

# QIFT: Unruh-like and Bubble wall friction

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- **QFT vs Inhomogeneous field theory (IFT)**
- **Construction of Supersymmetric IFTs**
- **Quantization of IFT**
- **Higgs Condensate Bubble Expansion**
- **Conclusion**

# QFT vs Inhomogeneous field theory (IFT)

- (canonical, renormalizable, local ) **QFT**

$$S = \int d^d x \mathcal{L}(\phi_a, \partial_\mu \phi_a; m, g_n) \quad \text{Constant} \rightarrow \text{Poincare symmetric}$$

- Classical and **quantum approaches**
- Preferred vacuum (Minkowski spacetime: global vacuum)
- Canonical quantization:

$$[\phi_a(t, \vec{x}), \pi_a(t, \vec{y})] = i\delta^{d-1}(\vec{x} - \vec{y}) \quad \pi_a(x) \equiv \frac{\partial \mathcal{L}}{\partial \dot{\phi}_a}$$

- **Inhomogeneous field theory (IFT)**

$$S = \int d^d x \mathcal{L}(\phi_a, \partial_\mu \phi_a; m(x), g_n(x))$$

# Inhomogeneous field theories (IFT)

- **High-Energy UV theory:**

- introducing the Poincare symmetry is natural because most fields are dynamical
- exemplified: string theory, super Yang-Mills theory, ABJM theory, etc

- **Intermediate energy scale:**

- some fields become non-dynamical
- examples: field theories on curved spacetime (FTCS)
- various non-dynamical field strengths in string or M-theory

- **Low-energy scale:**

- **Inhomogeneity becomes prevalent!!**
- examples: lattice structures, and experimental setups, specific phenomena like doping, superconductivity, and defects models in condensed matter physics

# Questions related with Inhomogeneous field theories

- **What are classical effects for IFTs?**

- Construction of Supersymmetric IFTs
- BPS solutions in various IFTs

- **How to quantize IFTs?**

- Renormalized Hadamard two-point functions and quantum EM tensor including leading quantum effects

- **To which physical systems can the quantum effects of IFT be applied?**

- Unruh-like effects
- Higgs condensate bubble expansion (assuming 1<sup>st</sup> order phase transition)

- **What are classical effects for IFTs?**
  - Construction of Supersymmetric IFTs
  - BPS solutions in various IFTs

# Supersymmetric IFT

- **N = 3 Inhomogeneous mABJM in 3d**  
**and N=1 inhomogeneous mSYM in 4d**

$$\mathcal{L}_{\text{ImABJM}} = \mathcal{L}_{\text{mABJM}} + m'(x) M_A^B Y^A Y_B^\dagger$$

**ImABJM** [Kim-OK, Kim-Kim-Kim-Kwon, 2018],

**ImSYM** [Arav-Gauntlett-Roberts-Rosen, Kim-OK-Tolla]

- Supersymmetric IFT models: Inhomogeneous CSH, AH in 3d  
and real scalar model in (1+1)-dim. [Kim-OK-Song-Kim, 2024, Kim-Jeon-OK-Song-Kim, 2024  
Kwon-Kim-Kim, 2022]

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} - \overline{D_\mu \phi} D^\mu \phi - \frac{g^2}{2} [|\phi|^2 - v_0^2 - \sigma(\mathbf{x})]^2 + sg\sigma(\mathbf{x})B$$

$$\mathcal{L} = -\frac{1}{2} \partial_\mu \phi \partial^\mu \phi + i \bar{\psi} \gamma^\mu \partial_\mu \psi + i \frac{\partial^2 W}{\partial \phi^2} \bar{\psi} \psi - \frac{1}{2} \left( \frac{\partial W}{\partial \phi} \right)^2 \mp \frac{\partial W}{\partial x}$$

# Inhomogeneous vacuum configuration

- Inhomogeneous Chern-Simons Higgs model

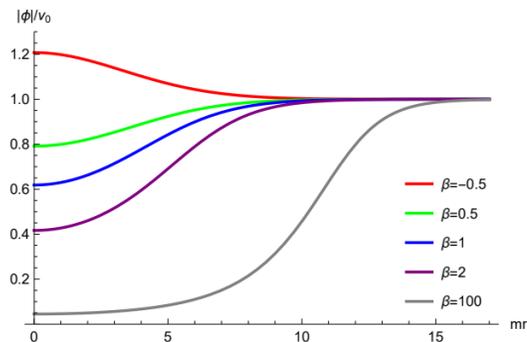
$$v^2(\mathbf{x}) = v_0^2 + \sigma(\mathbf{x})$$

$$\mathcal{L} = \frac{\kappa}{2} \epsilon^{\mu\nu\rho} A_\mu \partial_\nu A_\rho - \overline{D_\mu \phi} D^\mu \phi - \frac{1}{\kappa^2} |\phi|^2 (|\phi|^2 - v^2(\mathbf{x}))^2 + s\sigma(\mathbf{x})B$$

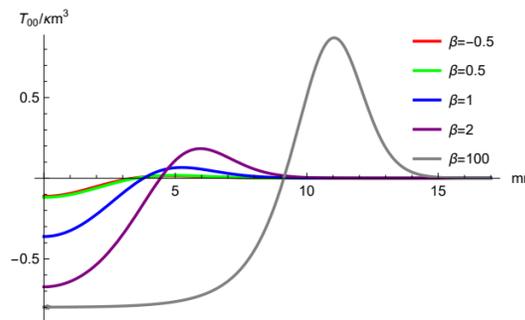
- Rotationally symmetric configuration:  $\sigma(\mathbf{x}) = -\beta v_0^2 e^{-\alpha^2 m^2 r^2}$

$$\frac{d^2 \ln |\phi|^2}{dr^2} + \frac{1}{r} \frac{d \ln |\phi|^2}{dr} = \frac{4}{\kappa^2} |\phi|^2 [|\phi|^2 - v_0^2 (1 - \beta e^{-\alpha^2 m^2 r^2})]$$

- E=0 inhomogeneous vacuum configuration: ImCSH, ImAH, etc



(a)



(b)

- **How to quantize IFTs? (one possible way)**

- Renormalized Hadamard two-point functions and quantum EM tensor including leading quantum effects

# Quantization of IFT

- In IFT, there is no preferred vacuum because Poincaré symmetry is absent.
- There is no guiding principle of quantization in IFT.
- On the other hand, in field theory on curved space (FTCS), a preferred vacuum does not exist, just as in IFT. Therefore, the canonical quantization method cannot be used, but the algebraic quantization method is employed instead. → Hadamard two-point function
- Classical conversion relation between FTCS and IFT in (1+1)-dimensions.
- We propose to promote this classical conversion relation to the quantum level.

# Quantization of IFT

- Conversion relation between FTCS and IFT in (1+1)-dimensions

## Scalar field in 2D spacetime background (FTCS)

EPJP138(2023)202 JH, O-K. Kwon, S.A Park, S.-H. Yi

$$ds^2 = \frac{1}{(a + e^{-bx})^2} (-dt^2 + dx^2) \quad ab = \frac{m_0}{2\xi}$$

$$S_{\text{FTCS}} = \int_{\mathcal{M}} d^2x \sqrt{-g} \left( -\frac{1}{2} g^{\mu\nu} \nabla_{\mu} \phi \nabla_{\nu} \phi - \frac{1}{2} m_0^2 \phi^2 - \frac{1}{2} \xi \mathcal{R} \phi^2 \right)$$

$$(-\square + m_0^2 + \xi \mathcal{R})\phi = 0, \quad \square = \frac{1}{\sqrt{-g}} \partial_{\mu} (\sqrt{-g} g^{\mu\nu} \partial_{\nu})$$

# Quantization of IFT

- Conversion relation between FTCS and IFT in (1+1)-dimensions

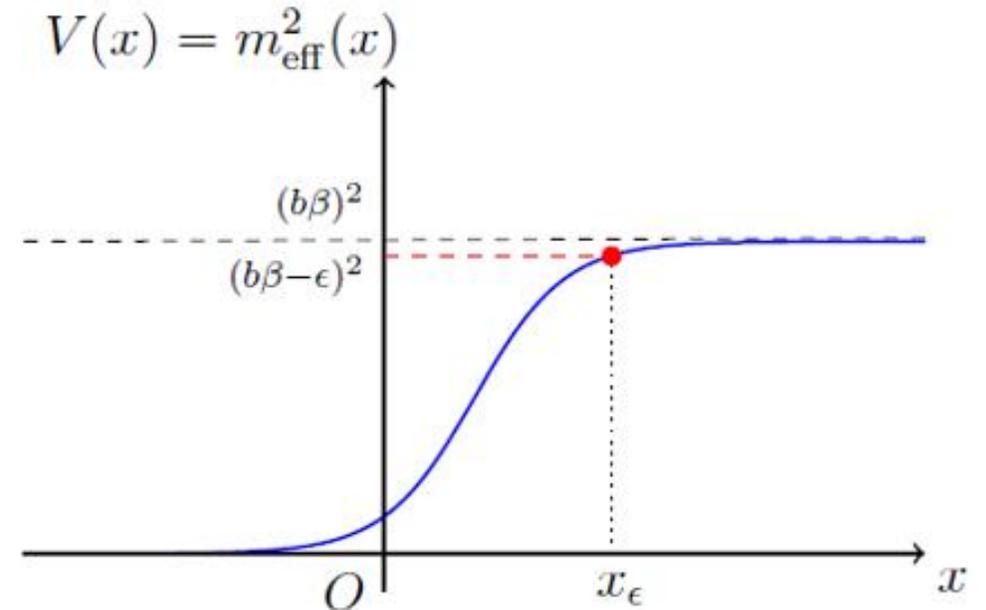
## Scalar field in 2D Minkowski spacetime with position dependent mass (IFT)

$$S_{\text{IFT}} = \int d^2x \left[ -\frac{1}{2} \eta^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{1}{2} m_{\text{eff}}^2(x) \phi^2 \right]$$

$$m_{\text{eff}}^2(x) = m^2(x) + m'(x), \quad m(x) = m_0 e^{\omega(x)}$$

$$e^{\omega(x)} = \frac{1}{a + e^{-bx}}, \quad m_{\text{eff}}^2(x) = \frac{(m_0^2 e^{bx} + 2\xi ab^2) e^{bx}}{(ae^{bx} + 1)^2}$$

$$(-\partial^2 + m_{\text{eff}}^2(x)) \phi = 0, \quad \partial^2 = \eta^{\mu\nu} \partial_\mu \partial_\nu$$



# Quantization of IFT

## Energy-Momentum Tensor in IFT

**Non-existence of Poincare symmetry in IFT :**

Defining the energy-momentum tensor in IFT as a conserved current is problematic.

Our IFT model possesses time translation symmetry  $\rightarrow$  the energy of the system is conserved.  $\rightarrow$  the (tt)-component of the energy-momentum tensor can be constructed canonically.

**No explicit criteria exist for determining the remaining components of the energy-momentum tensor.**

# Quantization of IFT

- **In the Theory of Gravity (without Poincaré symmetry):**
  - The energy-momentum tensor can be defined as the source of the gravitational field.
  - The energy-momentum tensor in IFT is directly read from the energy-momentum tensor in FTCS, using the **conversion relation** between FTCS and IFT.
  - This conversion relation provides a framework for **renormalization** of the energy-momentum tensor in QIFT.

## Positive frequency Wightman function at a position of $x_\epsilon$

Positive frequency Wightman function (R, $\epsilon$  - vacuum)

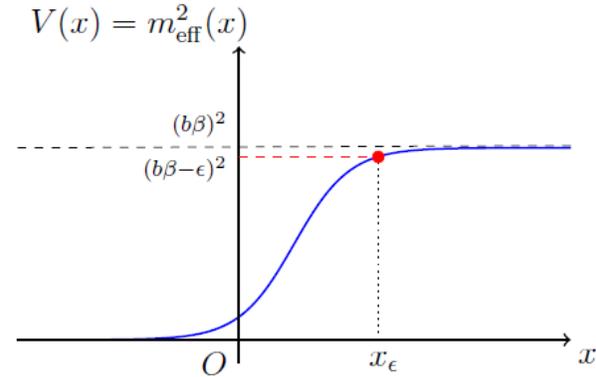
$$G_\epsilon^+(\mathbf{x}, \mathbf{x}') \simeq {}_\epsilon \langle 0 | \phi(\mathbf{x}) \phi(\mathbf{x}') | 0 \rangle_\epsilon \simeq \int_0^\infty \frac{dk}{4\pi\omega_k} \sum_{i=\pm} v_k^{(i)}(\mathbf{x}) \left( v_k^{(i)}(\mathbf{x}') \right)^*$$

$$\left( -\partial^2 + m_{\text{eff}}^2(x) \right) \phi = 0, \quad \partial^2 = \eta^{\mu\nu} \partial_\mu \partial_\nu \quad m_{\text{eff}}^2(x) = \frac{(m_0^2 e^{bx} + 2\xi ab^2) e^{bx}}{(a e^{bx} + 1)^2}$$

$$v_k^{(-)}(\mathbf{x}) = (1 + e^{-bx})^{2\xi} F \left( A, A - C + 1; A - B + 1 \mid -e^{-bx} \right) e^{-i(\omega t - kx)},$$

$$v_k^{(+)}(\mathbf{x}) = (1 + e^{-bx})^{2\xi} F \left( B, B - C + 1; B - A + 1 \mid -e^{-bx} \right) e^{-i(\omega t + kx)},$$

$$A = \frac{i}{b} (\omega - k) + \beta, \quad B = \frac{i}{b} (\omega + k) + \beta, \quad C = 1 + \frac{2i\omega}{b}, \quad k^2 \equiv \omega^2 - (2b\xi)^2 \quad \beta = 2\xi = \frac{m_0}{ab}$$



Singular structure of the Wightman function in 2D

$$G^+(\mathbf{x}, \mathbf{x}') = \frac{1}{4\pi} \left( V(\mathbf{x}, \mathbf{x}') \ln [\mu^2 \sigma(\mathbf{x}, \mathbf{x}')] + W(\mathbf{x}, \mathbf{x}'; \mu) \right)$$

Positive frequency Wightman function  $G^+(x, x') = \langle 0 | \phi(x) \phi(x') | 0 \rangle$

Hadamard two-point function:  $G_H(x, x') = \langle 0 | \{ \phi(x), \phi(x') \} | 0 \rangle$

## Hadamard functions

- ◆  $2\sigma(x, x')$  : the square of the geodesic distance between the points  $x$  and  $x'$ , Synge function
- ◆  $V(x, x'), W(x, x')$  : symmetric biscalar functions that remain regular as  $x'$  approaches  $x$ , which are determined by the geometry and the field equation.
- ◆  $W(x, x')$  encodes **quantum effect of  $|0\rangle_{\mathbb{R}}$** .

# Key Properties of Energy-Momentum Tensor $\langle T_{\mu\nu} \rangle_H$ in Hadamard State

## 1. Physical Validity

- The Hadamard state ensures that  $\langle T_{\mu\nu} \rangle_H$  provides physically meaningful and finite values.
- This avoids divergences and supports consistent physical interpretation in curved spacetimes.

## 2. Covariant Conservation

- The energy-momentum tensor satisfies the covariant conservation law:

$$\nabla^\mu \langle T_{\mu\nu} \rangle_H = 0$$

- This is essential for maintaining energy conservation and consistency with general relativity.

## Renormalized Positive frequency Wightman function

$$G_{\text{ren}}^+(x, x'; \mu) = G_{\epsilon}^+(x, x') - \frac{1}{4\pi} V(x, x') \ln [\mu^2 \sigma(x, x')] = \frac{1}{4\pi} W(x, x'; \mu)$$

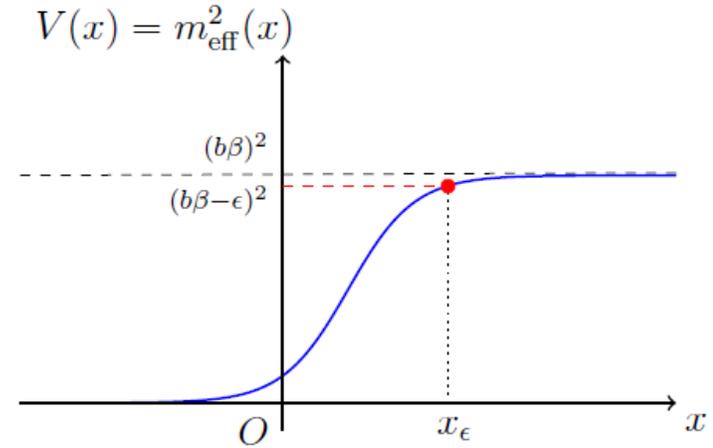
$$G_{\epsilon}^+(x, x') \simeq_{\text{R}} \langle 0 | \phi(x) \phi(x') | 0 \rangle_{\text{R}}^{\epsilon} \simeq \int_0^{\infty} \frac{dk}{4\pi\omega_k} \sum_{i=\pm} v_k^{(i)}(x) \left( v_k^{(i)}(x') \right)^*$$

$$v_k^{(-)}(\mathbf{x}) = (1 + e^{-bx})^{2\xi} F \left( A, A - C + 1; A - B + 1 \mid -e^{-bx} \right) e^{-i(\omega t - kx)},$$

$$v_k^{(+)}(\mathbf{x}) = (1 + e^{-bx})^{2\xi} F \left( B, B - C + 1; B - A + 1 \mid -e^{-bx} \right) e^{-i(\omega t + kx)},$$

## Calculation

$$G_{\text{ren}}^+(\mathbf{x}, \mathbf{x}'; \mu) = \frac{1}{4\pi} W(\mathbf{x}, \mathbf{x}'; \mu) = G_{\epsilon}^+(\mathbf{x}, \mathbf{x}') - \frac{1}{4\pi} V(\mathbf{x}, \mathbf{x}') \ln [\mu^2 \sigma(\mathbf{x}, \mathbf{x}')] ]$$



## Renormalized Wightman function

$$\begin{aligned} G_{\text{ren}}^+(\mathbf{x}, \mathbf{x}'; \mu) &= \frac{1}{4\pi} \left( 1 + \frac{b^2}{16} (-(t - t')^2 + (x - x')^2) \right) \ln \left( \frac{2\mu^2}{a^2 b^2} \right) \\ &+ \frac{1}{2\pi} (-\gamma + \ln 2) + \frac{b^2}{32\pi} (-(t - t')^2 + (x - x')^2) (1 - \gamma + \ln 2) \\ &- \frac{1}{2\pi} \left( 1 + \frac{b^2}{24} (x - x')^2 + \frac{b^2}{16} (-(t - t')^2 + (x - x')^2) \right) e^{-b \frac{x+x'}{2}} + \dots \end{aligned}$$

where the ellipsis  $\dots$  denotes higher-order terms in  $(t - t')$ ,  $(x - x')$ , and  $e^{-b \frac{x+x'}{2}}$ .

## Point-splitting method

$$S_{\text{FTCS}} = \int d^2x \sqrt{-g} \left[ -\frac{1}{2} g^{\mu\nu} \nabla_\mu \phi \nabla_\nu \phi - \frac{1}{2} m_0^2 \phi^2 - \xi \mathcal{R} \phi^2 \right] \quad P_x \equiv -\square_x + m_0^2 + \xi \mathcal{R}$$

$$T_{\mu\nu} = \nabla_\mu \phi \nabla_\nu \phi - \frac{1}{2} g_{\mu\nu} \left[ (\nabla \phi)^2 + m_0^2 \phi^2 \right] + \xi \left( -\nabla_\mu \nabla_\nu + g_{\mu\nu} \nabla^2 \right) \phi^2 - 2\xi \delta_\mu^{\mu'} \partial_{\mu'} \partial_{\nu'} + \frac{1}{4} g_{\mu\nu} P_x$$

$$\langle T_{\mu\nu}(\mathbf{x}; \mu) \rangle_{\text{R}}^\epsilon = \lim_{\mathbf{x}' \rightarrow \mathbf{x}} \mathcal{T}_{\mu\nu'} G_{\text{ren}}^+(\mathbf{x}, \mathbf{x}'; \mu)$$

## Calculation

$$\begin{aligned}\langle T_{tt} \rangle_{\text{R}}^{\epsilon} &= \frac{b^2}{16\pi} \left[ -\gamma + \ln 2 + \frac{1}{2} \ln \left( \frac{2\mu^2}{a^2 b^2} \right) \right] \\ &\quad + \frac{b^2}{8\pi} \left[ -\frac{1}{6} + \gamma - \ln 2 - \frac{1}{2} \ln \left( \frac{2\mu^2}{a^2 b^2} \right) \right] e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{xx} \rangle_{\text{R}}^{\epsilon} &= \frac{b^2}{16\pi} \left[ \gamma - \ln 2 - \frac{1}{2} \ln \left( \frac{2\mu^2}{a^2 b^2} \right) \right] \\ &\quad + \frac{b^2}{8\pi} \left[ \frac{1}{2} - \gamma + \ln 2 + \frac{1}{2} \ln \left( \frac{2\mu^2}{a^2 b^2} \right) \right] e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{tx} \rangle_{\text{R}}^{\epsilon} &= \langle T_{xt} \rangle_{\text{R}}^{\epsilon} = \mathcal{O}(e^{-2bx}).\end{aligned}$$

**Conservation of VEV of EM tensor**  $\nabla_{\mu} \langle T^{\mu\nu} \rangle_{\text{R}}^{\epsilon} = 0$ , up to the order of  $\mathcal{O}(e^{-2bx})$

## Fixing $\mu$ in the VEV of Minkowski vacuum

Requiring  ${}_{\text{M}}\langle 0|T_{\mu\nu}|0\rangle_{\text{M}} = 0$ . in this Minkowski metric,  $ds^2 = \frac{1}{a^2}(-dt^2 + dx^2)$ .

$$\langle T_{tt} \rangle_{\text{M}} = \frac{m_0^2}{4\pi a^2} \left[ -\gamma + \ln 2 + \frac{1}{2} \ln \left( \frac{\mu^2}{2m_0^2} \right) \right],$$

$$\langle T_{xx} \rangle_{\text{M}} = \frac{m_0^2}{4\pi a^2} \left[ \gamma - \ln 2 - \frac{1}{2} \ln \left( \frac{\mu^2}{2m_0^2} \right) \right],$$

$$\langle T_{tx} \rangle_{\text{M}} = \langle T_{xt} \rangle_{\text{M}} = 0.$$

$$\mu = \frac{m_0}{\sqrt{2}} e^\gamma$$

## Result of the VEV of Energy-Momentum Tensor

$$\langle T_{tt} \rangle_{\mathbb{R}}^{\epsilon} = -\frac{b^2}{48\pi} e^{-bx} + \mathcal{O}(e^{-2bx}),$$

$$\langle T_{xx} \rangle_{\mathbb{R}}^{\epsilon} = \frac{b^2}{16\pi} e^{-bx} + \mathcal{O}(e^{-2bx}),$$

$$\langle T_{tx} \rangle_{\mathbb{R}}^{\epsilon} = \langle T_{xt} \rangle_{\mathbb{R}}^{\epsilon} = \mathcal{O}(e^{-2bx}).$$

The non-vanishing quantities in the above result represent the quantum effects of the state  $|0\rangle_{\mathbb{R}}^{\epsilon}$

- **Which physical systems can the quantum effects of IFT be applied to?**

- Interpretation for the quantum effect of EM tensor
- Higgs condensate bubble expansion (assuming 1<sup>st</sup> order phase transition)

Negative quantity :  $\langle T_{tt} \rangle_{\text{R}}^{\epsilon} < 0$

$$\begin{aligned}\langle T_{tt} \rangle_{\text{R}}^{\epsilon} &= -\frac{b^2}{48\pi} e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{xx} \rangle_{\text{R}}^{\epsilon} &= \frac{b^2}{16\pi} e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{tx} \rangle_{\text{R}}^{\epsilon} &= \langle T_{xt} \rangle_{\text{R}}^{\epsilon} = \mathcal{O}(e^{-2bx}).\end{aligned}$$

### (Review) Unruh effect :

R.-M.-Wald,  $\Gamma$ Quantum Field Theory in Curved Space-Time and Black-Hole Thermodynamics,  $\Gamma$ University of Chicago, (1994);

S.-Hollands and R.-M.-Wald,  $\Gamma$ Quantum fields in curved spacetime,  $\Gamma$ Phys.-Rept.-574,-1-35-(2015)-[arXiv:1401.2026-[gr-qc]].

If  $\langle T_{tt} \rangle_{\text{M}}$  the VEV of  $T_{tt}$  for the Minkowski vacuum  $|0\rangle_{\text{M}}$ , is set to zero by the renormalization condition,

$\langle T_{tt} \rangle_{\text{Rindler}}$  the VEV of  $T_{tt}$  for the Rindler vacuum  $|0\rangle_{\text{Rindler}}$ , becomes negative.

→ The energy density of the Rindler vacuum state is lower than that of the Minkowski vacuum.

→ Minkowski observers : NO particles

**Negative quantity** :  $\langle T_{tt} \rangle_{\mathbf{R}}^{\epsilon} < 0$

$\langle T_{tt} \rangle_{\mathbf{R}(\mathbf{M})}$ , the VEV of  $T_{tt}$  for the Minkowski vacuum  $|0\rangle_{\mathbf{R}(\mathbf{M})}$ , has been **set to zero** by the renormalization condition, then,  $\langle T_{tt} \rangle_{\mathbf{R}}^{\epsilon}$ , the VEV of  $T_{tt}$  for the vacuum  $|0\rangle_{\mathbf{R}}^{\epsilon}$ , becomes **negative**.

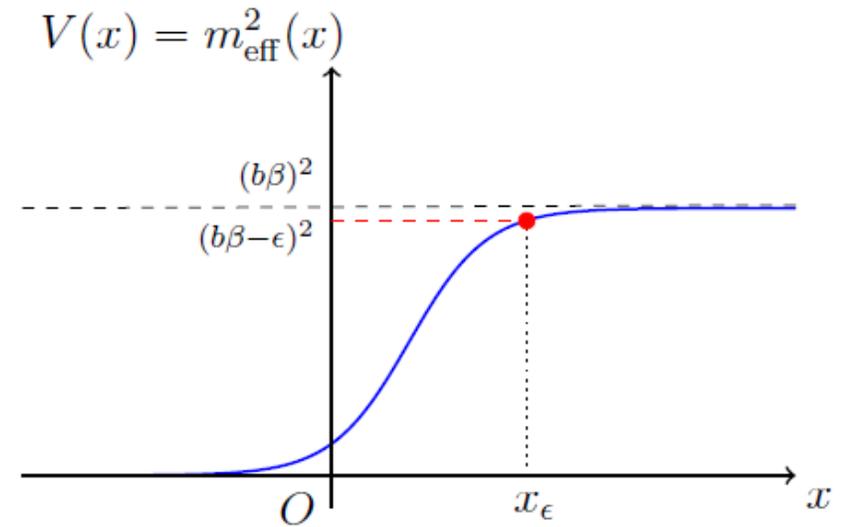
In the analogy with the Unruh effect,

**Unruh-like Effect** An observer located slightly out of the right asymptotic region would detect **a thermal-like particle distribution** for the field in the right Minkowski vacuum.

$$\langle T_{tt} \rangle_{\mathbf{R}}^{\epsilon} = -\frac{b^2}{48\pi} e^{-bx} + \mathcal{O}(e^{-2bx}),$$

$$\langle T_{xx} \rangle_{\mathbf{R}}^{\epsilon} = \frac{b^2}{16\pi} e^{-bx} + \mathcal{O}(e^{-2bx}),$$

$$\langle T_{tx} \rangle_{\mathbf{R}}^{\epsilon} = \langle T_{xt} \rangle_{\mathbf{R}}^{\epsilon} = \mathcal{O}(e^{-2bx}).$$



**Positive quantity :  $\langle T_{xx} \rangle_R^\epsilon > 0$**

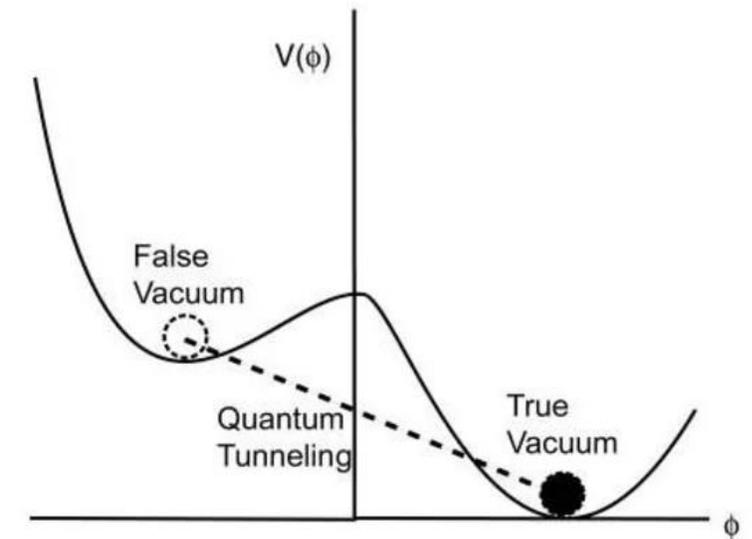
**Higgs condensate bubble expansion**

**Bubble nucleation, expansion, and plasma friction**

**Nucleation of true vacuum bubble within the false vacuum phase** during the first-order electroweak phase transition in the early universe S.R. Coleman (1977) ; C.G. Callan, Jr. and S.R. Coleman (1977), A.D. Linde (1983)

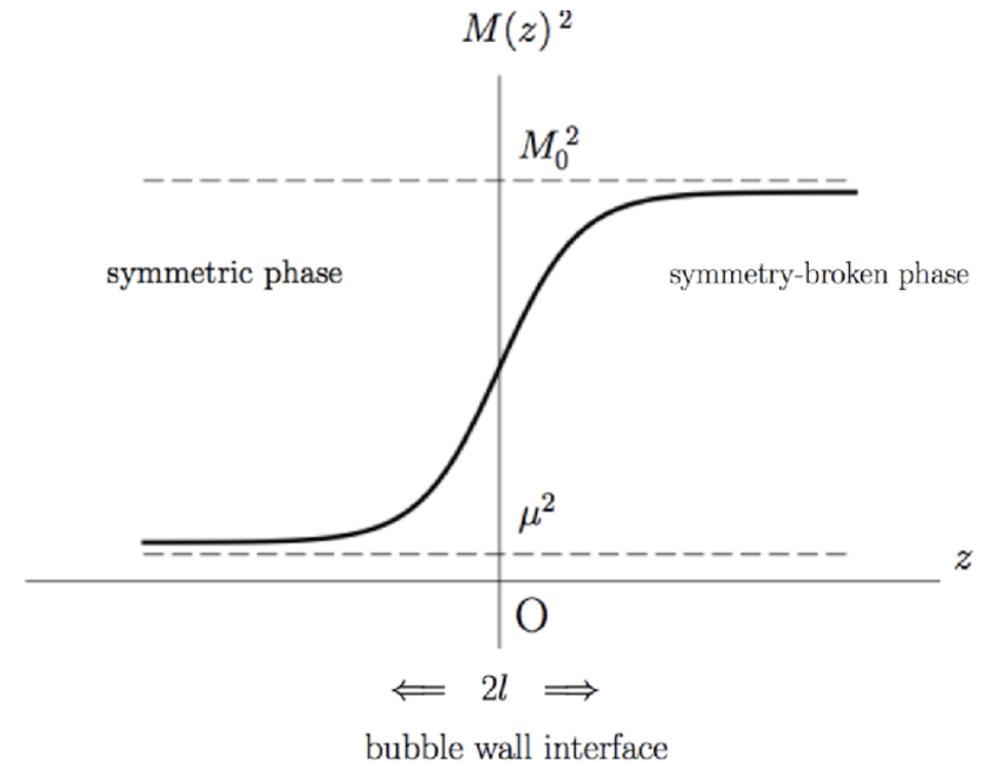
**Bubble expansion outward** driven by the energy difference between the true and false vacua.

$$\begin{aligned}\langle T_{tt} \rangle_R^\epsilon &= -\frac{b^2}{48\pi} e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{xx} \rangle_R^\epsilon &= \frac{b^2}{16\pi} e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{tx} \rangle_R^\epsilon &= \langle T_{xt} \rangle_R^\epsilon = \mathcal{O}(e^{-2bx}).\end{aligned}$$



## Bubble nucleation, expansion, and plasma friction

Plasma **friction** against bubble wall expansion through interaction with surrounding plasma of particles



## Terminal vel. of bubble walls and gravitational waves

Fictional effect → Terminal vel.

Ultra-relativistic or much slower than the speed of light

Study of **speed of the bubble wall expansion** is crucial for detecting **the gravitational wave as the evidence for the electroweak phase transition.**

Expansion speed ↑ → Collision Energy ↑

→ Gravitational Wave Amplitude, Peak frequency of gravitational waves ↑

## Quantum effect to Bubble wall expansion

A. D. Linde, Nucl. Phys. B 216, 421 (1983) [erratum: Nucl. Phys. B 223, 544 (1983)]

The total pressure exerted on the bubble wall

$$P_{\text{tot}} = -\Delta V + \Delta P,$$

In the context of our effective IFT model for bubble wall expansion (the bubble wall rest frame), the driving force  $\Delta V$  becomes irrelevant, while  $\Delta P$  needs to be taken into account.

$\Delta P$  includes both classical and quantum effects

$$\Delta P = \Delta P_{\text{classical}} + \Delta P_{\text{quantum}}$$

$$\Delta P_{\text{classical}} = 0 \quad \phi_{\text{classical}} = 0$$

## Quantum effect to Bubble wall expansion

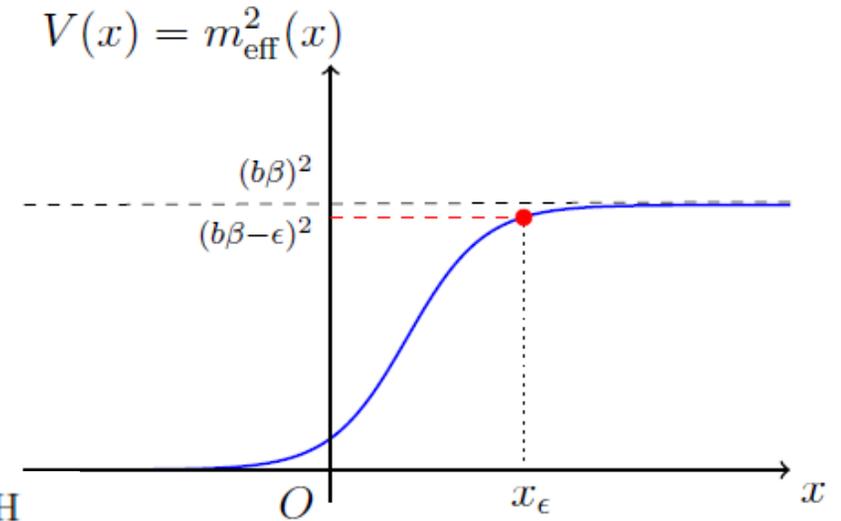
Vacuum  $|0\rangle_{\text{R}}^{\epsilon}$  Effect to the bubble wall to the observer at  $x_{\epsilon}$

$$: \Delta P_{\text{quantum}} = \Delta P_{\epsilon} \quad \Delta P_{\epsilon} = P_{\epsilon} - P_{\text{H}} = \langle T_{xx} \rangle_{\text{IFT}}^{\epsilon} - \langle T_{xx} \rangle_{\text{IFT}}^{\text{H}},$$

the pressure difference  $\rightarrow$  Frictional force

$$\Delta P_{\epsilon} = \frac{b^2}{16\pi} e^{-bx} + \mathcal{O}(e^{-2bx})$$

$\Delta P_{\epsilon} > 0$  : Quantum effect of the vacuum contributes to the friction opposing the bubble wall expansion.



$$\begin{aligned} \langle T_{tt} \rangle_{\text{R}}^{\epsilon} &= -\frac{b^2}{48\pi} e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{xx} \rangle_{\text{R}}^{\epsilon} &= \frac{b^2}{16\pi} e^{-bx} + \mathcal{O}(e^{-2bx}), \\ \langle T_{tx} \rangle_{\text{R}}^{\epsilon} &= \langle T_{xt} \rangle_{\text{R}}^{\epsilon} = \mathcal{O}(e^{-2bx}). \end{aligned}$$

# Conclusions

- Developed a supersymmetric inhomogeneous field theory for various models
- Obtained inhomogeneous BPS vacuum solutions for models like CSH, AH, and others.
- Calculated the renormalized Hadamard two-point function and the VEV of the EM tensor, showing quantum effects in IFTs
- Investigated the quantum frictional effect on bubble wall expansion during the electroweak first-order phase transition
- Trials: Extend the framework to higher-dimensional models and quantum effects including fermion fields

**Thank you for attention!!**