# Dispersive determination of neutrino mass ordering

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Presented at High 1

Jan. 22, 2024

2306.03463

### Unsolved issues in neutrino physics

- Today's talk will try to answer:
- Neutrino mass ordering

 $\Delta m_{21}^2 \equiv m_2^2 - m_1^2 = (7.55^{+0.20}_{-0.16}) \times 10^{-5} \text{ eV}^2 \qquad \Delta m_{32}^2 \equiv m_3^2 - m_2^2 = (2.424 \pm 0.03) \times 10^{-3} \text{ eV}^2$ 

but normal ordering or inverted ordering?

• Why small mixing in quark sector, but large mixing in lepton sector?

**CKM:**  $\theta_{12} = 13.04^{\circ} \pm 0.05^{\circ}, \ \theta_{13} = 0.201^{\circ} \pm 0.011^{\circ}, \ \theta_{23} = 2.38^{\circ} \pm 0.06^{\circ}$ 

Pontecorvo–Maki–Nakagawa–Sakata:  $\theta_{12} = 33.41^{\circ} + 0.75^{\circ}_{-0.72^{\circ}}$   $\theta_{13} = 8.54^{\circ} + 0.11^{\circ}_{-0.12^{\circ}}$ 

- Why lepton mixing has maximal angle  $\theta_{23} \approx 45^{\circ}$ ?
- Why neutrinos so light?

#### m<sub>3</sub>= EW scale x EW symmetry restoration scale



#### Speculation

- Fundamental parameters in theory (like Standard Model) usually constrained by symmetries at Lagrangian level
- Physical observables are analytical
- Dispersion relations must be respected
- $\Gamma_{12}$  involves CKM matrix elements and fermion masses
- Additional dynamical constraints imposed by dispersion relations, given  $M_{12}\,$  ?
- Turn out that dispersive constraints are so strong that Yukawa couplings in SM are in fact not free parameters

#### Idea

• Neutral state mixing disappears at high energy, where electroweak symmetry is restored (working assumption)

Li, 2304.05921





#### Proof of $M_{12}(s) \approx 0$

- Consider mixing of  $Q_L ar{q}_L$  ,  $ar{Q}_L q_L$  neutral states
- Before breaking, all particles are massless, quarks in flavor eigenstates
- Mixing occurs via exchanges of charged or neutral scalars, whose strengths described by Yukawa matrices



- After breaking, particles get masses, quarks turned to mass eigenstates
- Mixing occurs via W boson exchanges, whose strengths described by CKM matrix

### Mixing in symmetric phase at leading order

- Yukawa interaction  $\overline{Q_L}Y_u u_R \varphi + \overline{Q_L}Y_d d_R \tilde{\varphi} \qquad \varphi = \frac{1}{\sqrt{2}} \begin{pmatrix} \varphi^+ \\ \varphi^0 \end{pmatrix}$ , left-handed doublet
- In symmetric phase, implement quark field transformation adopted in broken phase  $u_L \rightarrow U_u u_L$   $u_R \rightarrow V_u u_R$   $d_L \rightarrow U_d d_L$   $d_R \rightarrow V_d d_R$
- Yukawa matrices diagonalized, but charged scalar currents exist
- down-type quarks, coupling to up-type quarks in mass eigenstates through charged scalar currents, are not in mass eigenstates



#### Box diagrams

- Heavy quark Q provides large s in box diagrams. Symmetry restores and intermediate particles become massless,  $M_{12}(s) \approx 0$
- s' can be low, so  $\Gamma_{12}(s')$  depends on CKM matrix elements associated with massive intermediate quarks in broken phase.

Cheng 1982 Buras et al 1984

$$\begin{split} \Gamma_{12}(s) &\propto -\frac{G_F^2}{16\pi} \sum_{i,j} \lambda_i \lambda_j \Gamma_{ij}(s) \\ \Gamma_{ij}(s) &= \frac{1}{s^2} \frac{\sqrt{s^2 - 2s(m_i^2 + m_j^2) + (m_i^2 - m_j^2)^2}}{(m_W^2 - m_i^2)(m_W^2 - m_j^2)} \\ &\times \left\{ \left( m_W^4 + \frac{m_i^2 m_j^2}{4} \right) [2s^2 - 4s(m_i^2 + m_j^2) + 2(m_i^2 - m_j^2)^2] + 3m_W^2 s(m_i^2 + m_j^2)(m_i^2 + m_j^2 - s) \right\} \end{split}$$

for D mixing i, j = d, s, b  $\lambda_i \equiv V_{ci}^* V_{ui}$ 

#### Constraints

How to diminish dispersive integral

$$\int ds' \frac{\Gamma_{12}(s')}{s-s'} ?$$

Asymptotic expansion

to have finite integral

 $\lambda_i \lambda_j g_{ij} \approx 0$ 

$$\begin{split} \Gamma_{ij}(s') &\approx \Gamma_{ij}^{(1)}s' + \Gamma_{ij}^{(0)} + \frac{\Gamma_{ij}^{(-1)}}{s'} + \cdots & \text{EW symmetry} \\ \text{restoration scale} \\ \Gamma_{ij}^{(1)} &= \frac{4m_W^4 - 6m_W^2(m_i^2 + m_j^2) + 4m_i^2m_j^2}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)}, \quad \clubsuit \Lambda^2/s \\ \Gamma_{ij}^{(0)} &= -\frac{3(m_i^2 + m_j^2)\left[4m_W^4 - 4m_W^2(m_i^2 + m_j^2) + m_i^2m_j^2\right]}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)} \quad \clubsuit (m_i^2 + m_j^2)\Lambda/s \\ \Gamma_{ij}^{(-1)} &= \frac{3(m_i^4 + m_j^4)\left[4m_W^4 - 2m_W^2(m_i^2 + m_j^2) + m_i^2m_j^2\right]}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)}. \quad \clubsuit (m_i^4 + m_j^4)\ln\Lambda/s \\ \Gamma_{ij}^{(-1)} &= \frac{3(m_i^4 + m_j^4)\left[4m_W^4 - 2m_W^2(m_i^2 + m_j^2) + m_i^2m_j^2\right]}{2(m_W^2 - m_i^2)(m_W^2 - m_j^2)}. \quad \clubsuit (m_i^4 + m_j^4)\ln\Lambda/s \\ \text{to diminish integral} \end{split}$$

EW symmetry

$$\int ds' \frac{\Gamma_{12}(s')}{s-s'} \approx \frac{1}{s} \sum_{i,j} \lambda_i \lambda_j g_{ij} \qquad g_{ij} \equiv \int_{t_{ij}}^{\infty} ds' \left[ \Gamma_{ij}(s') - \Gamma_{ij}^{(1)}s' - \Gamma_{ij}^{(0)} - \frac{\Gamma_{ij}^{(-1)}}{s'} \right]$$

#### Minimization

• Use unitarity to eliminate  $\lambda_b$  and to rewrite constrains

$$r^{2}R_{dd}^{(m)} + 2rR_{ds}^{(m)} + 1 \approx 0, \quad m = 1, 0, -1, i$$

$$R_{dd}^{(m)} = \frac{\Gamma_{dd}^{(m)} - 2\Gamma_{db}^{(m)} + \Gamma_{bb}^{(m)}}{\Gamma_{ss}^{(m)} - 2\Gamma_{sb}^{(m)} + \Gamma_{bb}^{(m)}}, \quad R_{ds}^{(m)} = \frac{\Gamma_{ds}^{(m)} - \Gamma_{db}^{(m)} - \Gamma_{sb}^{(m)} + \Gamma_{bb}^{(m)}}{\Gamma_{ss}^{(m)} - 2\Gamma_{sb}^{(m)} + \Gamma_{bb}^{(m)}} \qquad m = 1, 0, -1$$

- Expression for m = i similar, but with  $g_{ij}$
- Ratio of CKM elements  $r = \frac{\lambda_d}{\lambda_s} = \frac{V_{cd}^* V_{ud}}{V_{cs}^* V_{us}} \equiv u + iv,$
- Tune u and v to minimize the sum (real parts of constraints)

$$\sum_{m=1,-1,i} \left[ (u^2 - v^2) R_{dd}^{(m)} + 2u R_{ds}^{(m)} + 1 \right]^2$$

#### **Results** $m_d = 0.005 \text{ GeV}$ $m_s = 0.12 \text{ GeV}$ $m_b = 4.0 \text{ GeV}$ $m_W = 80.377 \text{ GeV}$

#### 3. × 10<sup>-6</sup> $3. \times 10^{-6}$ v=0.00062 v=0 minimum reached 2. × 10<sup>-6</sup> 2. × 10<sup>-6</sup> m=0,-1 1.×10<sup>-6</sup> $1. \times 10^{-6}$ - m=1 -0.9985 -1\_0000 -0 9995 -1.0000-0.9990 -0.9995 U U $-1. \times 10^{-6}$ $-1. \times 10^{-6}$ $-2. \times 10^{-6}$ $-2. \times 10^{-6}$ m=i PDG

 $r = \frac{V_{cd}^* V_{ud}}{V_{cs}^* V_{us}} = -1.0 + (6.2^{+1.2}_{-1.0}) \times 10^{-4} i \qquad u = -1.00029 \pm 0.00002, \qquad v = 0.00064 \pm 0.00002$ variation of ms by 0.01 GeV they agree well

#### Pontecorvo–Maki–Nakagawa–Sakata matrix

$$\begin{bmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{bmatrix} \begin{bmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{13}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{13}} & 0 & c_{13} \end{bmatrix} \begin{bmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{bmatrix}$$
 Chau-Keung  
= 
$$\begin{bmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{13}} \\ -s_{12}c_{23} - c_{12}s_{23}s_{13}e^{i\delta_{13}} & c_{12}c_{23} - s_{12}s_{23}s_{13}e^{i\delta_{13}} & s_{23}c_{13} \\ s_{12}s_{23} - c_{12}c_{23}s_{13}e^{i\delta_{13}} & -c_{12}s_{23} - s_{12}c_{23}s_{13}e^{i\delta_{13}} & c_{23}c_{13} \end{bmatrix}.$$

#### PDG



	Ref. $[188]$ w/o SK-ATM		Ref. [188] w SK-ATM		Ref. [189] w SK-ATM		Ref. [190] w SK-ATM	
NO	Best Fit Ordering		Best Fit Ordering		Best Fit Ordering		Best Fit Ordering	
Faram	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range	bfp $\pm 1\sigma$	$3\sigma$ range
$\frac{\sin^2 \theta_{12}}{10^{-1}}$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.04^{+0.14}_{-0.13}$	$2.65 \rightarrow 3.46$	$3.20^{+0.20}_{-0.16}$	$2.73 \rightarrow 3.79$
$\theta_{12}/^{\circ}$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.46^{+0.87}_{-0.88}$	$30.98 \rightarrow 36.03$	$34.5^{+1.2}_{-1.0}$	$31.5 \rightarrow 38.0$
$\frac{\sin^2 \theta_{23}}{10^{-1}}$	$5.58^{+0.20}_{-0.33}$	$4.27 \rightarrow 6.09$	$5.63^{+0.18}_{-0.24}$	$4.33 \rightarrow 6.09$	$5.51^{+0.19}_{-0.80}$	$4.30 \rightarrow 6.02$	$5.47^{+0.20}_{-0.30}$	$4.45 \rightarrow 5.99$
$\theta_{23}/^{\circ}$	$48.3^{+1.2}_{-1.9}$	$40.8 \rightarrow 51.3$	$48.6^{+1.0}_{-1.4}$	$41.1 \rightarrow 51.3$	$47.9^{+1.1}_{-4.0}$	$41.0 \rightarrow 50.9$	$47.7^{+1.2}_{-1.7}$	$41.8 \rightarrow 50.7$
$\frac{\sin^2 \theta_{13}}{10^{-2}}$	$2.241^{+0.066}_{-0.065}$	$2.046 \rightarrow 2.440$	$2.237^{+0.066}_{-0.065}$	$2.044 \rightarrow 2.435$	$2.14^{+0.09}_{-0.07}$	$1.90 \rightarrow 2.39$	$2.160^{+0.083}_{-0.069}$	$1.96 \rightarrow 2.41$
$\theta_{13}/^{\circ}$	$8.61^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.99$	$8.60^{+0.13}_{-0.13}$	$8.22 \rightarrow 8.98$	$8.41^{+0.18}_{-0.14}$	$7.9 \rightarrow 8.9$	$8.45_{-0.14}^{+0.16}$	$8.0 \rightarrow 8.9$
$\delta_{\rm CP}/^{\circ}$	$222_{-28}^{+38}$	$141 \rightarrow 370$	$221_{-28}^{+39}$	$144 \rightarrow 357$	$238^{+41}_{-33}$	$149 \rightarrow 358$	$218^{+38}_{-27}$	$157 \rightarrow 349$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.34_{-0.14}^{+0.17}$	$6.92 \rightarrow 7.91$	$7.55_{-0.16}^{+0.20}$	$7.05 \rightarrow 8.24$
$\frac{\Delta m_{32}^2}{10^{-3} \text{ eV}^2}$	$2.449^{+0.032}_{-0.030}$	$2.358 \rightarrow 2.544$	$2.454_{-0.031}^{+0.029}$	$2.362 \rightarrow 2.544$	$2.419^{+0.035}_{-0.032}$	$2.319 \rightarrow 2.521$	$2.424 \pm 0.03$	$2.334 \rightarrow 2.524$
IO	$\Delta \chi^2 = 6.2$		$\Delta \chi^2 = 10.4$		$\Delta \chi^2 = 9.5$		$\Delta \chi^2 = 11.7$	
$\frac{\sin^2 \theta_{12}}{10^{-1}}$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.10^{+0.13}_{-0.12}$	$2.75 \rightarrow 3.50$	$3.03^{+0.14}_{-0.13}$	$2.64 \rightarrow 3.45$	$3.20^{+0.20}_{-0.16}$	$2.73 \rightarrow 3.79$
$\theta_{12}/^{\circ}$	$33.82^{+0.78}_{-0.76}$	$31.61 \rightarrow 36.27$	$33.82^{+0.78}_{-0.75}$	$31.62 \rightarrow 36.27$	$33.40^{+0.87}_{-0.81}$	$30.92 \rightarrow 35.97$	$34.5^{+1.2}_{-1.0}$	$31.5 \rightarrow 38.0$
$\frac{\sin^2 \theta_{23}}{10^{-1}}$	$5.63^{+0.19}_{-0.26}$	$4.30 \rightarrow 6.12$	$5.65^{+0.17}_{-0.22}$	$4.36 \rightarrow 6.10$	$5.57^{+0.17}_{-0.24}$	$4.44 \rightarrow 6.03$	$5.51^{+0.18}_{-0.30}$	$4.53 \rightarrow 5.98$
$\theta_{23}/^{\circ}$	$48.6^{+1.1}_{-1.5}$	$41.0 \rightarrow 51.5$	$48.8^{+1.0}_{-1.2}$	$41.4 \rightarrow 51.3$	$48.2^{+1.0}_{-1.4}$	$41.8 \rightarrow 50.9$	$47.9^{+1.0}_{-1.7}$	$42.3 \rightarrow 50.7$
$\frac{\sin^2 \theta_{13}}{10^{-2}}$	$2.261^{+0.067}_{-0.064}$	$2.066 \rightarrow 2.461$	$2.259^{+0.065}_{-0.065}$	$2.064 \rightarrow 2.457$	$2.18^{+0.08}_{-0.07}$	$1.95 \rightarrow 2.43$	$2.220^{+0.074}_{-0.076}$	$1.99 \rightarrow 2.44$
$\theta_{13}/^{\circ}$	$8.65^{+0.13}_{-0.12}$	$8.26 \rightarrow 9.02$	$8.64^{+0.12}_{-0.13}$	$8.26 \rightarrow 9.02$	$8.49^{+0.15}_{-0.14}$	$8.0 \rightarrow 9.0$	$8.53^{+0.14}_{-0.15}$	$8.1 \rightarrow 9.0$
$\delta_{\rm CP}/^{\circ}$	$285^{+24}_{-26}$	$205 \rightarrow 354$	$282^{+23}_{-25}$	$205 \rightarrow 348$	$247^{+26}_{-27}$	$193 \rightarrow 346$	$281^{+23}_{-27}$	$202 \rightarrow 349$
$\frac{\Delta m_{21}^2}{10^{-5} \text{ eV}^2}$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.39^{+0.21}_{-0.20}$	$6.79 \rightarrow 8.01$	$7.34_{-0.14}^{+0.17}$	$6.92 \rightarrow 7.91$	$7.55_{-0.16}^{+0.20}$	$7.05 \rightarrow 8.24$
$\frac{\Delta m_{32}^2}{10^{-3} \text{ eV}^2}$	$-2.509^{+0.032}_{-0.032}$	$-2.603 \rightarrow -2.416$	$-2.510^{+0.030}_{-0.031}$	$-2.601 \rightarrow -2.419$	$-2.478^{+0.035}_{-0.033}$	$-2.577 \rightarrow -2.375$	$-2.50\pm^{+0.04}_{-0.03}$	$-2.59 \rightarrow -2.39$

#### Lepton mixing

- Apply the same formalism to lepton  $\mu^-e^+-\mu^+e^-$  mixing through similar box diagrams with intermediate neutrino channels **PMNS**
- Correspondence  $m_{d,s,b} \leftrightarrow m_{1,2,3}$   $V_{cd}^*V_{ud}/(V_{cs}^*V_{us}) \leftrightarrow r = U_{\mu 1}^*U_{e1}/(U_{\mu 2}^*U_{e2})$
- Normal hierarchy (NH)  $m_1^2 = 10^{-6} \text{ eV}^2$  de Salas et al, 2018

 $\Delta m_{21}^2 \equiv m_2^2 - m_1^2 = (7.55^{+0.20}_{-0.16}) \times 10^{-5} \text{ eV}^2 \qquad \Delta m_{32}^2 \equiv m_3^2 - m_2^2 = (2.424 \pm 0.03) \times 10^{-3} \text{ eV}^2$ 

• Predict LO analysis so far

$$r = \frac{U_{\mu 1}^* U_{e1}}{U_{\mu 2}^* U_{e2}} \approx -1.0 - 0.02i$$

$$r = -(0.738_{-0.048}^{+0.050}) - (0.179_{-0.125}^{+0.136})i$$

- Inverted hierarchy (IH)  $r \approx -1.0 O(10^{-5})i$   $r = -(1.03^{+0.05}_{-0.16}) (0.356^{+0.015}_{-0.048})i$
- NH and observed PMNS matrix satisfy constraint at order of magnitude

## Why neutrinos so light?

So far, connections between mixing angles and mass ratios Haven't addressed absolute neutrino masses

alternative to see-saw mechanism

### NLO in symmetric phase

• Mixing disappears above EW restoration scale only at LO, in fact



does not vanish, and determines smallness of neutrino masses

#### Estimate m<sub>3</sub>

• Consider 2-3 generation mixing to reach maximal neutrino mass

50 TeV, between b' mass 2.7 TeV and t' mass 200 TeV

Li, 2309.15602

measured value 
$$\Delta m^2_{32} \equiv m^2_3 - m^2_2 = (2.424 \pm 0.03) \times 10^{-3} \text{ eV}^2$$

Analyticity dictates scalar sector and explains SM flavor structure?