

K-DRIFT: KASI Deep Rolling Imaging Fast Telescope

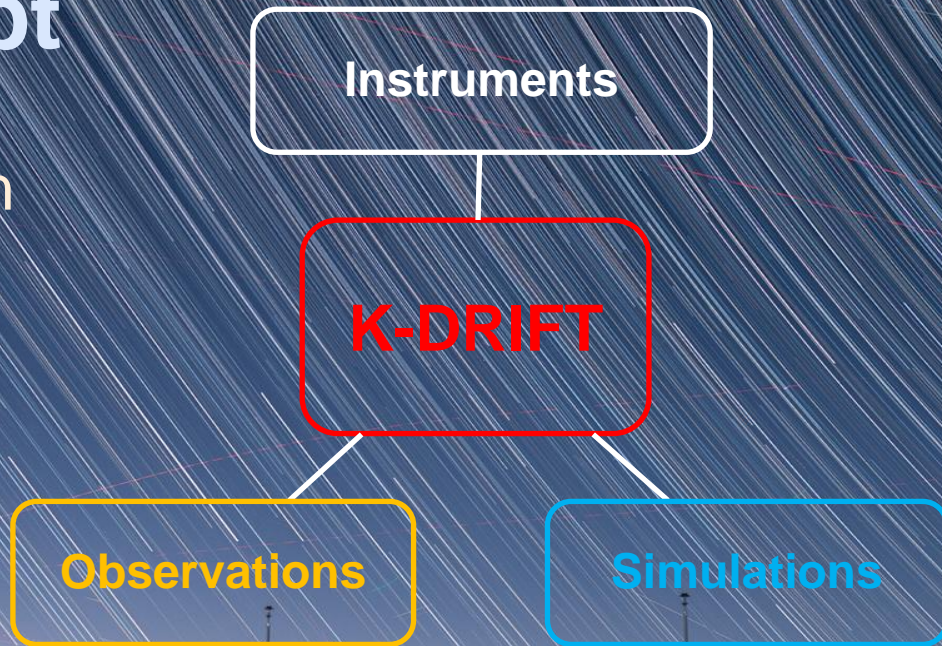
K-DRIFT update: 2024 project status report

고종완(Jongwan Ko) & K-DRIFT Team

Korea Astronomy and Space Science Institute

Basic Concept

Exploring the LSB
universe with our own
cutting-edge
telescopes



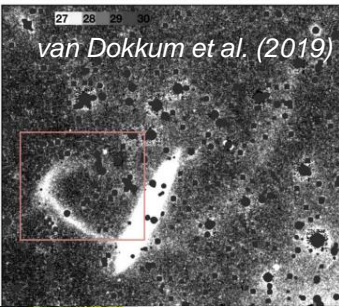
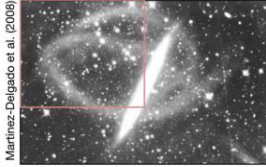
Outline

- K-DRIFT project status report
- K-DRIFT pathfinder upgrade
- K-DRIFT 1st (main survey) preparation

Motivation: why exploring the LSB Universe

Observations

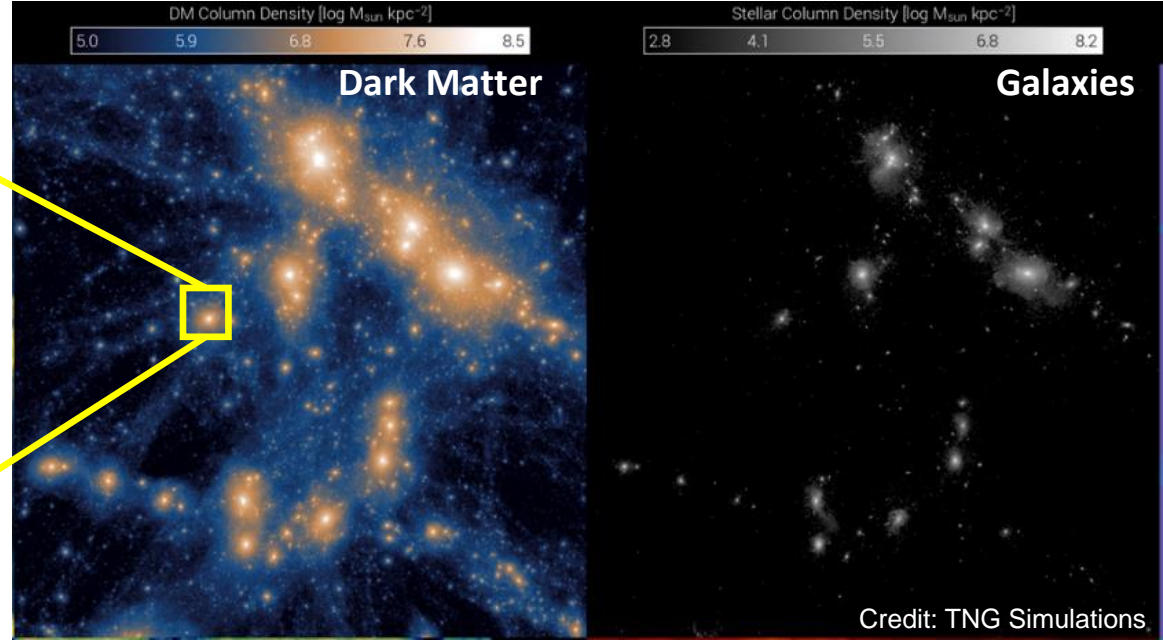
K-DRIFT Pathfinder:
Byun, Ko et al.
(2022)



Martinez-Delgado et al. (2008)



Simulations



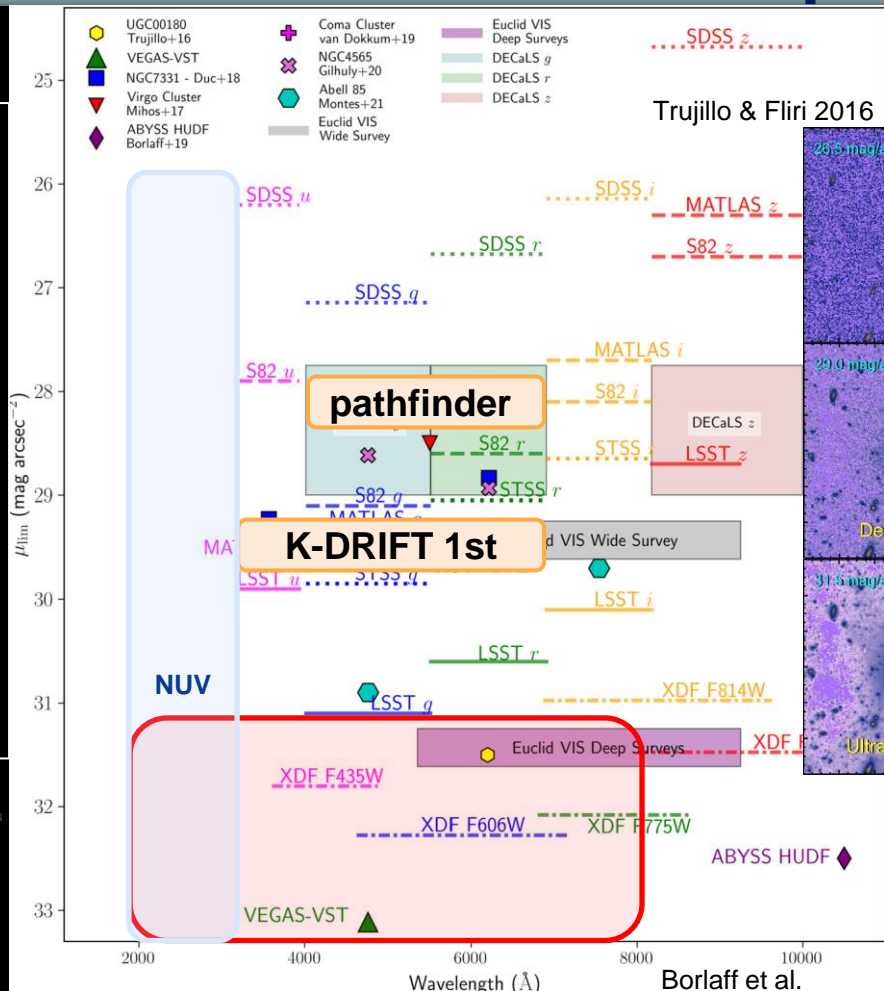
ΛCDM ⇒ all galaxies are surrounded by **vast and complex network of LSB filaments**

But, LSB regime (*generally, <0.1 % of the night sky level*) that is undetectable in past wide-area surveys

Science & technical challenges ⇒ **A step change in our understanding of the Universe!**

K-DRIFT to Space: why space?

SB limit (3σ , $10 \times 10 \text{ arcsec}^2$)



night sky level $\sim 22.0 \text{ mag arcsec}^{-2}$

$(\sim 27.0 \text{ mag arcsec}^{-2})$

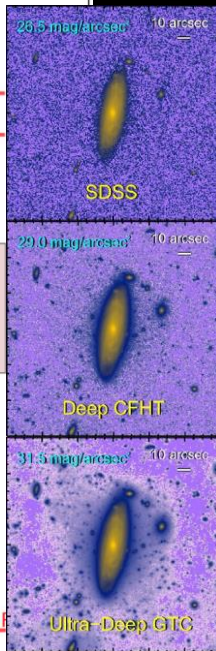
$(\sim 29.0 \text{ mag arcsec}^{-2})$

LSB universe will be explored with Rubin Obs., Euclid, ... K-DRIFT

$(> 31.0 \text{ mag arcsec}^{-2})$

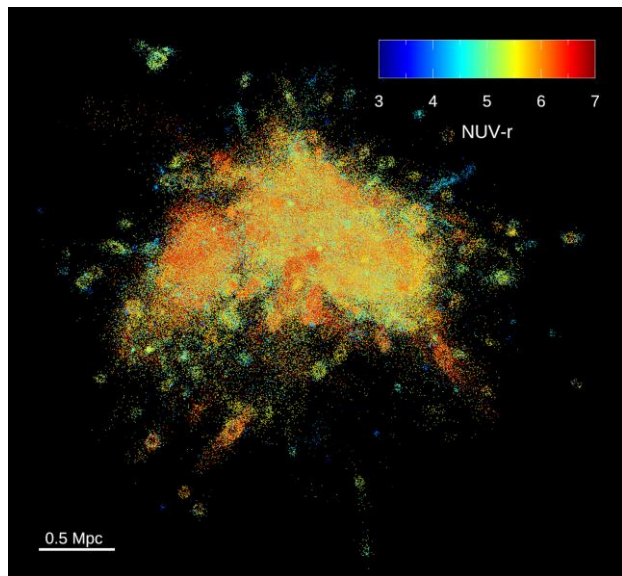
Ultra-LSB universe is almost unexplored!

Space telescope!

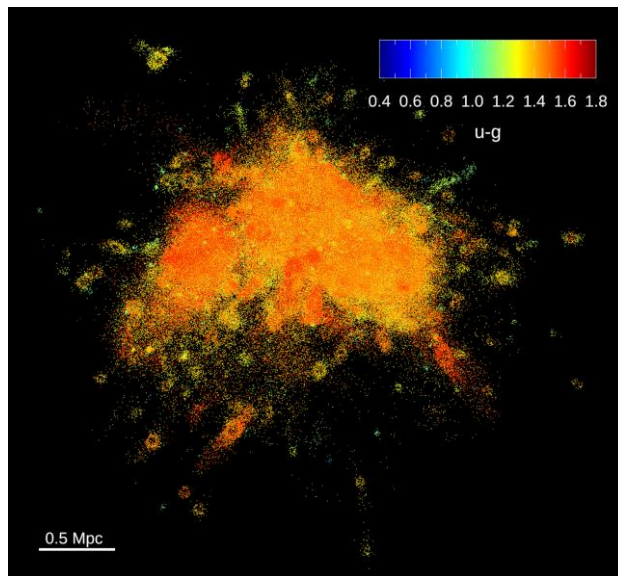


Case 1: ICL

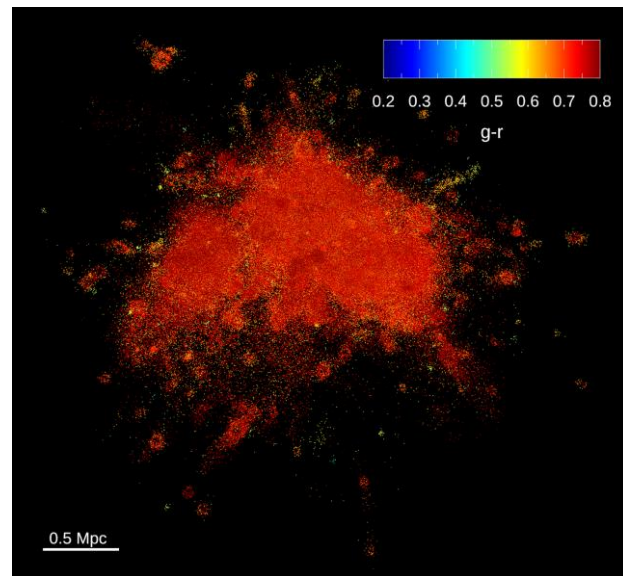
NUV-r



u-g


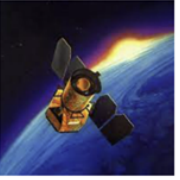
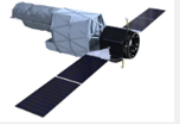



g-r

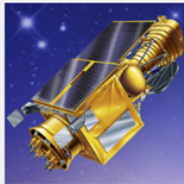
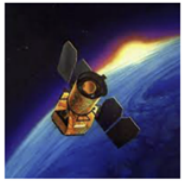
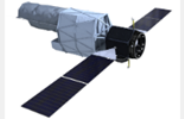



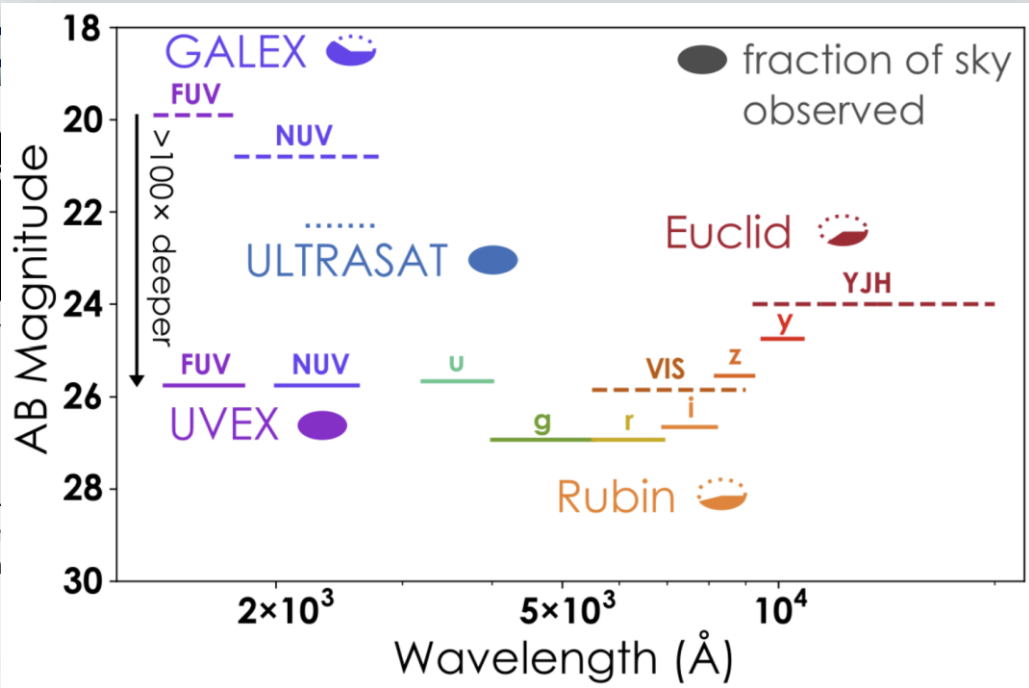
Color gradient + recent SFA → ICL origin

UV Mission Comparison

	MISSION	UV INSTRUMENT APERTURE DIAMETER	BANDPASS IN THE UV REGION	POINT SPREAD FUNCTION	DETECTOR	FIELD OF VIEW	LAUNCH DATE
	ULTRASAT Ultraviolet Transient Astronomy Satellite	0.33 m	230-290 nm	8.3 arcseconds	CMOS	204 degrees ²	2026 (planned)
	GALEX Galaxy Evolution Explorer	0.5 m	170-280 nm 135-175 nm	5.5 arcseconds	MCP	1.13 degrees ²	Launched April 28, 2003 (Decommissioned June 28, 2013)
	UVEX The Ultraviolet Explorer	0.75 m	139-190 nm and 203-270 nm	2.25 arcseconds	CMOS	12.25 degrees ²	2028 (pending downselect)
	STAR-X Survey and Time-domain Astrophysical Research Explorer	0.3 m	160-200 nm and 250-300nm	4.0 arcseconds	MCP	1.0 degrees ²	2028 (pending downselect)

UV Missions

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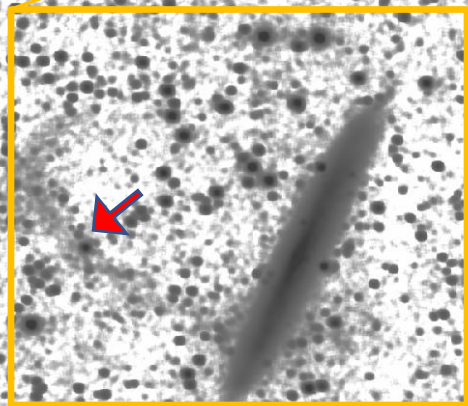
K-DRIFT from 2019 to Present

2019.03.01: 과제 시작
2020.01.07: 광학 설계 최종 결정 (LAF-TMS)
2021.04.19: K-DRIFT pathfinder's first light @본원
2021.05.21: BOAO 이송
2021.06.01-15: BOAO 시험관측, NGC 5907 은하 관측
2021.10.13: K-DRIFT 1st 특별세션 @2021 KAS Fall Meeting
2022.05.17: pathfinder 본원 이송
2022.09.01: M2 교체+재정렬
2022.11.15: 시험관측 @본원
2022.11.30: BOAO로 이송
2022.12.19: BOAO 시험관측 (M2 교체+재정렬 후)
2023.03.21: 가대 교체 (PlaneWave L600)
2023.05.14: 남쪽 돔 관측 준비 완료
2023.05.15~: 원격관측 & 스카이 모니터링 시스템 구축 시작
2023.10.04~: 관측 재개
2023.10.18: K-DRIFT 2nd 특별세션 @2023 KAS Fall Meeting
2023.10.25: K-DRIFT 1st 특별세션 @2023 KSSS Fall Meeting
2023.12.01: 스카이 모니터링 시스템 1차 구축(SB map)

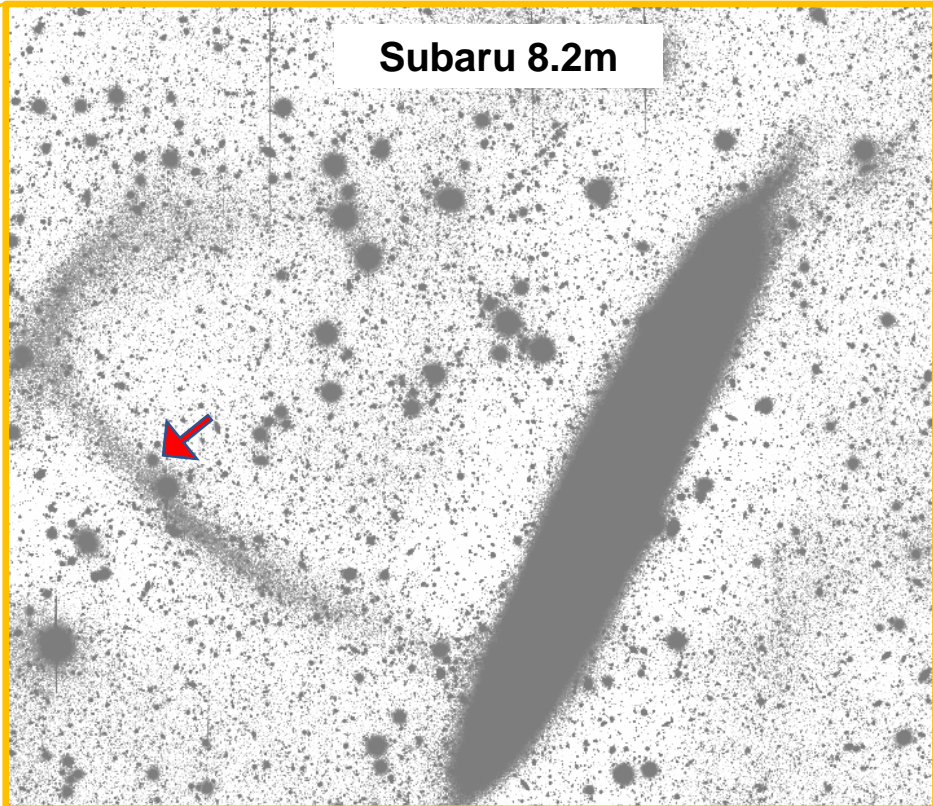


K-DRIFT Dome @BOAO

K-DRIFT Pathfinder 0.3m



Subaru 8.2m



K-DRIFT Pathfinder

Diameter: 0.3m

Exp. Time: 6000s



Subaru

Diameter: 8.2m

Exp. time: 1000s

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K-DRIFT Dome @BOAO



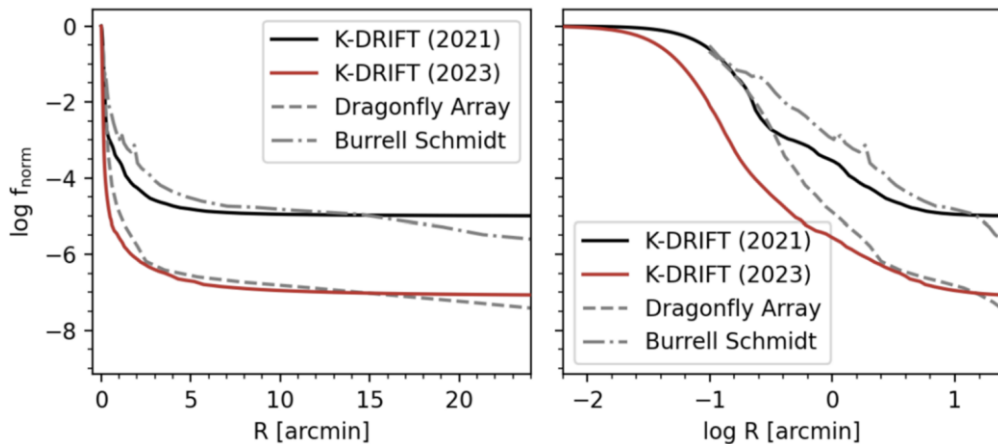
2022.12.15



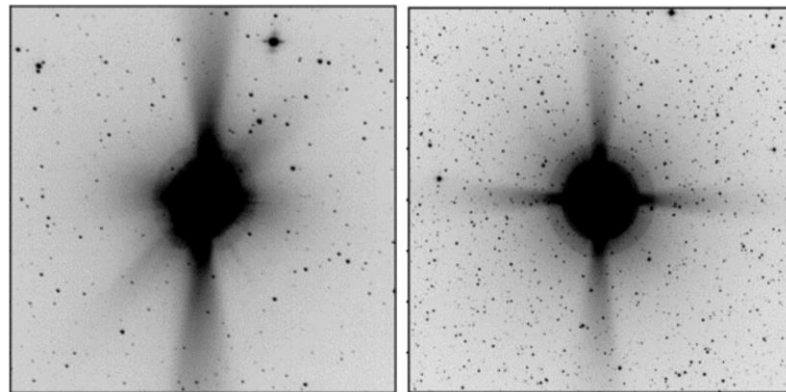
2023.03.21

K-DRIFT pathfinder upgrade: error control for off-axis freeform three-mirror system

PSF profiles



Observed bright stars in 2021 vs 2023



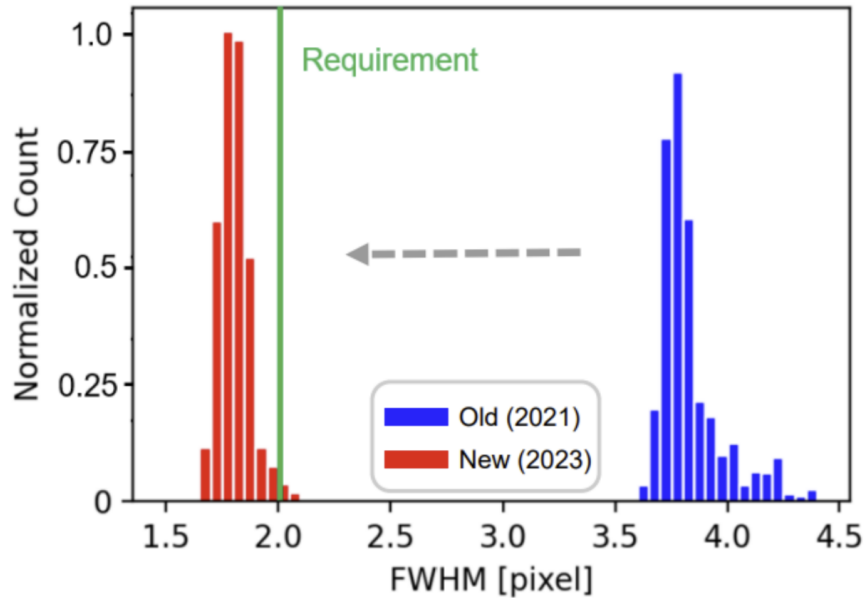
We have developed K-DRIFT's optical performance analysis algorithm!

K-DRIFT pathfinder upgrade

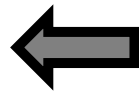
M2 replacing + Realigning + Mount replacing

On-sky test observations: 2022.11.15 @KASI & 2022.12.14, **2023.03.21 @BOAO**

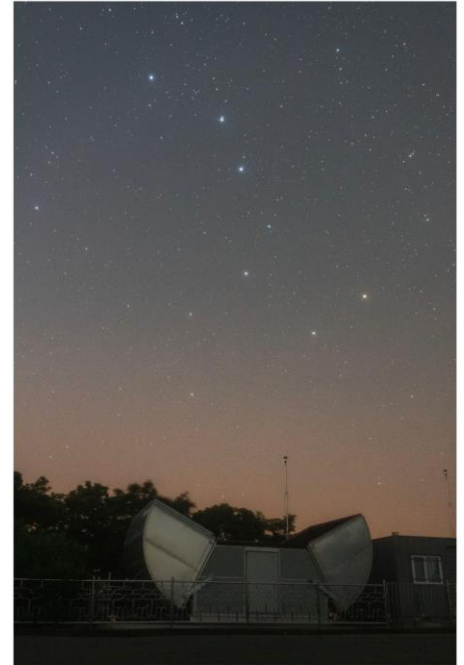
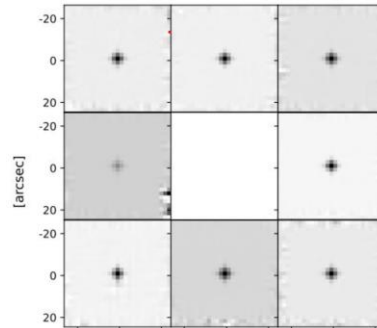
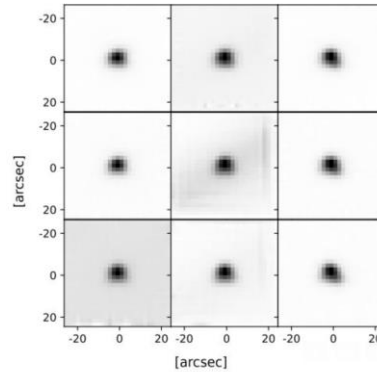
PSF FWHM



1.8 pixel

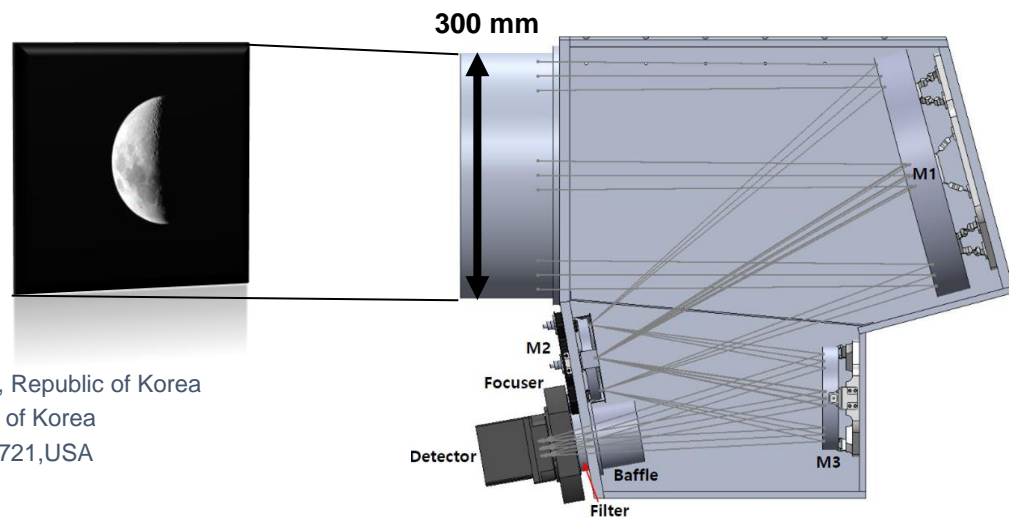


3.8 pixel



On-sky Point Spread Function Analysis for the Off-axis Freeform Three-mirror Telescope

Gayoung Lee^{1,2}, Yunjong Kim², Daewook Kim³, Woowon Byun², Seunghyuk Chang⁴
and Jongwan Ko^{2,5}



¹ Department of Astronomy, Kyungpook National University, Daegu 41566, Republic of Korea

² Korea Astronomy and Space Science Institute, Daejeon 34055, Republic of Korea

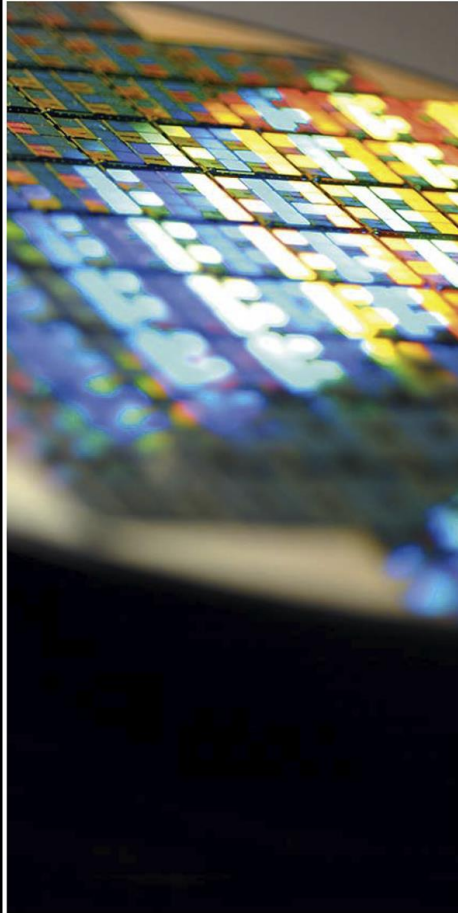
³ Wyant College of Optical Sciences, University of Arizona, Tucson, AZ 85721, USA

⁴ Center for Integrated Smart Sensors, Daejeon 34141, Republic of Korea

⁵ University of Science and Technology, Daejeon 34055, Republic of Korea

**On-sky
for the**

**Gayoung Lee^{1,2},
and Jongwan Ko**



Student Paper Award

Optical Fabrication and Testing (OF&T)

Presented to

Gayoung Lee

Kyungpook National University, Korea
Astronomy and Space Science Institute,
Republic of Korea

For the paper

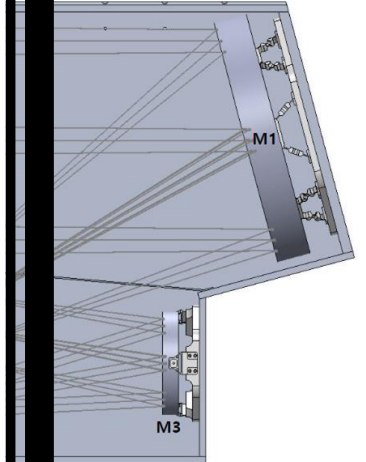
**On-Sky Point Spread Function Analysis
for the Off-Axis Freeform Three-Mirror
Telescope**

Optica Design and Fabrication
Congress

04 - 08 June 2023

Québec City, Québec, Canada

OPTICA | Formerly **OSA**



¹ Department of Astr

² Korea Astronomy a

³ Wyant College of Opt

⁴ Center for Integrat

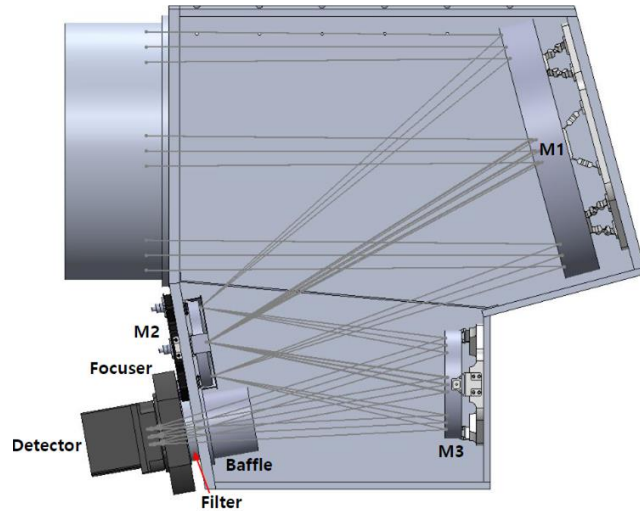
⁵ University of Scien

K-DRIFT pathfinder upgrade:

stray light control for off-axis freeform three-mirror system

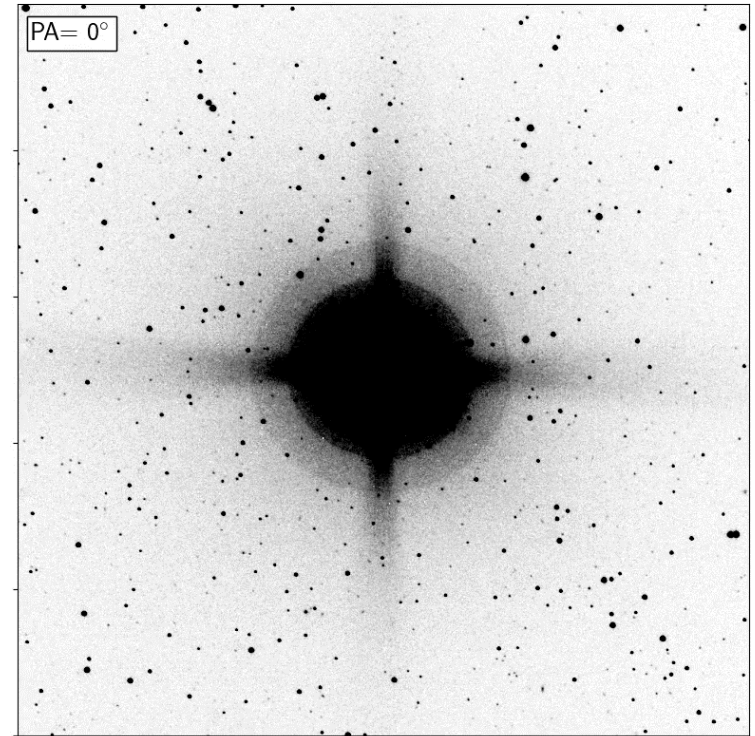
Camera rotation test:

position angle is rotating by $\sim 30^\circ$



We have developed K-DRIFT's
stray light analysis algorithm!

Observed bright star

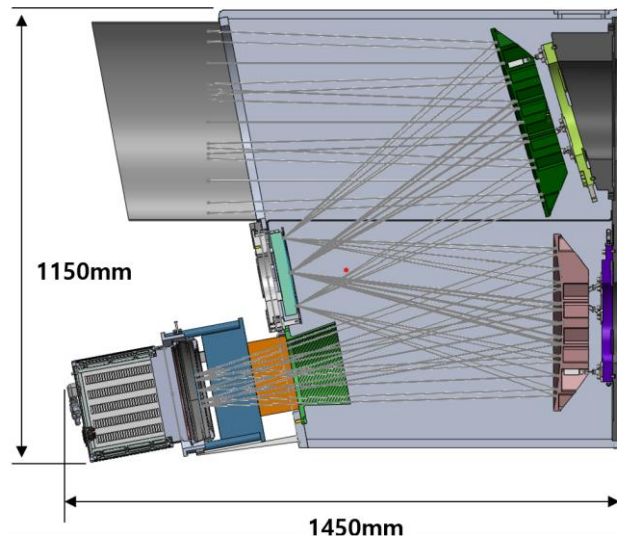
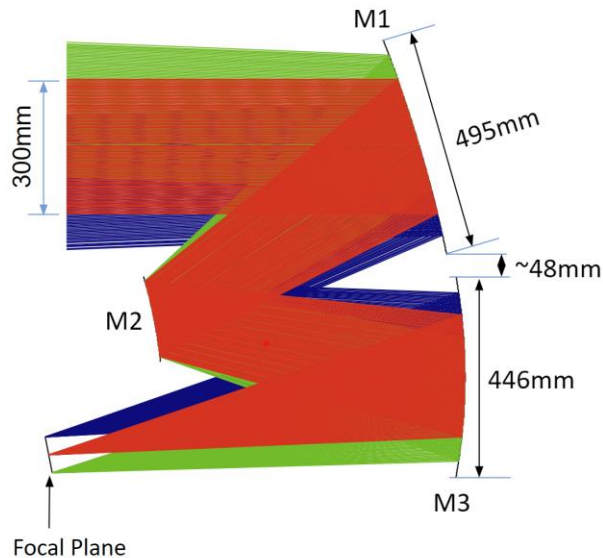


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K-DRIFT 1st (G1) Optical Design

LAF-TMS (Linear-Astigmatism-Free Three-Mirror System)



Parameter	Measurement
Aperture/ Entrance pupil diameter	300/500mm
Focal Length	1050mm
Focal Ratio	3.5
Field of View	4.43°(H) x 4.43°(V)
Image Area	81.2mm x 81.2mm
Pixel Scale	1.96" / 10μm

(large FoV, homogeneous PSF!)

Optical design:
Seunghyuk Chang

19.6 deg² vs 9.6 deg² (Rubin Obs.)

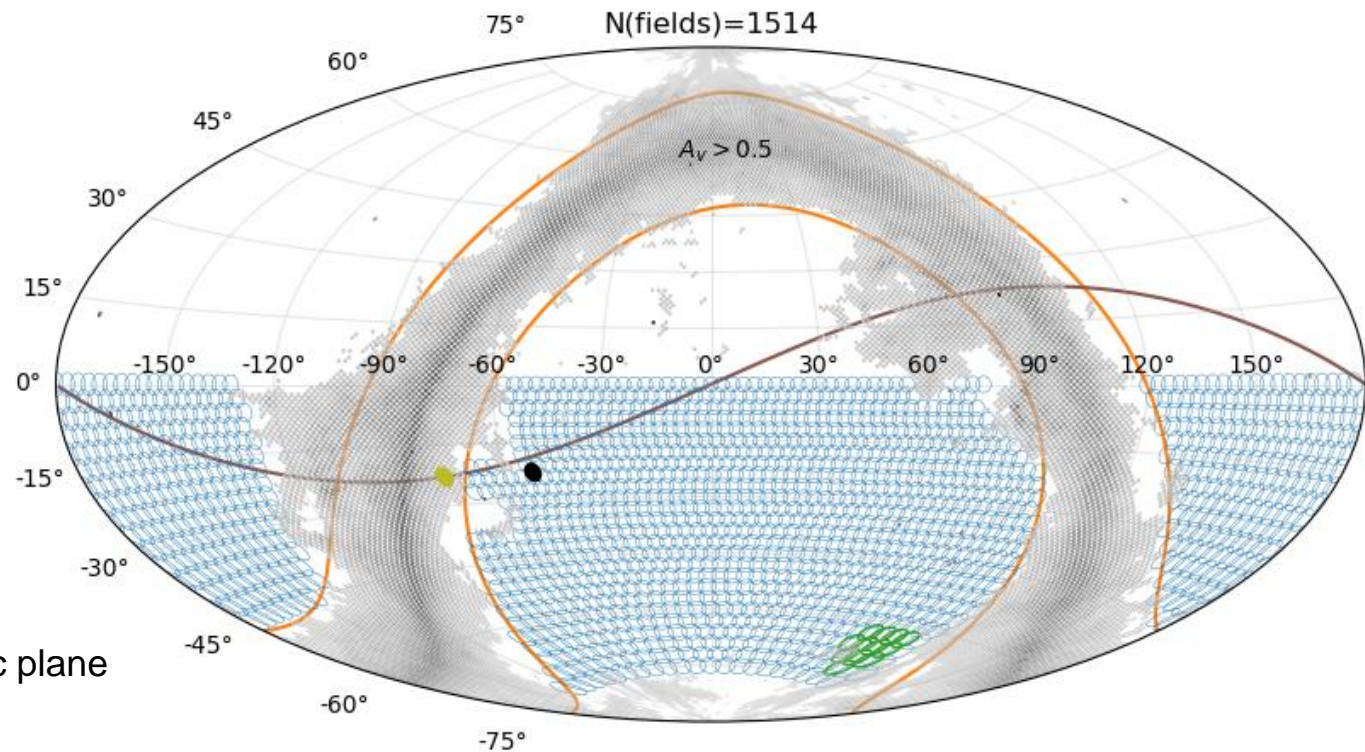
보름달 ~81개 크기!



Southern sky survey (2025~)

- >29.5 mag/arcsec² in u, g, r
- + narrow bands (H α , OIII, ...)

K-DRIFT 1st (G1) survey plan: rolling-dithering



Blue: the rolling-dithering fields

Orange: $\pm 15^\circ$ from the Galactic plane

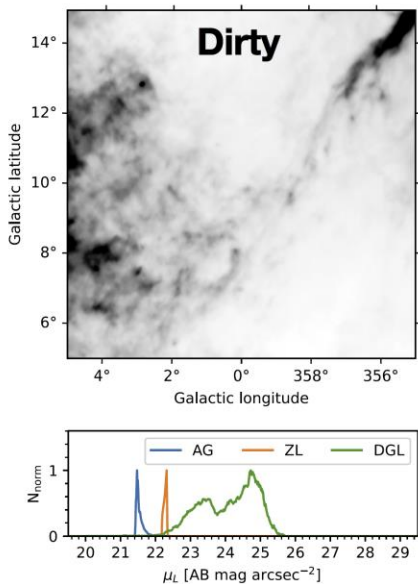
Brown: the ecliptic plane

Green: the SEP regions

Yellow (black): the sun (the moon)

Grayscale: Galactic cirrus ($A_V > 0.5$)

Survey preparation: sky simulation & rolling dithering



*“However, a dark-sky flat may not be perfect because it contains **airglow**, **zodiacal light**, and **diffuse galactic light**.”*

How can we remove these effects?

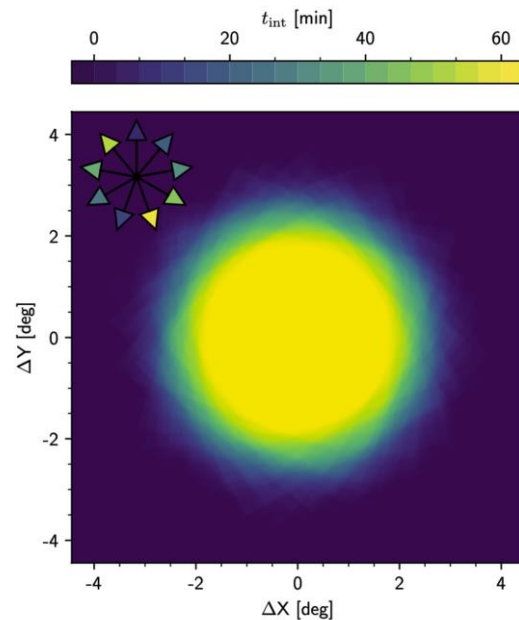
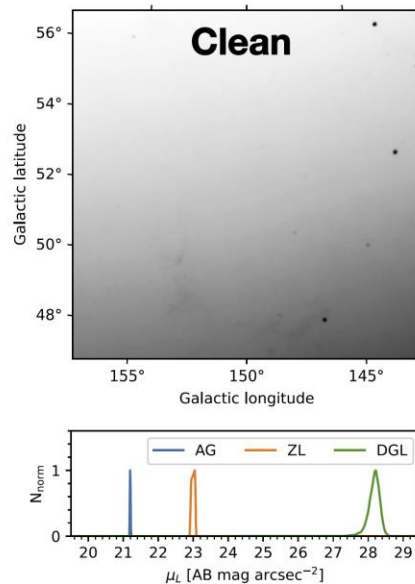
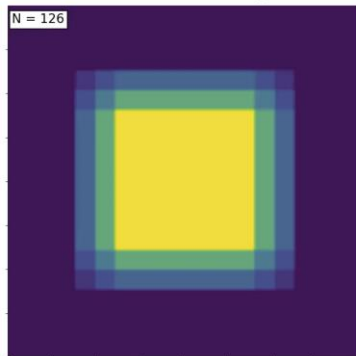


Figure 1. Coverage of a single observation sequence using the rolling dithering method. For each exposure, the offset is randomly determined within a $30' \times 30'$ box relative to the original coordinate. The camera's position angle rotates by 160° after several exposures, and the angular direction is displayed in the upper left of the panel. The effective area with the longest integration time is $\sim 12 \text{ deg}^2$.

Survey preparation: sky simulation & rolling dithering

Offset dithering



Rolling dithering

