

Covariant formulation of classical electromagnetism

The <u>covariant formulation of classical electromagnetism</u> refers to ways of writing the laws of classical electromagnetism (in particular, <u>Maxwell's equations</u> and the <u>Lorentz force</u>) in a form that is manifestly invariant under <u>Lorentz transformations</u>, in the formalism of <u>special relativity</u> using rectilinear <u>inertial coordinate systems</u>. These expressions both make it simple to prove that the laws of classical electromagnetism take the same form in any inertial coordinate system, and also provide a way to translate the fields and forces from one frame to another. However, this is not as general as <u>Maxwell's equations in curved spacetime</u> or non-rectilinear coordinate systems. [a]

Covariant objects

Preliminary four-vectors

Lorentz tensors of the following kinds may be used in this article to describe bodies or particles:

• four-displacement:

$$x^{lpha}=(ct,\mathbf{x})=(ct,x,y,z)$$
 .

■ Four-velocity:

$$u^{lpha}=\gamma(c,\mathbf{u}),$$

where $y(\mathbf{u})$ is the Lorentz factor at the 3-velocity \mathbf{u} .

■ Four-momentum:

$$p^{lpha}=(E/c,{f p})=m_0u^{lpha}$$

where ${\bf p}$ is 3-momentum, ${\bf \it E}$ is the total energy, and ${\bf \it m}_{0}$ is rest mass.

■ Four-gradient:

$$\partial^{
u} = rac{\partial}{\partial x_{
u}} = \left(rac{1}{c}rac{\partial}{\partial t}, -
abla
ight)\,,$$

■ The <u>d'Alembertian</u> operator is denoted ∂^2 ,

$$\partial^2 = rac{1}{c^2}rac{\partial^2}{\partial t^2} -
abla^2.$$

The signs in the following tensor analysis depend on the <u>convention</u> used for the <u>metric tensor</u>. The convention used here is (+ - - -), corresponding to the Minkowski metric tensor:

$$\eta^{\mu
u} = egin{pmatrix} 1 & 0 & 0 & 0 \ 0 & -1 & 0 & 0 \ 0 & 0 & -1 & 0 \ 0 & 0 & 0 & -1 \end{pmatrix}$$

Electromagnetic tensor

The electromagnetic tensor is the combination of the electric and magnetic fields into a covariant antisymmetric tensor whose entries are \mathbf{B} -field quantities. [1]

$$F_{lphaeta} = egin{pmatrix} 0 & E_x/c & E_y/c & E_z/c \ -E_x/c & 0 & -B_z & B_y \ -E_y/c & B_z & 0 & -B_x \ -E_z/c & -B_y & B_x & 0 \end{pmatrix}$$

and the result of raising its indices is

$$F^{\mu
u} \stackrel{
m def}{=} \eta^{\mulpha} \, F_{lphaeta} \, \eta^{eta
u} = egin{pmatrix} 0 & -E_x/c & -E_y/c & -E_z/c \ E_x/c & 0 & -B_z & B_y \ E_y/c & B_z & 0 & -B_x \ E_z/c & -B_y & B_x & 0 \end{pmatrix},$$

where E is the electric field, B the magnetic field, and c the speed of light.

Four-current

The four-current is the contravariant four-vector which combines electric charge density ρ and electric current density \mathbf{j} :

$$J^{lpha}=(c
ho,\mathbf{j})$$
 .

Four-potential

The electromagnetic four-potential is a covariant four-vector containing the <u>electric potential</u> (also called the scalar potential) ϕ and magnetic vector potential (or vector potential) **A**, as follows:

$$A^{lpha}=(\phi/c,{f A})$$
 .

The differential of the electromagnetic potential is

$$F_{lphaeta} = \partial_lpha A_eta - \partial_eta A_lpha$$
 .

In the language of differential forms, which provides the generalisation to curved spacetimes, these are the components of a 1-form $A = A_{\alpha} dx^{\alpha}$ and a 2-form $F = dA = \frac{1}{2} F_{\alpha\beta} dx^{\alpha} \wedge dx^{\beta}$ respectively. Here, d is the exterior derivative and \wedge the wedge product.

Electromagnetic stress-energy tensor

The electromagnetic stress—energy tensor can be interpreted as the flux density of the momentum four-vector, and is a contravariant symmetric tensor that is the contribution of the electromagnetic fields to the overall stress—energy tensor:

$$T^{lphaeta} = egin{pmatrix} arepsilon_0 E^2/2 + B^2/2\mu_0 & S_x/c & S_y/c & S_z/c \ S_x/c & -\sigma_{xx} & -\sigma_{xy} & -\sigma_{xz} \ S_y/c & -\sigma_{yx} & -\sigma_{yy} & -\sigma_{yz} \ S_z/c & -\sigma_{zx} & -\sigma_{zy} & -\sigma_{zz} \end{pmatrix},$$

where ε_0 is the <u>electric permittivity of vacuum</u>, μ_0 is the <u>magnetic permeability of vacuum</u>, the <u>Poynting vector</u> is

$$\mathbf{S} = rac{1}{\mu_0} \mathbf{E} imes \mathbf{B}$$

and the Maxwell stress tensor is given by

$$\sigma_{ij} = arepsilon_0 E_i E_j + rac{1}{\mu_0} B_i B_j - \left(rac{1}{2}arepsilon_0 E^2 + rac{1}{2\mu_0} B^2
ight) \delta_{ij} \,.$$

The electromagnetic field tensor F constructs the electromagnetic stress—energy tensor T by the equation: [2]

$$T^{lphaeta} = rac{1}{\mu_0} \left(\eta^{lpha
u} F_{
u\gamma} F^{\gammaeta} + rac{1}{4} \eta^{lphaeta} F_{\gamma
u} F^{\gamma
u}
ight)$$

where η is the Minkowski metric tensor (with signature (+ - - -)). Notice that we use the fact that

$$\varepsilon_0 \mu_0 c^2 = 1$$
,

which is predicted by Maxwell's equations.

Maxwell's equations in vacuum

In vacuum (or for the microscopic equations, not including macroscopic material descriptions), Maxwell's equations can be written as two tensor equations.

The two inhomogeneous Maxwell's equations, <u>Gauss's Law</u> and <u>Ampère's law</u> (with Maxwell's correction) combine into (with (+---) metric): [3]

$$rac{ ext{Gauss-Ampère}}{\partial_{lpha}F^{lphaeta}}=\mu_{0}J^{eta}$$

The homogeneous equations – Faraday's law of induction and Gauss's law for magnetism combine to form $\partial^{\sigma} F^{\mu\nu} + \partial^{\mu} F^{\nu\sigma} + \partial^{\nu} F^{\sigma\mu} = 0$, which may be written using Levi-Civita duality as:

$$rac{ ext{Gauss-} ext{Faraday}}{\partial_{lpha}\left(arepsilon^{lphaeta\gamma\delta}F_{\gamma\delta}
ight)=0}$$

where $F^{\alpha\beta}$ is the <u>electromagnetic tensor</u>, J^{α} is the <u>four-current</u>, $\varepsilon^{\alpha\beta\gamma\delta}$ is the <u>Levi-Civita symbol</u>, and the indices behave according to the Einstein summation convention.

Each of these tensor equations corresponds to four scalar equations, one for each value of β .

Using the <u>antisymmetric tensor</u> notation and comma notation for the partial derivative (see <u>Ricci calculus</u>), the second equation can also be written more compactly as:

$$F_{[\alpha\beta,\gamma]}=0.$$

In the absence of sources, Maxwell's equations reduce to:

$$\partial^
u\partial_
u F^{lphaeta} \stackrel{
m def}{=} \partial^2 F^{lphaeta} \stackrel{
m def}{=} rac{1}{c^2} rac{\partial^2 F^{lphaeta}}{\partial t^2} -
abla^2 F^{lphaeta} = 0 \, ,$$

which is an electromagnetic wave equation in the field strength tensor.

Maxwell's equations in the Lorenz gauge

The <u>Lorenz gauge condition</u> is a Lorentz-invariant gauge condition. (This can be contrasted with other gauge <u>conditions</u> such as the <u>Coulomb gauge</u>, which if it holds in one <u>inertial frame</u> will generally not hold in any other.) It is expressed in terms of the four-potential as follows:

$$\partial_{\alpha}A^{\alpha}=\partial^{\alpha}A_{\alpha}=0$$
.

In the Lorenz gauge, the microscopic Maxwell's equations can be written as:

$$\partial^2 A^\sigma = \mu_0 J^\sigma$$
 .

Lorentz force

Charged particle

Electromagnetic (EM) fields affect the motion of <u>electrically charged</u> matter: due to the <u>Lorentz force</u>. In this way, EM fields can be <u>detected</u> (with applications in <u>particle physics</u>, and natural occurrences such as in <u>aurorae</u>). In relativistic form, the Lorentz force uses the field strength tensor as follows. [4]

Expressed in terms of coordinate time *t*, it is:

$$rac{dp_{lpha}}{dt}=q\,F_{lphaeta}\,rac{dx^{eta}}{dt},$$

where p_a is the four-momentum, q is the charge, and x^{β} is the position.

Expressed in frame-independent form, we have the four-force

$$rac{dp_{lpha}}{d au}\,=q\,F_{lphaeta}\,u^{eta},$$

where u^{β} is the four-velocity, and τ is the particle's <u>proper time</u>, which is related to coordinate time by $dt = \gamma d\tau$.

$q (\mathbf{v} \times \mathbf{B})$ $q \in \mathbf{B}$

Lorentz force **f** on a charged particle (of charge *q*) in motion (instantaneous velocity **v**). The **E** field and **B** field vary in space and time.

Charge continuum

The density of force due to electromagnetism, whose spatial part is the Lorentz force, is given by

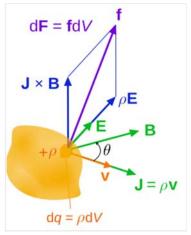
$$f_{lpha}=F_{lphaeta}J^{eta}.$$

and is related to the electromagnetic stress-energy tensor by

$$f^lpha = -T^{lphaeta}_{,eta} \equiv -rac{\partial T^{lphaeta}}{\partial x^eta}.$$

Conservation laws

Electric charge



Lorentz force per spatial volume f on a continuous charge distribution (charge density ρ) in motion.

The continuity equation:

$$J^{eta}_{\;\;,eta}\stackrel{ ext{def}}{=}\partial_{eta}J^{eta}=\partial_{eta}\partial_{lpha}F^{lphaeta}/\mu_0=0.$$

expresses charge conservation.

Electromagnetic energy-momentum

Using the Maxwell equations, one can see that the <u>electromagnetic stress-energy tensor</u> (defined above) satisfies the following differential equation, relating it to the electromagnetic tensor and the current four-vector

$$T^{lphaeta}_{,eta}+F^{lphaeta}J_eta=0$$

or

$$\eta_{lpha
u}T^{
ueta}_{\ \ ,eta}+F_{lphaeta}J^eta=0,$$

which expresses the conservation of linear momentum and energy by electromagnetic interactions.

Covariant objects in matter

Free and bound four-currents

In order to solve the equations of electromagnetism given here, it is necessary to add information about how to calculate the electric current, J^{ν} . Frequently, it is convenient to separate the current into two parts, the free current and the bound current, which are modeled by different equations;

$$J^{\nu} = J^{\nu}_{\text{free}} + J^{\nu}_{\text{bound}}$$
,

where

$$egin{aligned} J^{
u}_{ ext{ free}} &= \left(\, c
ho_{ ext{free}}, \quad \mathbf{J}_{ ext{free}} \,
ight) = \left(\, c
abla \cdot \mathbf{D}, \quad -rac{\partial \mathbf{D}}{\partial t} +
abla imes \mathbf{H} \,
ight), \ J^{
u}_{ ext{ bound}} &= \left(\, c
ho_{ ext{bound}}, \quad \mathbf{J}_{ ext{bound}} \,
ight) = \left(\, -c
abla \cdot \mathbf{P}, \quad rac{\partial \mathbf{P}}{\partial t} +
abla imes \mathbf{M} \,
ight). \end{aligned}$$

 $\underline{\text{Maxwell's macroscopic equations}}$ have been used, in addition the definitions of the $\underline{\text{electric displacement}}$ \mathbf{D} and the magnetic intensity \mathbf{H} :

$$\mathbf{D} = arepsilon_0 \mathbf{E} + \mathbf{P}, \ \mathbf{H} = rac{1}{\mu_0} \mathbf{B} - \mathbf{M} \,.$$

where ${\bf M}$ is the magnetization and ${\bf P}$ the electric polarization.

Magnetization-polarization tensor

The bound current is derived from the **P** and **M** fields which form an antisymmetric contravariant magnetization-polarization tensor $\frac{[1]}{[5]} \frac{[6][7]}{[6][7]}$

$$\mathcal{M}^{\mu
u} = egin{pmatrix} 0 & P_x c & P_y c & P_z c \ -P_x c & 0 & -M_z & M_y \ -P_y c & M_z & 0 & -M_x \ -P_z c & -M_y & M_x & 0 \end{pmatrix},$$

which determines the bound current

$$J^{
u}_{\mathrm{bound}} = \partial_{\mu} \mathcal{M}^{\mu
u}$$
 .

Electric displacement tensor

If this is combined with $F^{\mu\nu}$ we get the antisymmetric contravariant electromagnetic displacement tensor which combines the **D** and **H** fields as follows:

$$\mathcal{D}^{\mu
u} = egin{pmatrix} 0 & -D_x c & -D_y c & -D_z c \ D_x c & 0 & -H_z & H_y \ D_y c & H_z & 0 & -H_x \ D_z c & -H_y & H_x & 0 \end{pmatrix}.$$

The three field tensors are related by:

$$\mathcal{D}^{\mu
u} = rac{1}{\mu_0} F^{\mu
u} - \mathcal{M}^{\mu
u}$$

which is equivalent to the definitions of the **D** and **H** fields given above.

Maxwell's equations in matter

The result is that Ampère's law,

$$abla imes \mathbf{H} - rac{\partial \mathbf{D}}{\partial t} = \mathbf{J}_{ ext{free}},$$

and Gauss's law,

$$\nabla \cdot \mathbf{D} = \rho_{\text{free}}$$

combine into one equation:

$$rac{ ext{Gauss-Ampère}}{J^{
u}_{ ext{free}} = \partial_{\mu} \mathcal{D}^{\mu
u}}$$

The bound current and free current as defined above are automatically and separately conserved

$$egin{aligned} \partial_{
u} J^{
u}_{
m \ bound} &= 0 \ \partial_{
u} J^{
u}_{
m \ free} &= 0 \,. \end{aligned}$$

Constitutive equations

Vacuum

In vacuum, the constitutive relations between the field tensor and displacement tensor are:

$$\mu_0 \mathcal{D}^{\mu
u} = \eta^{\mulpha} F_{lphaeta} \eta^{eta
u} \, .$$

Antisymmetry reduces these 16 equations to just six independent equations. Because it is usual to define $F^{\mu\nu}$ by

$$F^{\mu
u}=\eta^{\mulpha}F_{lphaeta}\eta^{eta
u},$$

the constitutive equations may, in vacuum, be combined with the Gauss-Ampère law to get:

$$\partial_{eta}F^{lphaeta}=\mu_{0}J^{lpha}.$$

The electromagnetic stress—energy tensor in terms of the displacement is:

$$T_{lpha}{}^{\pi} = F_{lphaeta}\mathcal{D}^{\pieta} - rac{1}{4}\delta^{\pi}_{lpha}F_{\mu
u}\mathcal{D}^{\mu
u},$$

where $\delta_{\alpha}^{\ \pi}$ is the Kronecker delta. When the upper index is lowered with η , it becomes symmetric and is part of the source of the gravitational field.

Linear, nondispersive matter

Thus we have reduced the problem of modeling the current, J^{ν} to two (hopefully) easier problems — modeling the free current, J^{ν}_{free} and modeling the magnetization and polarization, $\mathcal{M}^{\mu\nu}$. For example, in the simplest materials at low frequencies, one has

$$egin{aligned} \mathbf{J}_{ ext{free}} &= \sigma \mathbf{E} \ \mathbf{P} &= arepsilon_0 \chi_e \mathbf{E} \ \mathbf{M} &= \chi_m \mathbf{H} \end{aligned}$$

where one is in the instantaneously comoving inertial frame of the material, σ is its electrical conductivity, χ_e is its electric susceptibility, and χ_m is its magnetic susceptibility.

The constitutive relations between the \mathcal{D} and F tensors, proposed by Minkowski for a linear materials (that is, \mathbf{E} is proportional to \mathbf{D} and \mathbf{B} proportional to \mathbf{H}), are:

$$\mathcal{D}^{\mu
u}u_
u = c^2 arepsilon F^{\mu
u}u_
u \ \star \mathcal{D}^{\mu
u}u_
u = rac{1}{\mu} \star F^{\mu
u}u_
u$$

where u is the four-velocity of material, ε and μ are respectively the proper <u>permittivity</u> and <u>permeability</u> of the material (i.e. in rest frame of material), \star and denotes the Hodge star operator.

Lagrangian for classical electrodynamics

Vacuum

The <u>Lagrangian</u> density for classical electrodynamics is composed by two components: a field component and a source component:

$${\cal L} \,=\, {\cal L}_{
m field} + {\cal L}_{
m int} = -rac{1}{4\mu_0} F^{lphaeta} F_{lphaeta} - A_lpha J^lpha \,.$$

In the interaction term, the four-current should be understood as an abbreviation of many terms expressing the electric currents of other charged fields in terms of their variables; the four-current is not itself a fundamental field.

The <u>Lagrange equations</u> for the electromagnetic lagrangian density $\mathcal{L}(A_{\alpha}, \partial_{\beta} A_{\alpha})$ can be stated as follows:

$$\partial_eta \left[rac{\partial \mathcal{L}}{\partial (\partial_eta A_lpha)}
ight] - rac{\partial \mathcal{L}}{\partial A_lpha} = 0 \, .$$

Noting

$$F^{\lambda\sigma} = F_{\mu
u}\eta^{\mu\lambda}\eta^{
u\sigma}, \ F_{\mu
u} = \partial_{\mu}A_{
u} - \partial_{
u}A_{\mu} \ rac{\partial\left(\partial_{\mu}A_{
u}
ight)}{\partial\left(\partial_{
ho}A_{\sigma}
ight)} = \delta^{
ho}_{\mu}\delta^{\sigma}_{
u}$$

the expression inside the square bracket is

$$egin{aligned} rac{\partial \mathcal{L}}{\partial (\partial_{eta} A_{lpha})} &= -rac{1}{4\mu_0} \; rac{\partial \left(F_{\mu
u} \eta^{\mu\lambda} \eta^{
u\sigma} F_{\lambda\sigma}
ight)}{\partial \left(\partial_{eta} A_{lpha}
ight)} \ &= -rac{1}{4\mu_0} \; \eta^{\mu\lambda} \eta^{
u\sigma} \left(F_{\lambda\sigma} \left(\delta^{eta}_{\mu} \delta^{lpha}_{
u} - \delta^{eta}_{
u} \delta^{lpha}_{\mu}
ight) + F_{\mu
u} \left(\delta^{eta}_{\lambda} \delta^{lpha}_{\sigma} - \delta^{eta}_{\sigma} \delta^{lpha}_{\lambda}
ight)
ight) \ &= -rac{F^{etalpha}}{\mu_0} \; . \end{aligned}$$

The second term is

$$rac{\partial \mathcal{L}}{\partial A_{lpha}} = -J^{lpha} \, .$$

Therefore, the electromagnetic field's equations of motion are

$$rac{\partial F^{etalpha}}{\partial x^eta} = \mu_0 J^lpha \ .$$

which is the Gauss-Ampère equation above.

Matter

Separating the free currents from the bound currents, another way to write the Lagrangian density is as follows:

$${\cal L} \, = \, -rac{1}{4\mu_0} F^{lphaeta} F_{lphaeta} - A_lpha J_{
m free}^lpha + rac{1}{2} F_{lphaeta} {\cal M}^{lphaeta} \, .$$

Using Lagrange equation, the equations of motion for $\mathcal{D}^{\mu\nu}$ can be derived.

The equivalent expression in vector notation is:

$$\mathcal{L} \,=\, rac{1}{2} \left(arepsilon_0 E^2 - rac{1}{\mu_0} B^2
ight) - \phi \,
ho_{ ext{free}} + \mathbf{A} \cdot \mathbf{J}_{ ext{free}} + \mathbf{E} \cdot \mathbf{P} + \mathbf{B} \cdot \mathbf{M} \,.$$

See also

- Covariant classical field theory
- Electromagnetic tensor
- Electromagnetic wave equation
- Liénard–Wiechert potential for a charge in arbitrary motion
- Moving magnet and conductor problem
- Inhomogeneous electromagnetic wave equation
- Proca action
- Quantum electrodynamics
- Relativistic electromagnetism

- Stueckelberg action
- Wheeler–Feynman absorber theory

Notes

a. This article uses the <u>classical treatment of tensors</u> and <u>Einstein summation convention</u> throughout and the <u>Minkowski metric</u> has the form diag(+1, -1, -1, -1). Where the equations are specified as holding in vacuum, one could instead regard them as the formulation of Maxwell's equations in terms of *total* charge and current.

References

- 1. Vanderlinde, Jack (2004), <u>classical electromagnetic theory</u> (https://books.google.com/books?id=HWrME T9_VpUC&q=electromagnetic+field+tensor+vanderlinde&pg=PA316), Springer, pp. 313–328, ISBN 9781402026997
- 2. Classical Electrodynamics, Jackson, 3rd edition, page 609
- 3. Classical Electrodynamics by Jackson, 3rd Edition, Chapter 11 Special Theory of Relativity
- 4. The assumption is made that no forces other than those originating in **E** and **B** are present, that is, no gravitational, weak or strong forces.
- 5. However, the assumption that $M^{\mu\nu}$, $D^{\mu\nu}$, and even $F^{\mu\nu}$, are relativistic tensors in a polarizable medium, is without foundation. The quantity

$$A^{\alpha}=(\phi/c,\mathbf{A})$$

is not a four vector in a polarizable medium, so

$$F_{lphaeta}=\partial_lpha A_eta-\partial_eta A_lpha$$

does not produce a tensor.

- 6. Franklin, Jerrold, <u>Can electromagnetic fields form tensors in a polarizable medium?</u> (https://www.acade mia.edu/104227822)
- 7. Gonano, Carlo, *Definition for Polarization P and Magnetization M Fully Consistent with Maxwell's Equations* (https://www.jpier.org/PIERB/pier.php?paper=15100606)

Further reading

- The Feynman Lectures on Physics Vol. II Ch. 25: Electrodynamics in Relativistic Notation (https://feynmanlectures.caltech.edu/II_25.html)
- Einstein, A. (1961). *Relativity: The Special and General Theory* (https://archive.org/details/relativityspecia 00eins_0). New York: Crown. ISBN 0-517-02961-8.
- Misner, Charles; Thorne, Kip S.; Wheeler, John Archibald (1973). *Gravitation*. San Francisco: W. H. Freeman. ISBN 0-7167-0344-0.
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