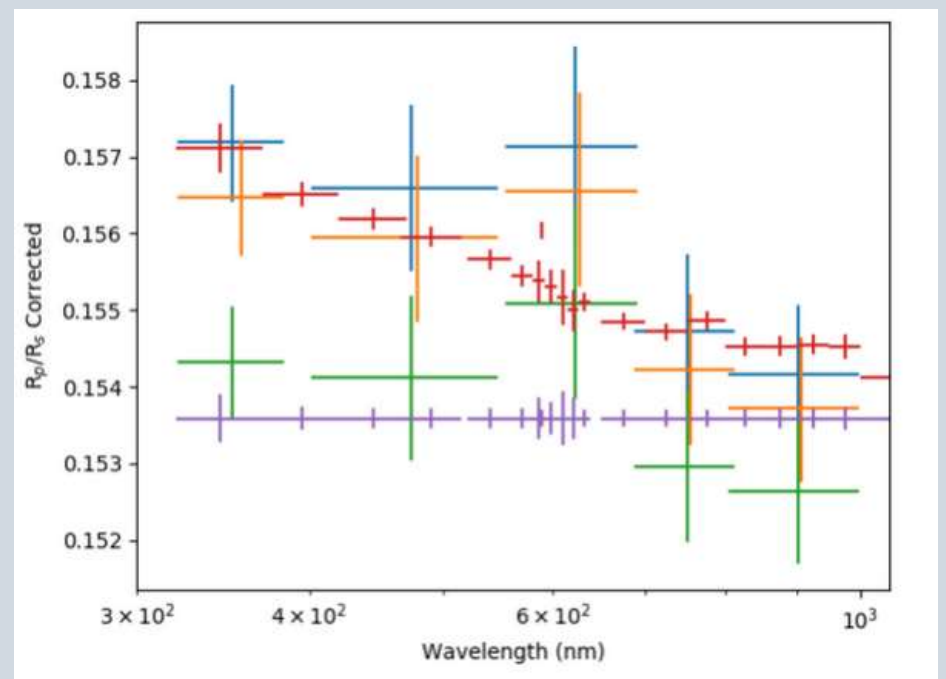
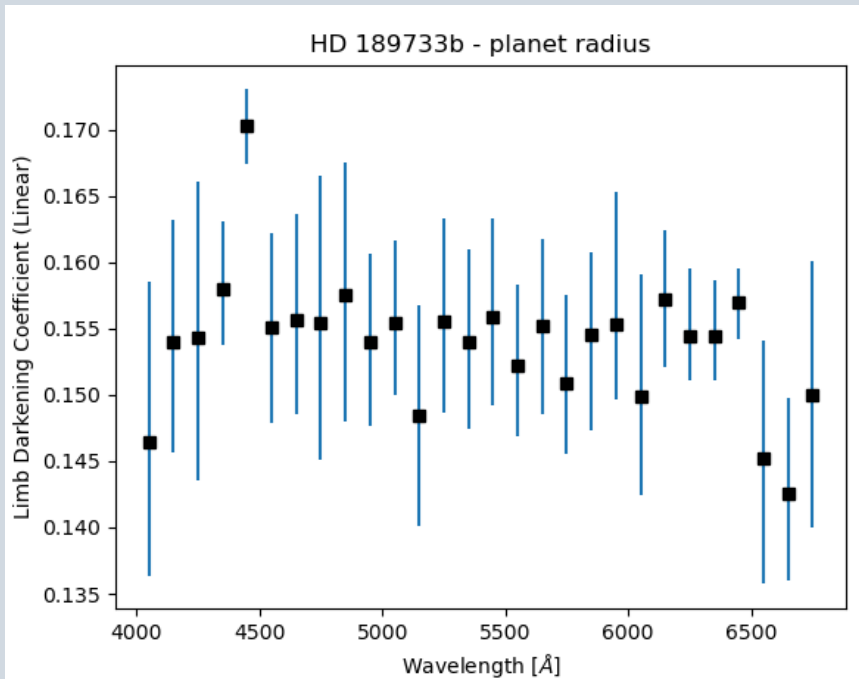
STD of 1.2 mmag

# Gaussian Process

HD 189733



# HD 189733b



Kasper et al. 2019 MNRAS



# Qatar-8

2022.04.08 11:55 - 18:44 (UT)

$V = 11.7$

$a = \text{sma}/r_* = 7.911$

$p = r_p/r_* = 0.100$

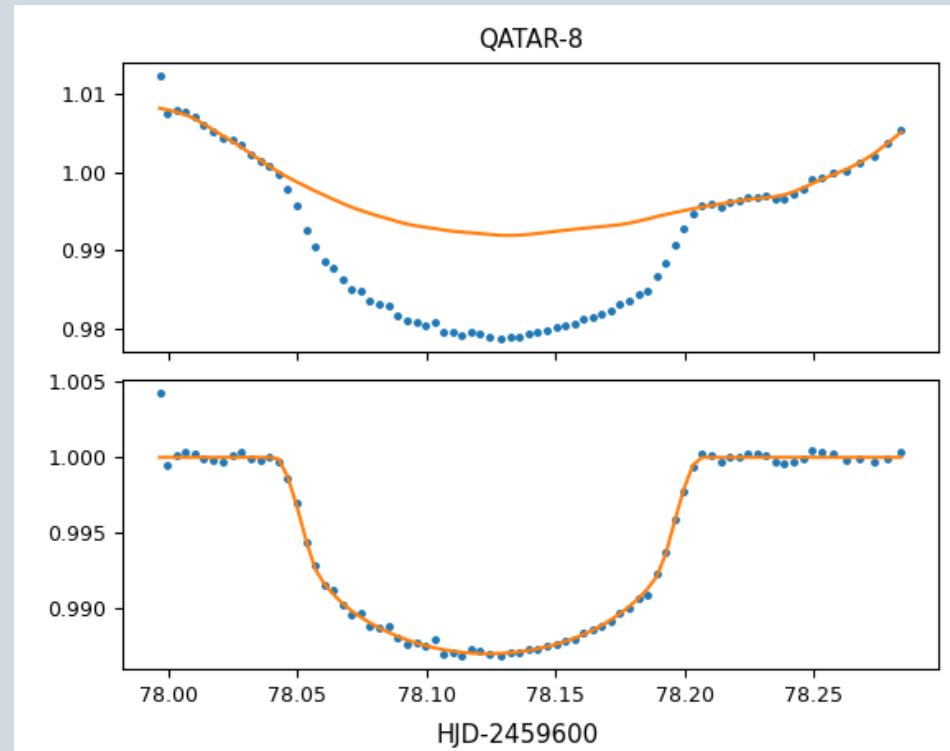
$b = 0.098$

$i = \arccos(b/a) = 89.29^\circ$

## Primary Transit Parameters:

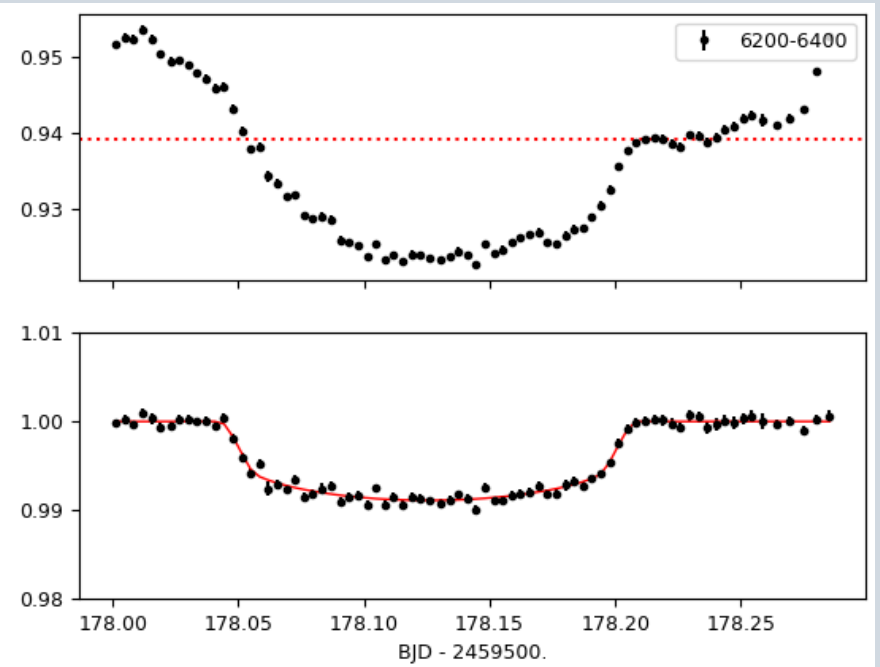
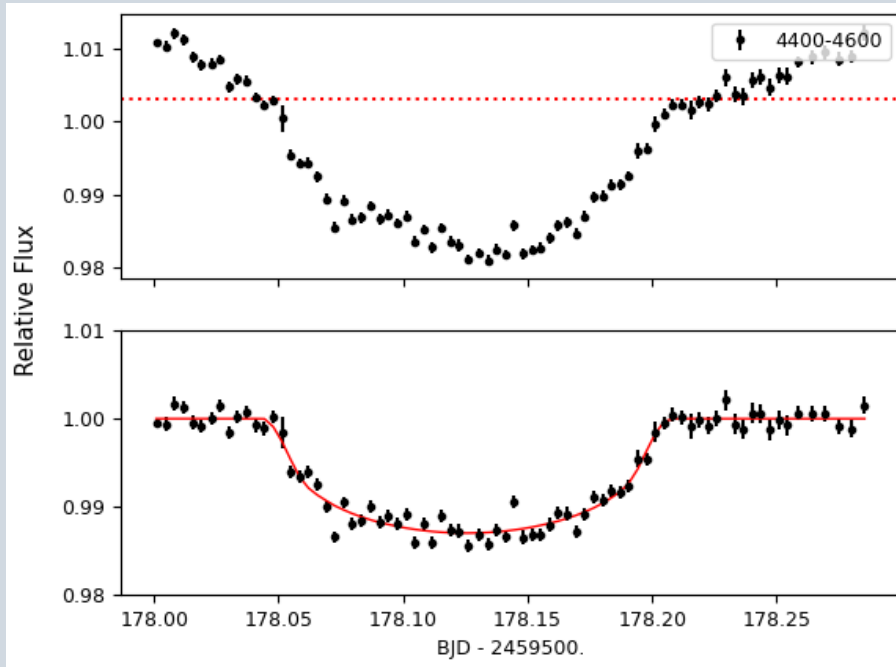
$T_C$	Time of transit (BJD <sub>TDB</sub> )	$2458210.83980 \pm 0.00085$
$R_p/R_*$	Radius of planet in stellar radii	<u><math>0.1005 \pm 0.0008</math></u>
$a/R_*$	Semimajor axis in stellar radii	$7.761 \pm 0.100$
$i$	Inclination ( $^\circ$ )	$89.29 \pm 0.70$
$b$	Impact Parameter	<u><math>0.096 \pm 0.093</math></u>
$T_{14}$	Total duration (days)	$0.1678 \pm 0.0017$

Alsubai et al. 2019



# Qatar-8

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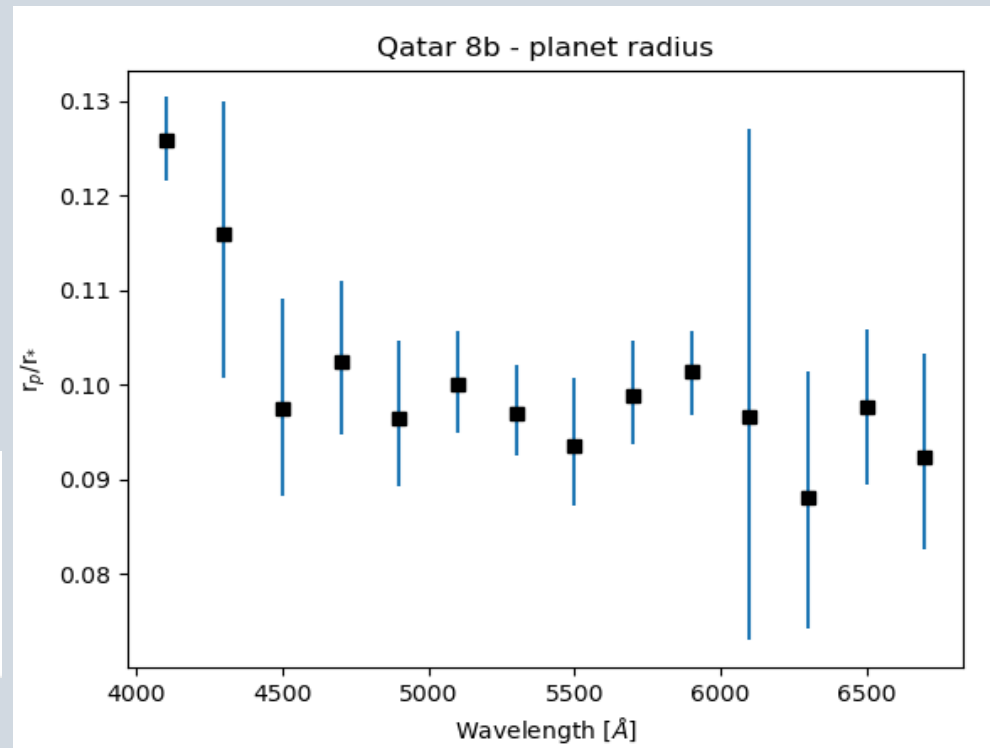


# Qatar-8b

## Primary Transit Parameters:

$T_C$	Time of transit (BJD <sub>TDB</sub> )	$2458210.83980 \pm 0.00085$
$R_p/R_*$	Radius of planet in stellar radii	$0.1005 \pm 0.0008$
$a/R_*$	Semimajor axis in stellar radii	$7.761 \pm 0.100$
$i$	Inclination (°)	$89.29 \pm 0.70$
$b$	Impact Parameter	$0.096 \pm 0.093$
$T_{14}$	Total duration (days)	$0.1678 \pm 0.0017$

Alsubai et al. 2019, AJ





S P E C T R E

# SPECTR

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- Wide FOV, multi-purpose low-resolution spectrophotometer
- Direct view
- FOV: 10 x 4.6 arcmin
- VPH grism
- Wavelength: 380 – 685 nm
- Resolution:  $R > 400$  ( $RS \sim 2000$  [“])
- Slit widths: 1.4“, 2.0“, 2.8“, 4.0“, 6.0“, 9.0“, 15“, 25“
- Andor iKon L936 CCD (2k x 2k)