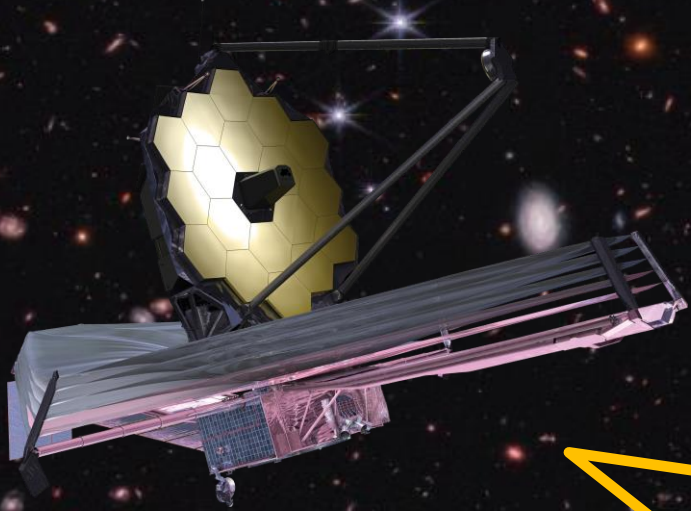


Survey Science Group Workshop 2025

(Session I . Space Missions)

Recent Studies on High-redshift Galaxies with JWST



Jeong Hwan Lee
(Seoul National Univ.)

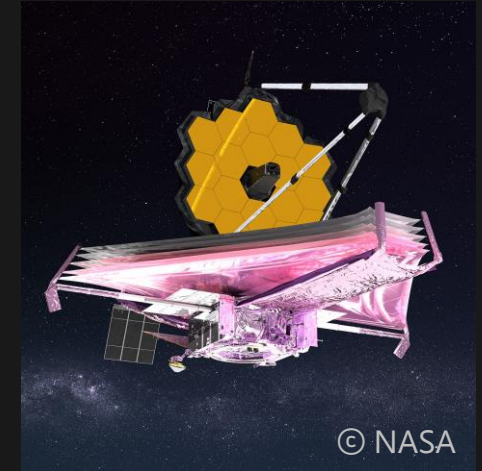
CEERS Deep Field
(NASA, ESA, CSA, and S. Finkelstein)

Webb's Science Themes
(© E. Wheatley)

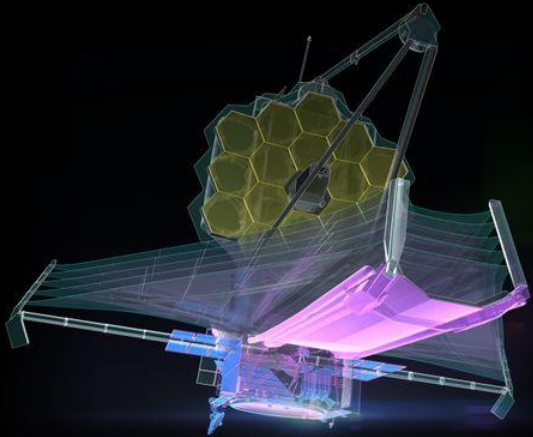
JWST: The Most Powerful Space Telescope

❖ James Webb Space Telescope (JWST)

- Launched on 2021 Christmas, Data released from July 2022
- The largest (6.5m) space telescope
- Covering near- to mid-infrared wavelengths



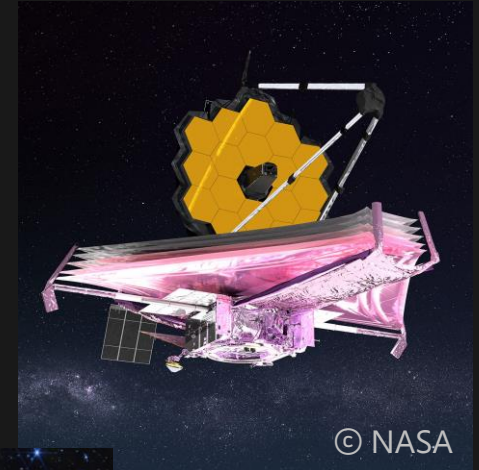
INTO
UNKNOWN
THE JAMES WEBB SPACE TELESCOPE



JWST: The Most Powerful Space Telescope

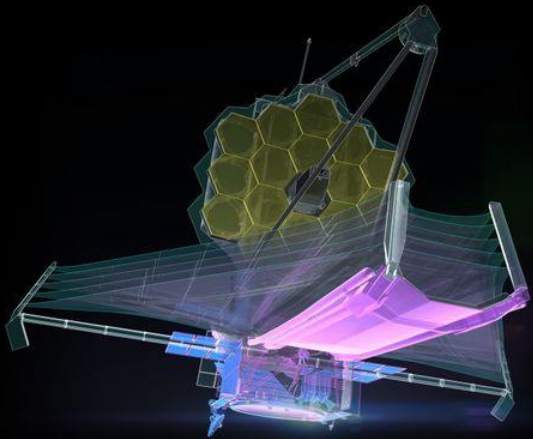
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© NASA

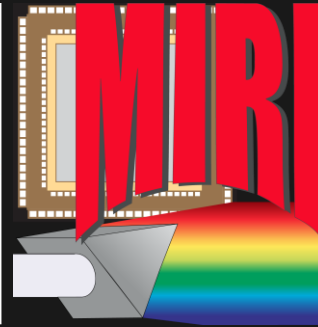
INTO THE UNKNOWN
THE JAMES WEBB SPACE TELESCOPE



Instruments on JWST

❖ Four main instruments on JWST:

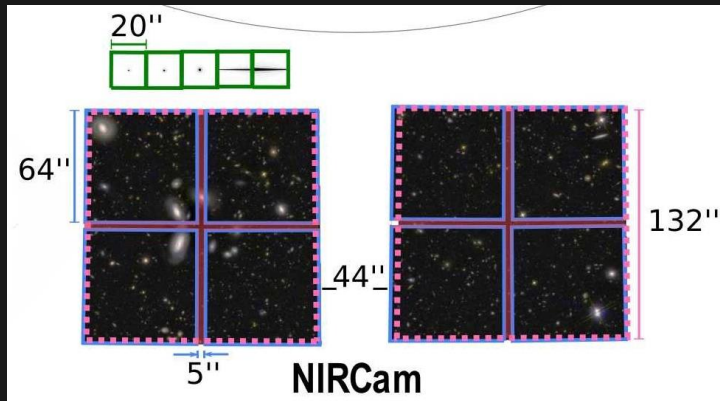
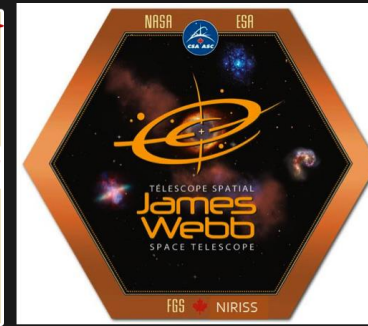
- Near Infrared Camera (**NIRCam**)
- Near Infrared Spectrograph (**NIRSpec**)
- Near Infrared Imager and Slitless Spectrograph (**NIRISS**)
- Mid Infrared Instrument (**MIRI**)



Instruments on JWST

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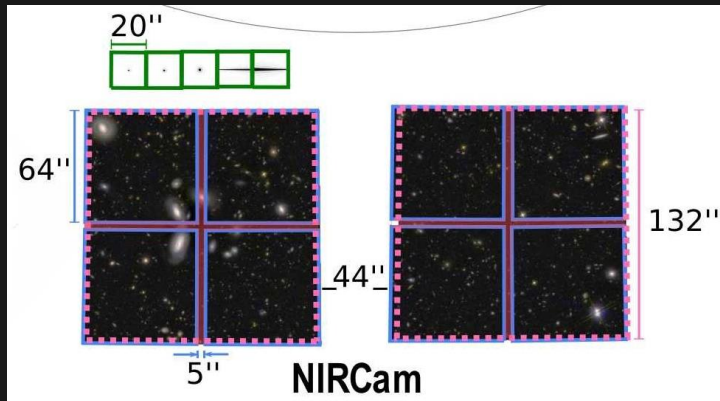
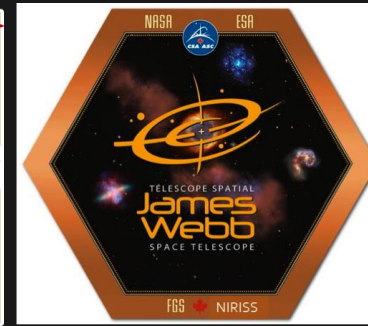
NIRCam

- ✓ Imaging
- ✓ Coronagraphic imaging
- ✓ Wide field slitless spectroscopy

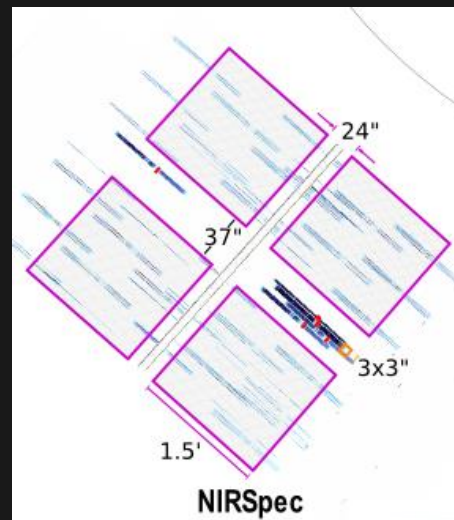
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NIRCam



NIRSpec

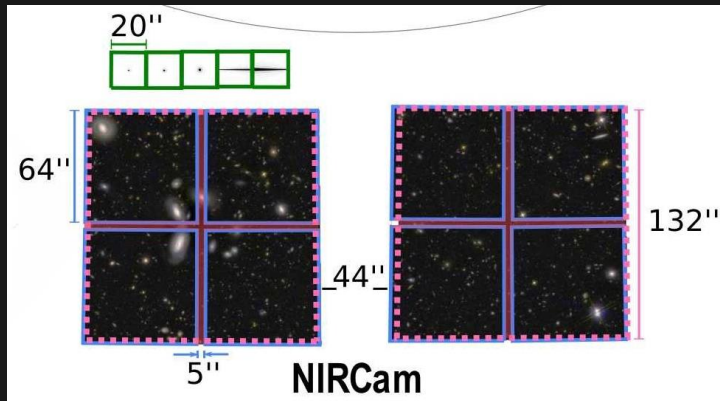
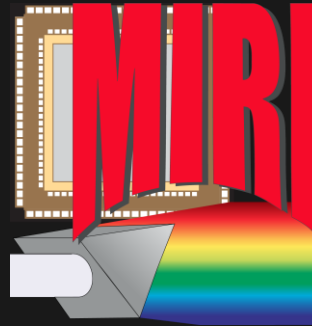
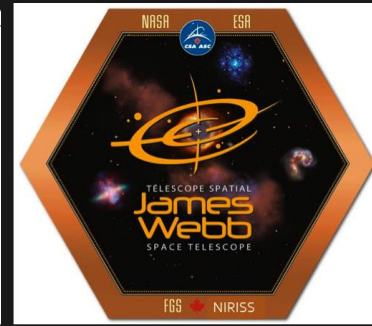
- ✓ Imaging
- ✓ Coronagraphic imaging
- ✓ Wide field slitless spectroscopy

- ✓ MOS spectroscopy
- ✓ IFU spectroscopy
- ✓ Fixed slit spectroscopy

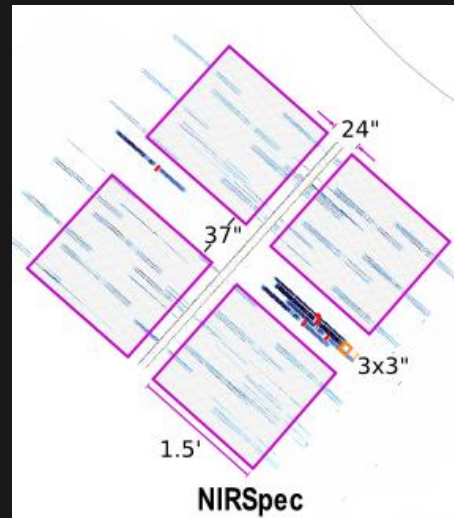
Instruments on JWST

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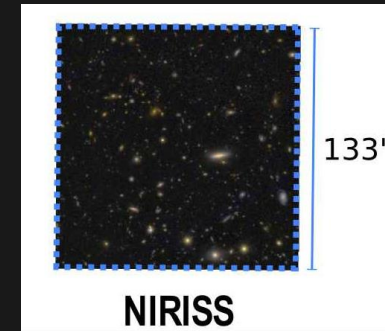
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NIRCam



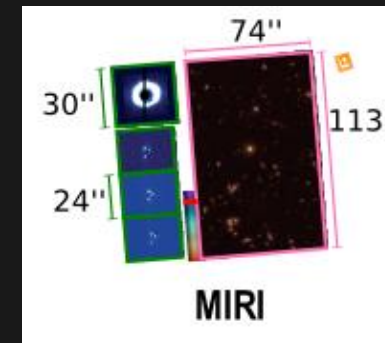
NIRSpec



NIRISS

NIRISS

- ✓ Imaging
- ✓ Single spectroscopy
- ✓ Wide field spectroscopy



MIRI

MIRI

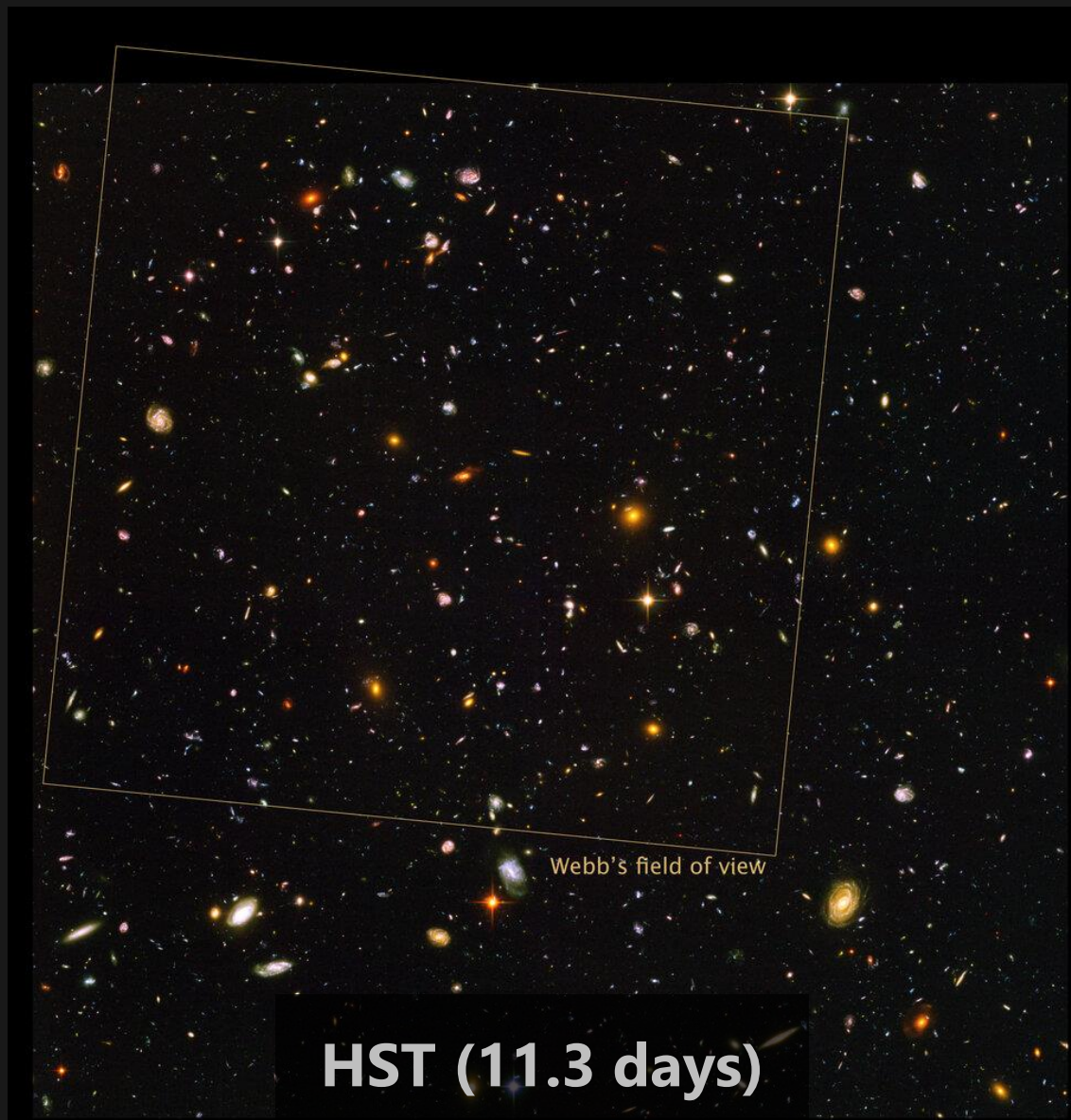
- ✓ Imaging
- ✓ Low-R spectroscopy
- ✓ Mid-R spectroscopy

- ✓ Imaging
- ✓ Coronagraphic imaging
- ✓ Wide field slitless spectroscopy

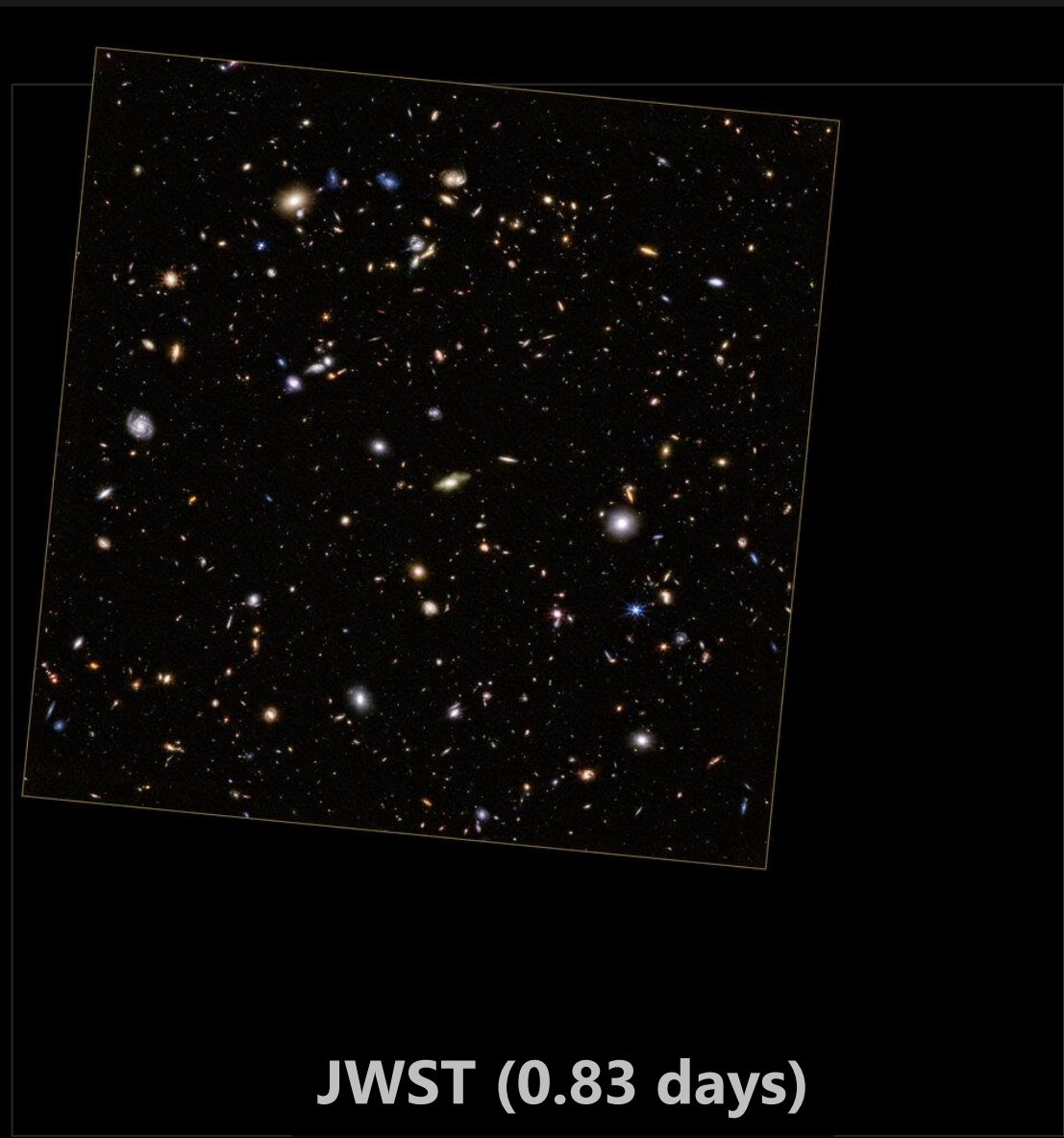
- ✓ MOS spectroscopy
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Performance of JWST: Hubble UDF

Hubble Ultra Deep Field
(NASA, ESA, CSA, and J. DePasquale)



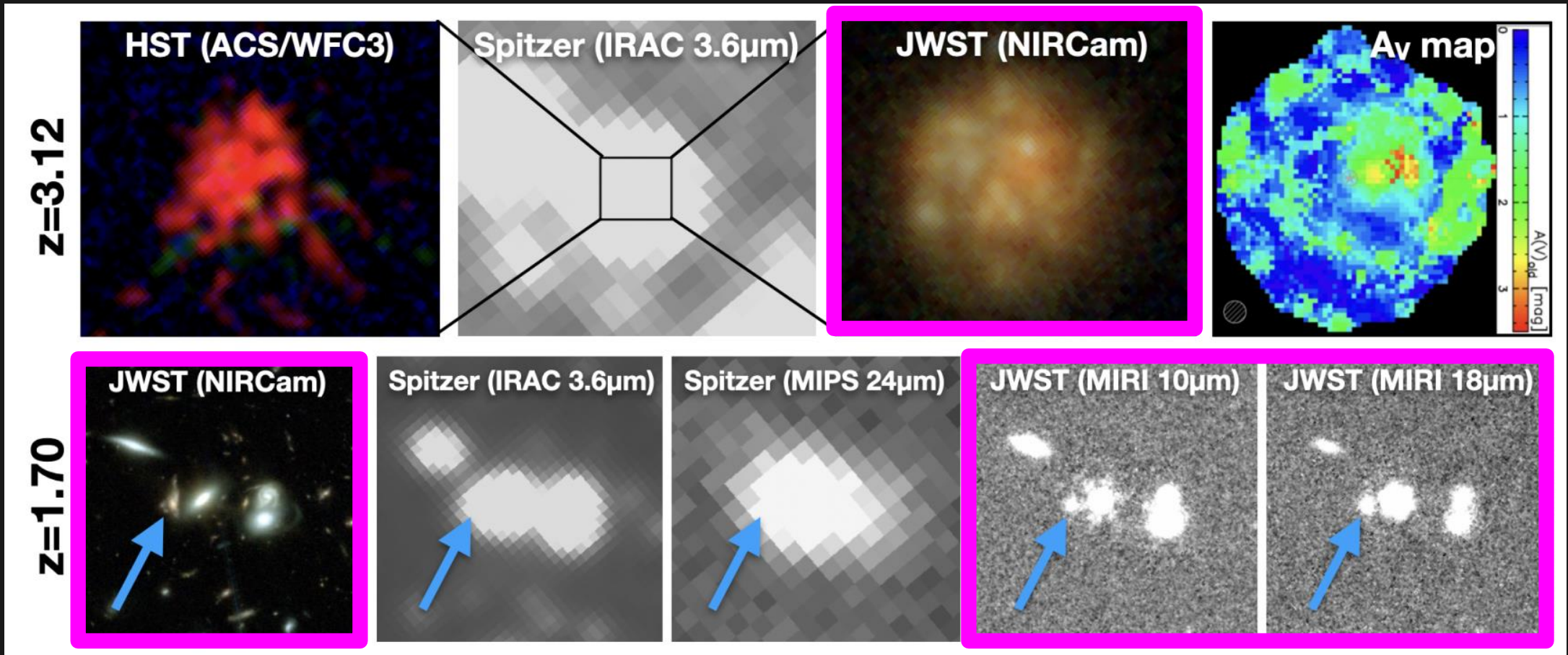
Hubble UDF (exposure time: 11.3 days)



Webb (exposure time: 0.83 days)

Performance of JWST: CEERS

Finkelstein et al. (2025)



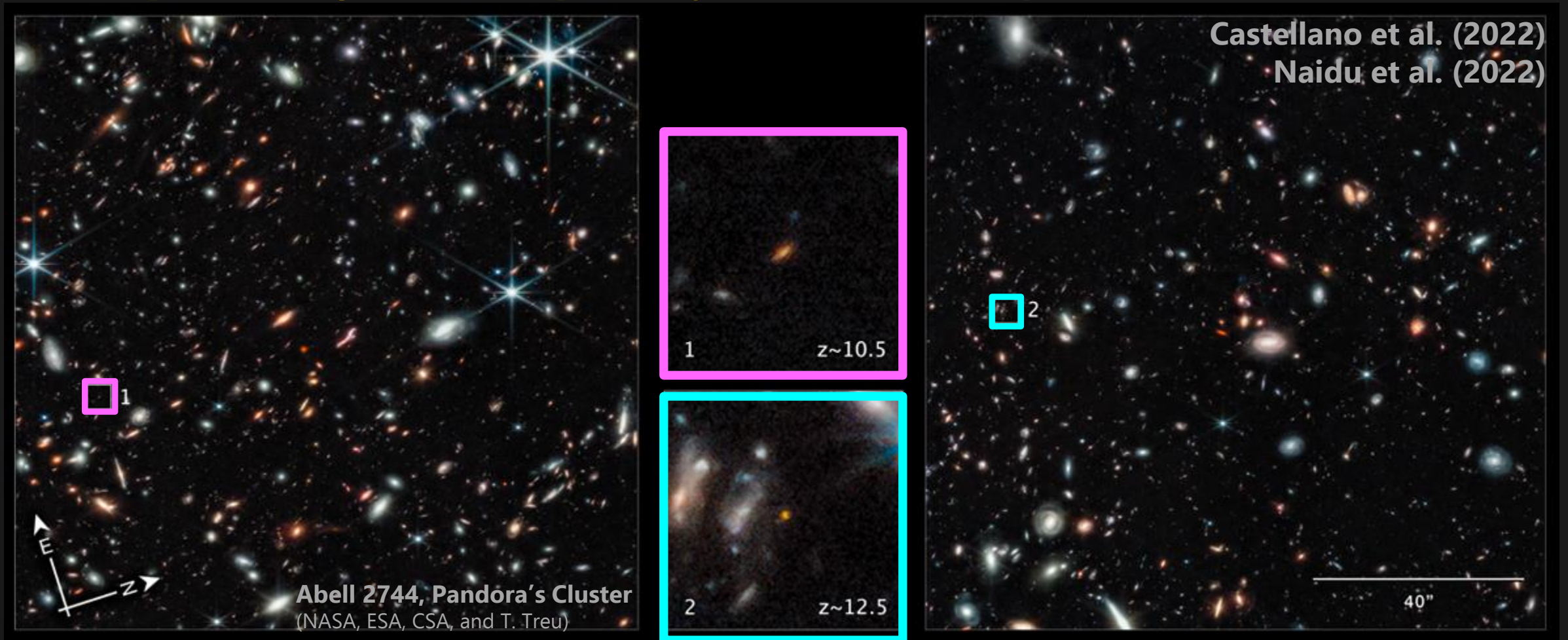
- ❖ Sensitivity of JWST: Unprecedentedly higher than other space telescopes
- ❖ Spatial resolution of JWST: Similar to HST, much better than Spitzer
- ❖ Advantage in wavelength coverage (NIR to MIR) compared to HST (up to 1.6 μm)
→ **Beneficial for studying the high- z universe!**

High-redshift Studies with JWST Instruments

- **Systematic Search for Galaxies in the “Cosmic Dawn” ($z \gtrsim 10$)**
- **Morphological Distributions of Galaxies in the “Cosmic Morning” ($z \sim 4 - 8$)**
- **Detailed Properties of AGNs in the High- z Universe**
- **Detection of Lensed Stars, Star Clusters, and Galaxies**
- **Emission-line Galaxies in the Cosmic Reionization Era**

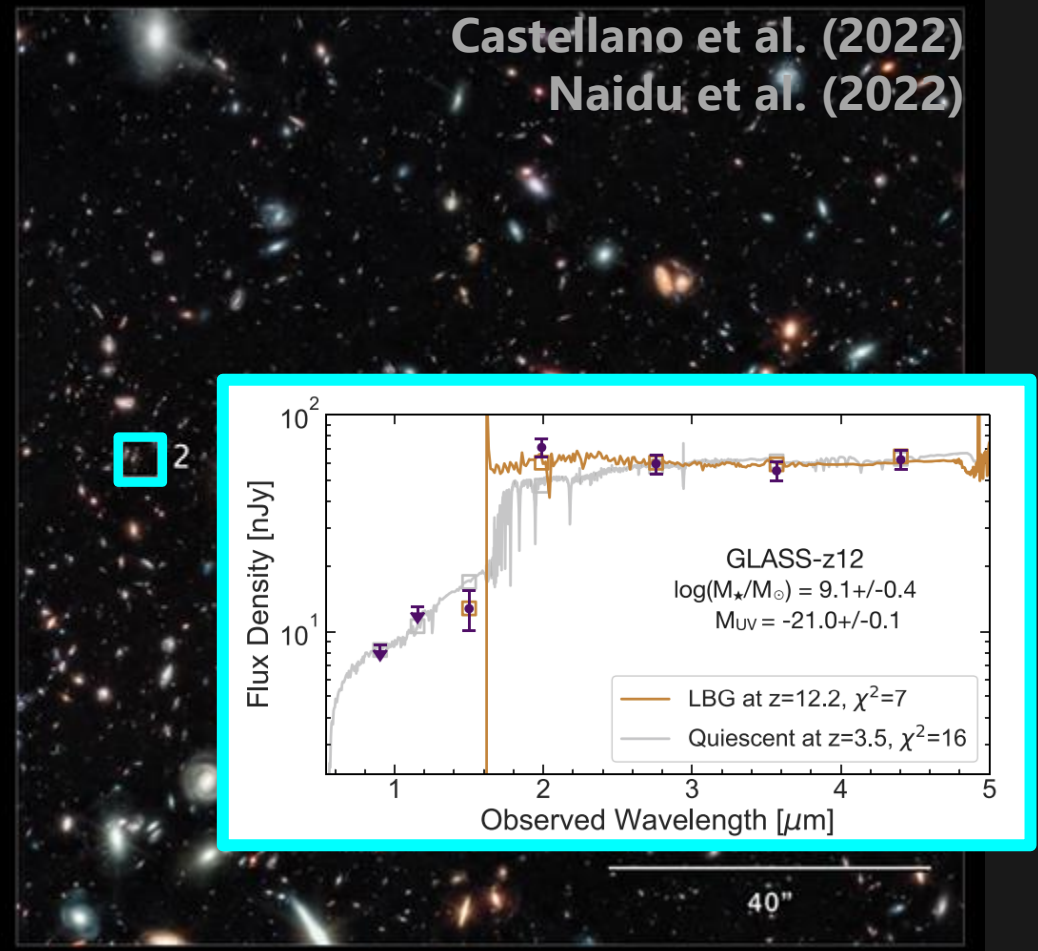
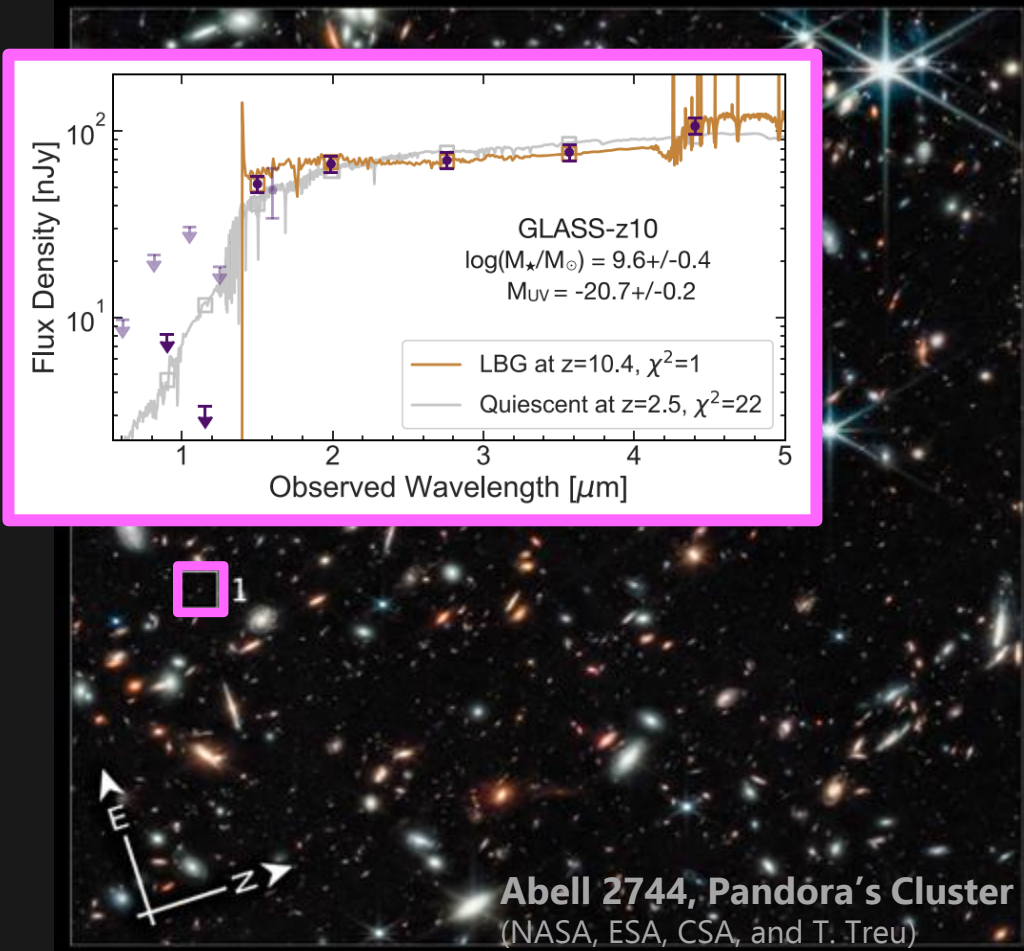
1. Cosmic Dawn Galaxies

- ❖ Extremely bright and massive galaxies at ~ 400 Myr after Big Bang
- ❖ Helpful to understand early galaxy formation!
 - **Rapid mass growth? Top-heavy IMFs due to Pop III stars?**



1. Cosmic Dawn Galaxies

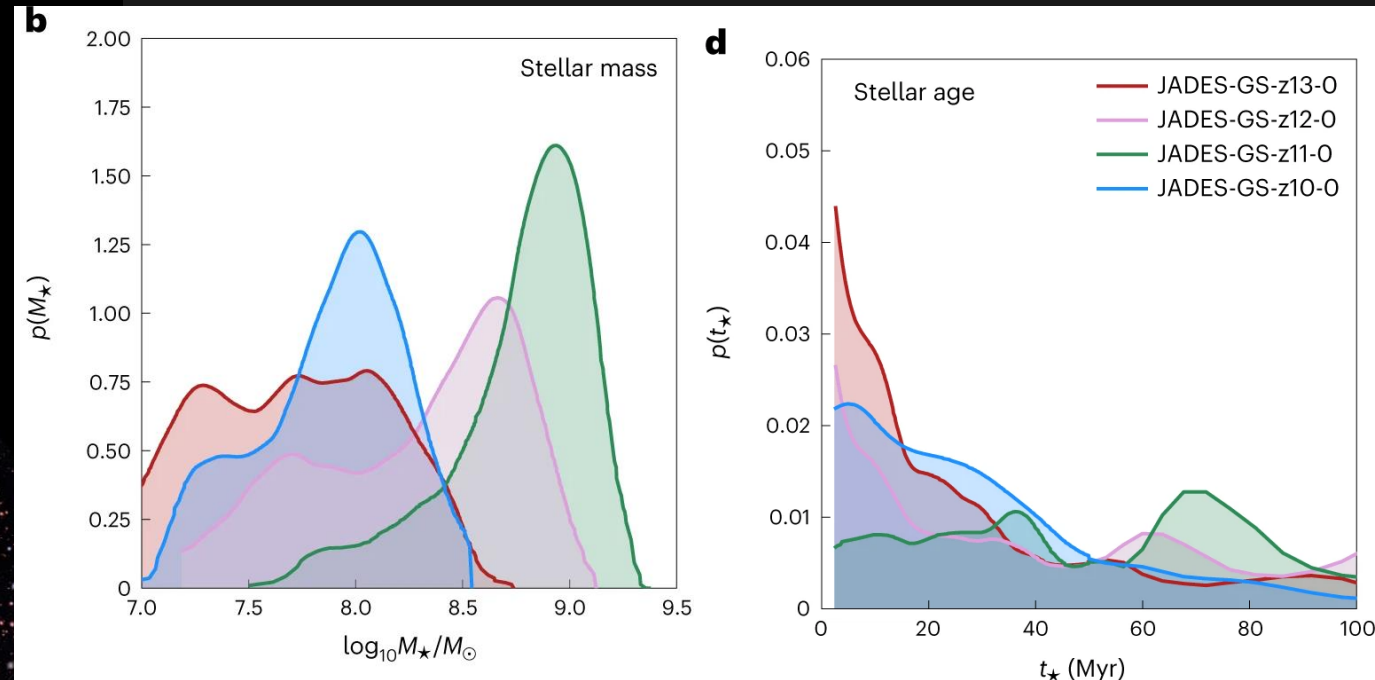
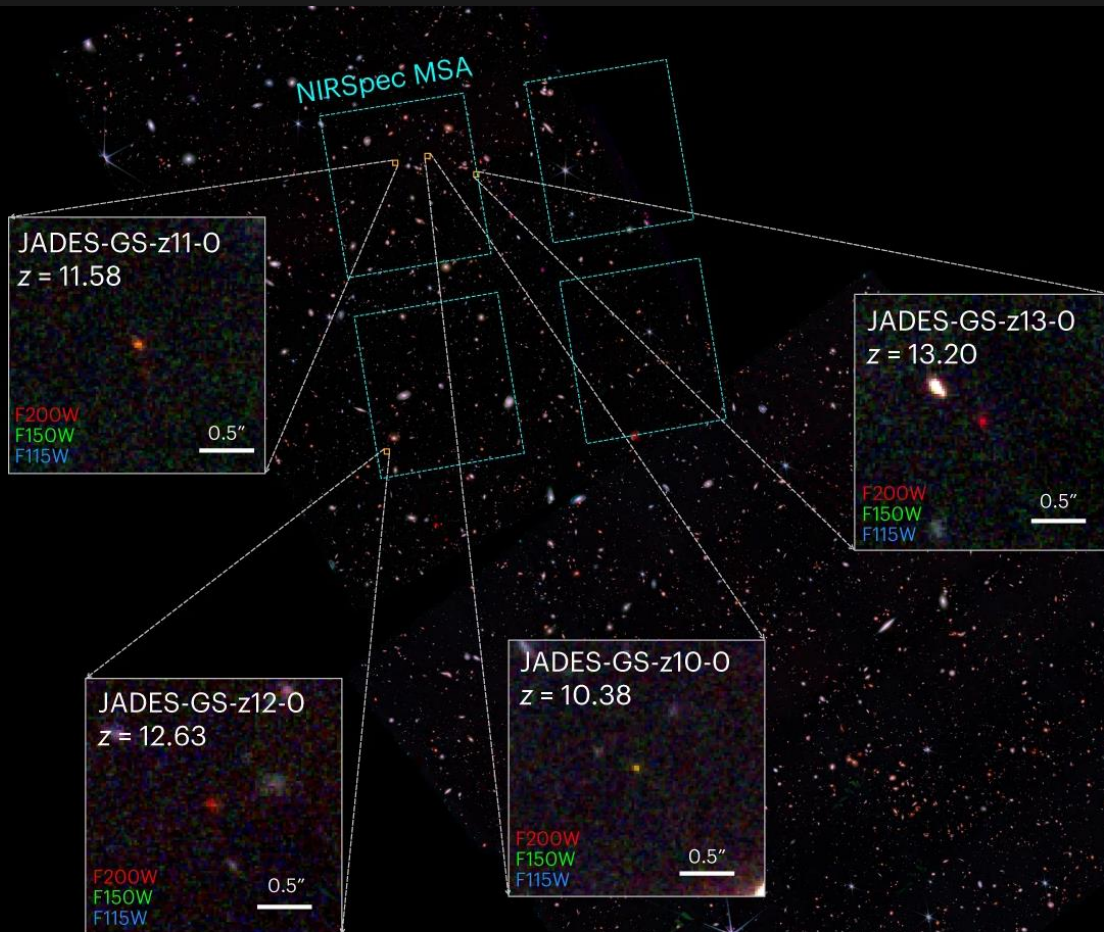
- ❖ Extremely bright and massive galaxies at ~ 400 Myr after Big Bang
- ❖ Helpful to understand early galaxy formation!
 - **Rapid mass growth? Top-heavy IMFs due to Pop III stars?**



1. Cosmic Dawn Galaxies

- ❖ Spectroscopic confirmation of cosmic dawn galaxies with JWST/NIRSpec
- ❖ Metal-poor, young ages, and stellar masses with $10^7 - 10^9 M_{\odot}$

→ **Emergence of the first stars & Large contribution to cosmic reionization**



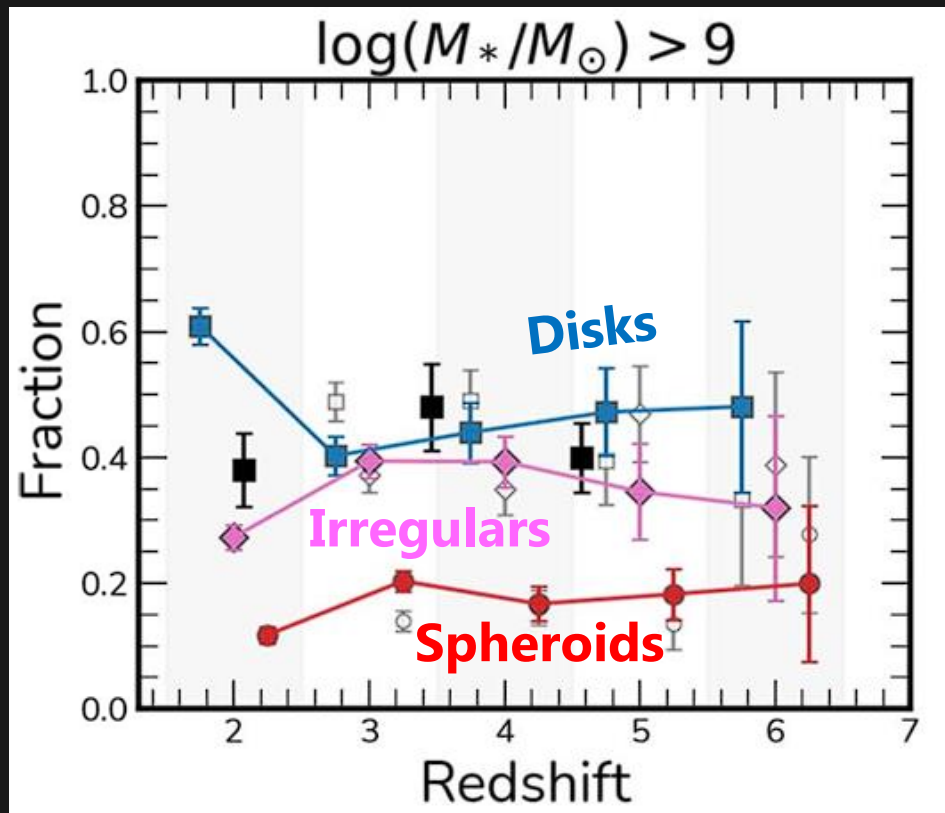
Curtis-Lake et al. (2023)
Robertson et al. (2023)

2. Morphologies of High-z Galaxies

❖ JWST is able to explore rest-frame optical galaxy morphologies beyond $z > 3$!

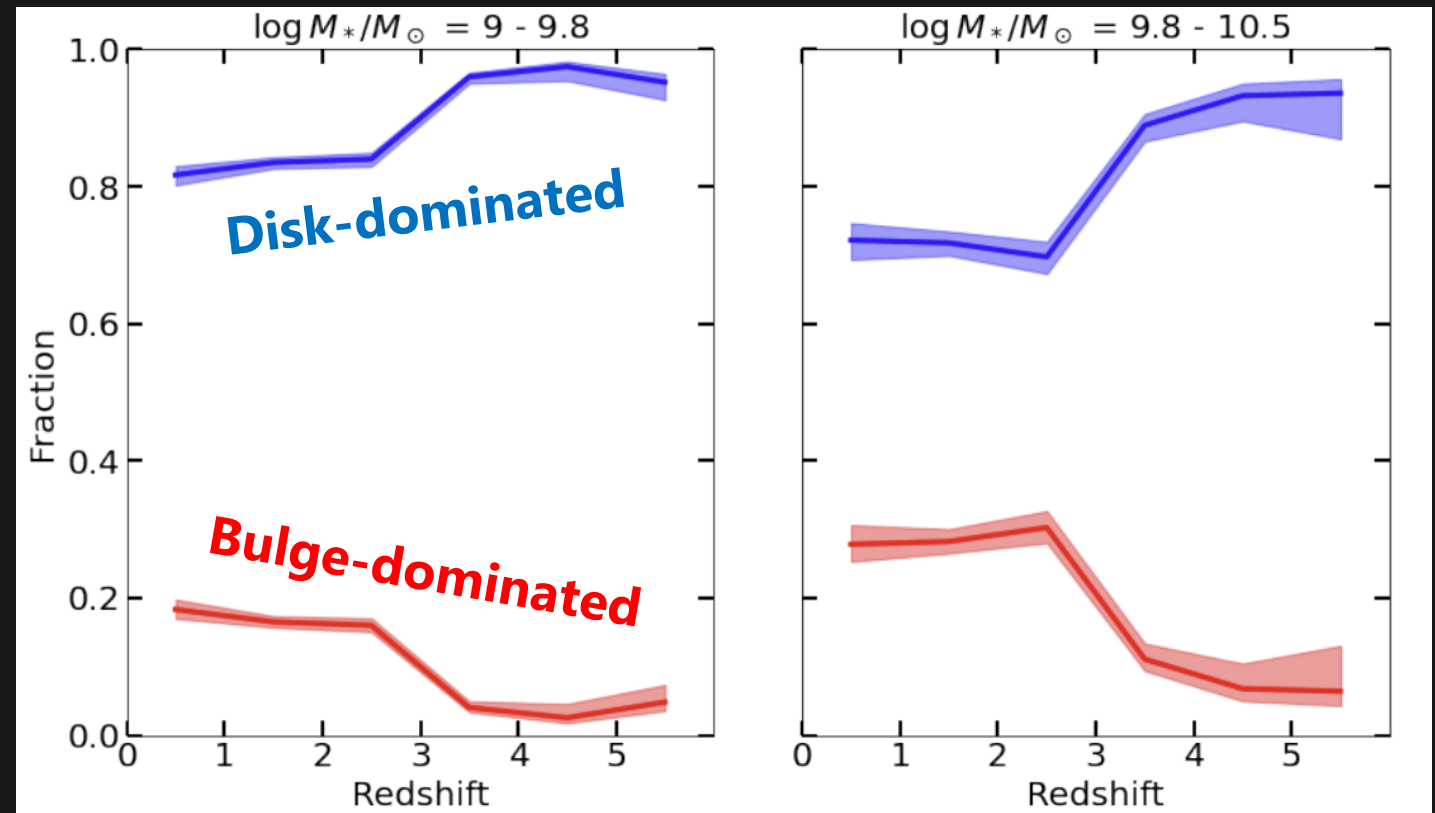
CEERS field (visual classification)

Ferreira et al. (2023)



CEERS field (CNN + Visual classification)

Huertas-Company et al. (2023)



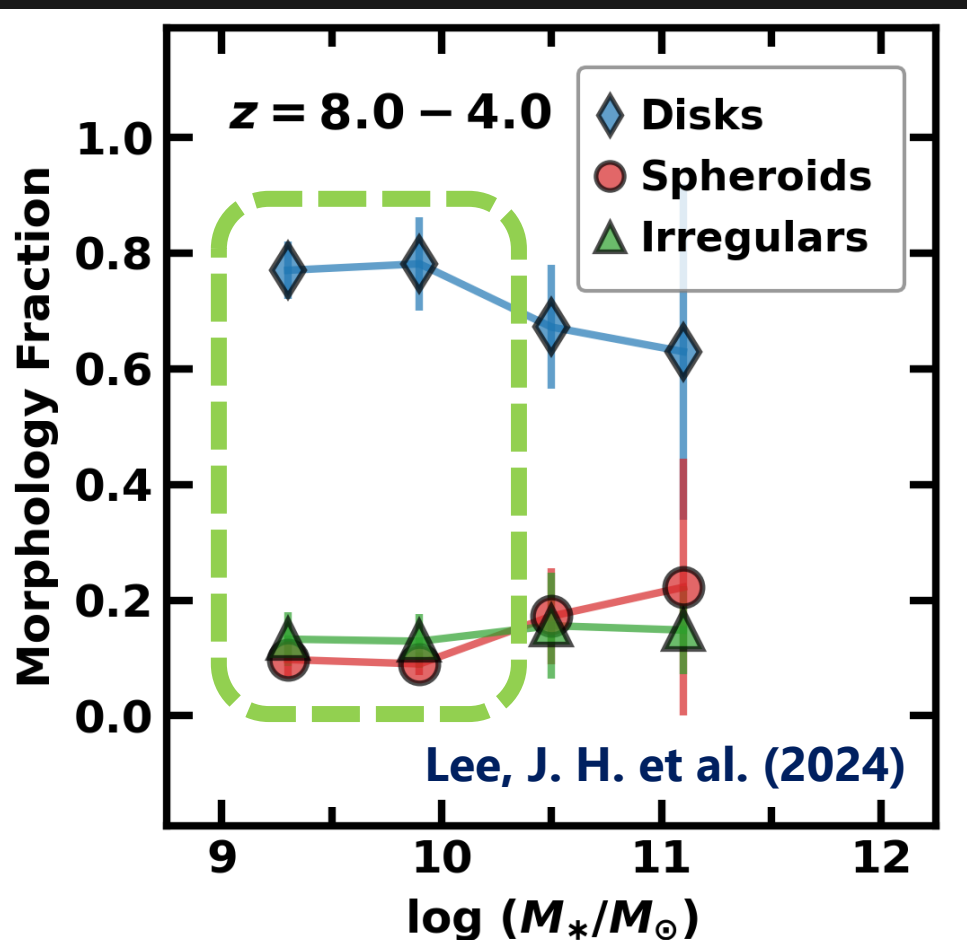
❖ Most JWST studies agree the disk dominance in the early universe.

→ **Early disk formation due to initial angular momentum**

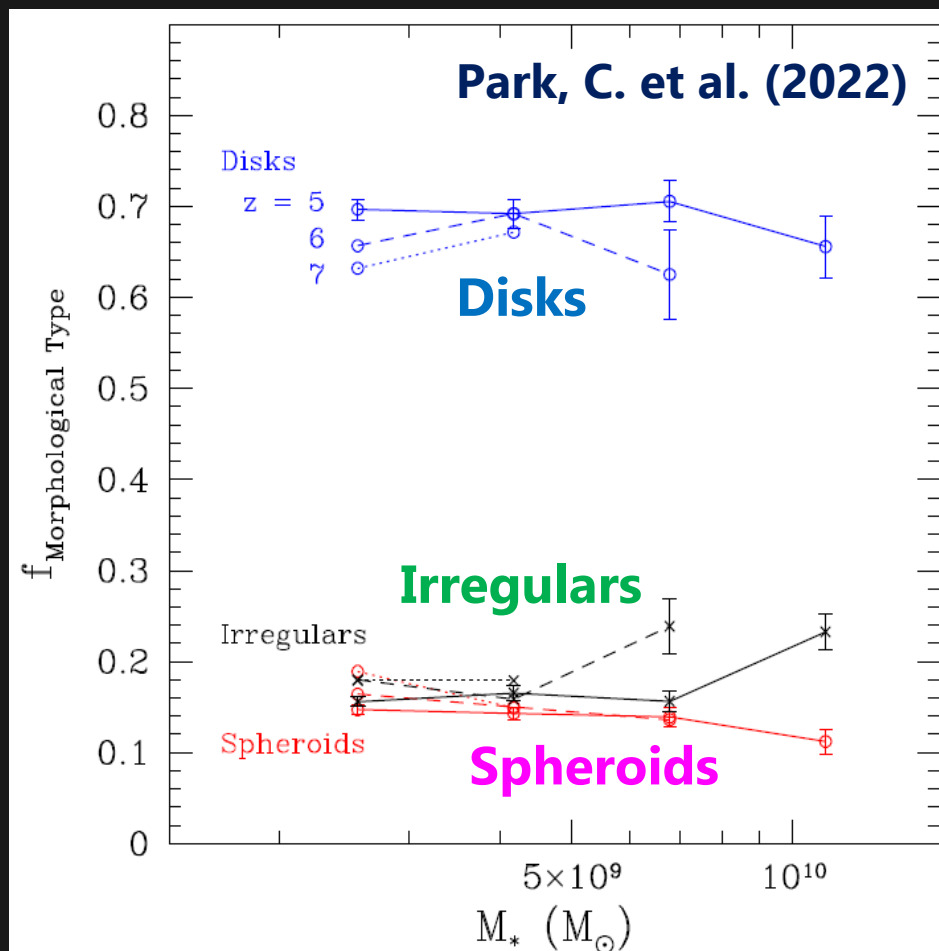
2. Morphologies of High-z Galaxies

❖ JWST observational results are consistent with HR5 cosmological simulation!

JWST (observation)



HR5 (simulation)

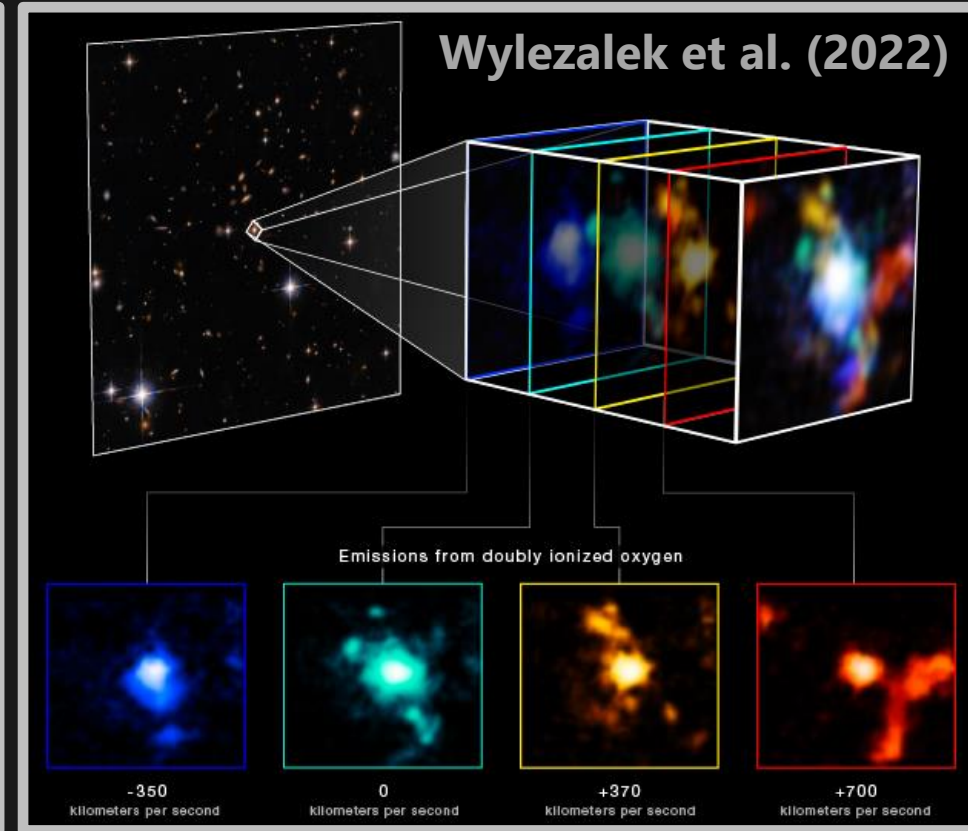
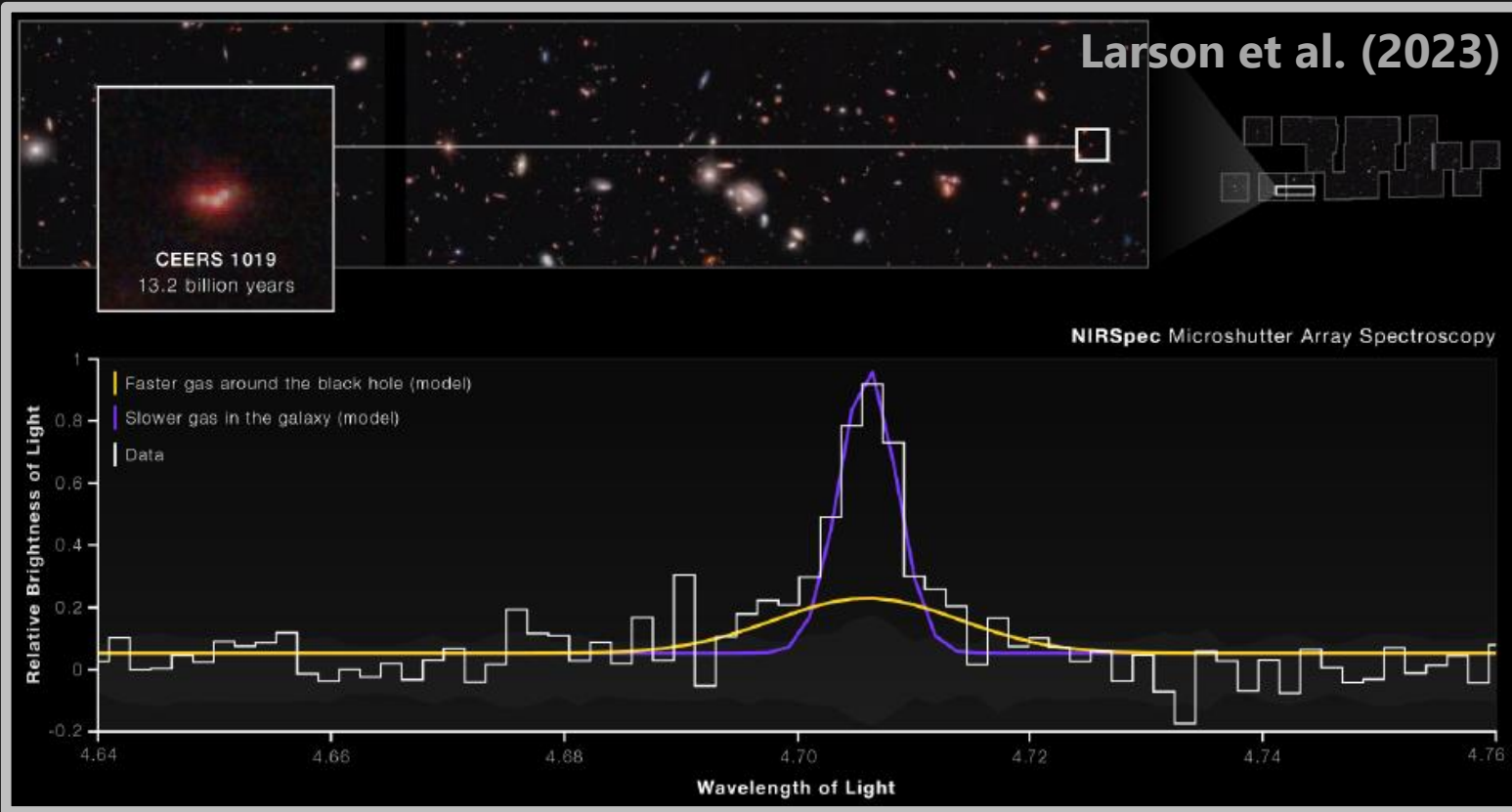


3. Detailed AGN Properties at High Redshifts

❖ Combining JWST imaging and spectroscopy reveals the physics of high-*z* AGNs.

CEERS_1019 (z=8.679)

SDSSJ165202+172852 (z=2.94)



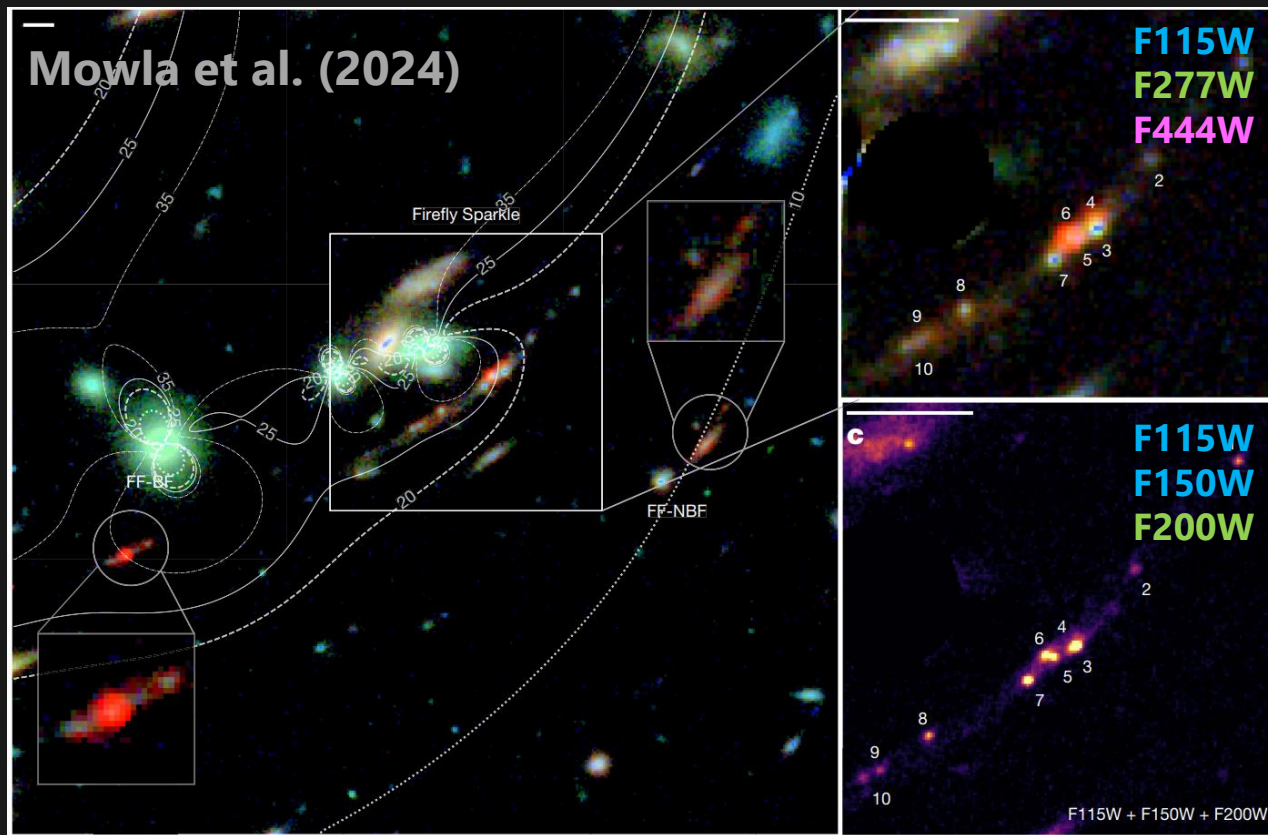
- Highest-redshift AGN
- $\log M_{BH} \sim 7$ with **super-Eddington accretion**

- Complex [OIII] gas kinematics
- **Dense environments**

4. Detection of Lensed Objects

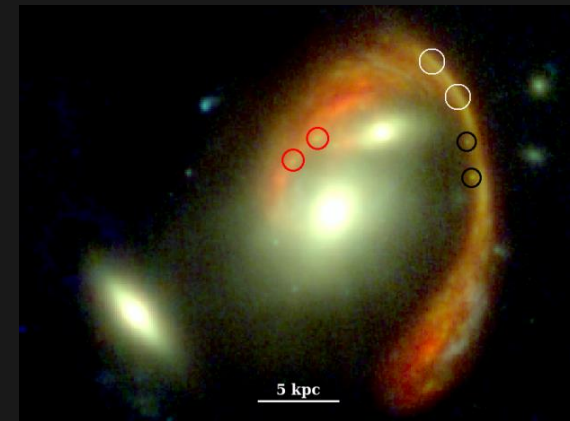
❖ Gravitational lensing provides good opportunity to study distant objects, by magnifying their appearance and brightness.

“Firefly Sparkle” ($z=8.30$)



- Progenitors of massive star clusters

Lensed objects in “El Gordo”



“En Anzuelo”
($z=3.58$)



- A lensed RSG star ($z=2.19$)

5. Emission-line Galaxies in the Reionization Era

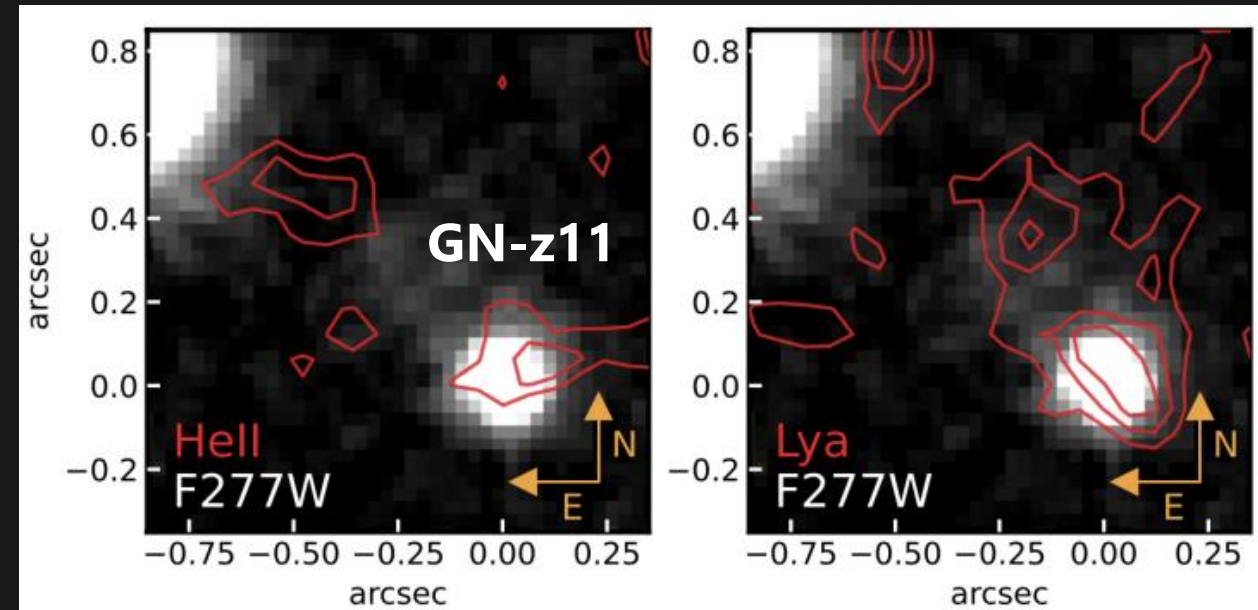
- ❖ Emission-line galaxies found by JWST/NIRCam and NIRSpect are main contributors to cosmic reionization.

EIGER Survey (NIRCam WFSS)



- 117 [OIII] emitters at $z \sim 5-7$
- Young ages, High EWs, High U , Low A_V

Ly α Halo around GN-z11 ($z=10.6$)

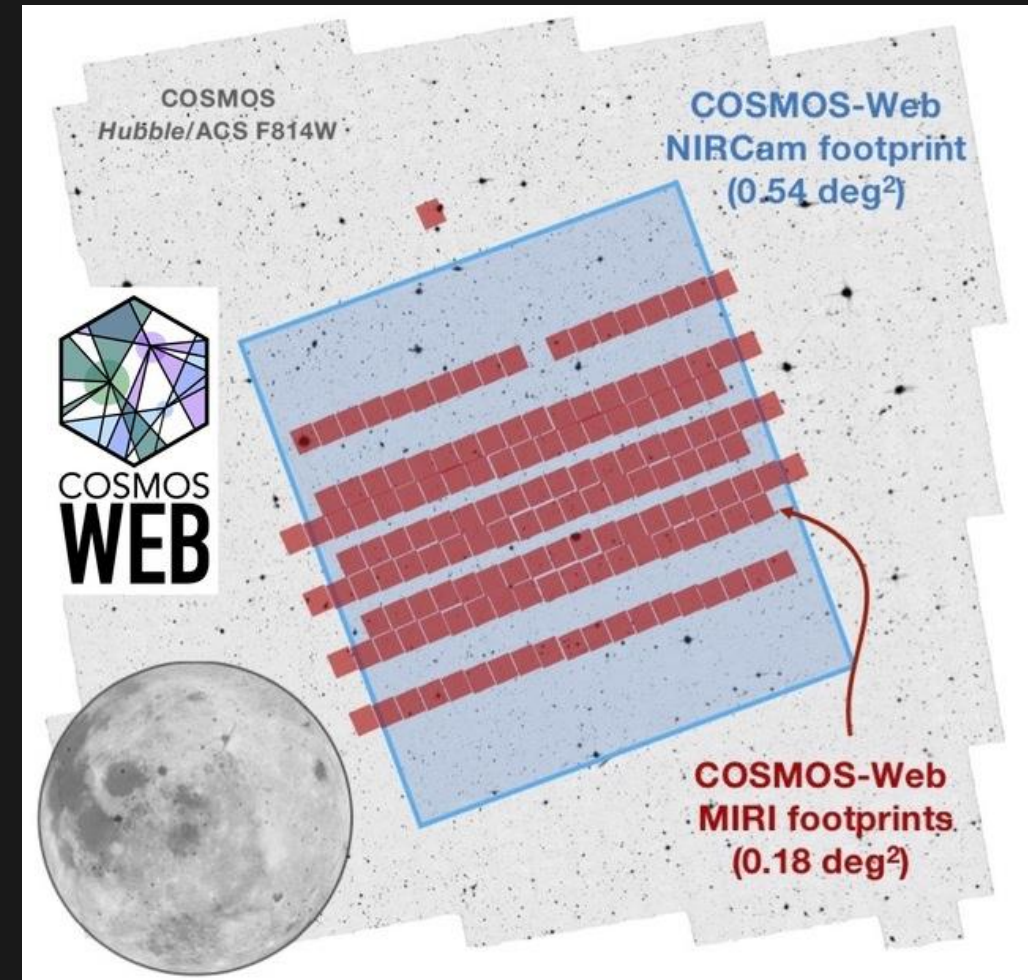


Maiolino et al. (2024)

- High HeII EW, Weak metal lines
→ **Pop III star signature?**

Future Researches using JWST: with COSMOS-Web

- ❖ **Limitation of JWST: narrow field of view**
 - Challenging to study the environmental effect of high-z galaxies
- ❖ **COSMOS-Web survey**
 - Largest JWST field
(**$\sim 0.54 \text{ deg}^2$** \leftrightarrow $\sim 100 \text{ arcmin}^2$ for CEERS)
 - Multiwavelength data from UV to IR
- ❖ Possible sciences
 - Finding protoclusters at high redshifts
 - Environment-Morphology relation
 - Environment-SFR relation
- ❖ Tuning for SED fitting is necessary...

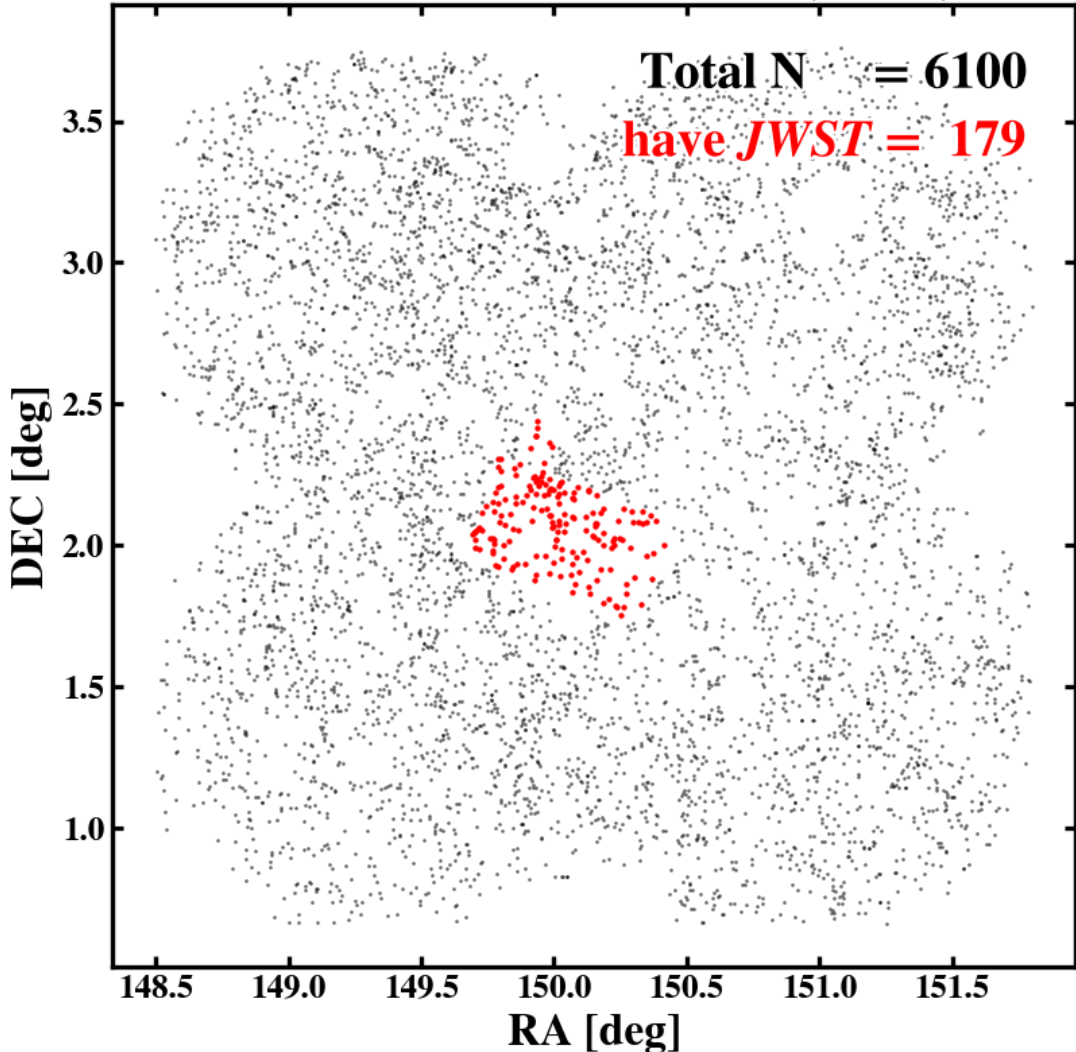


Future Researches using JWST: with ODIN

❖ **ODIN survey:** DECam narrow-band survey for finding LAEs at $z=2.4, 3.1,$ and 4.5

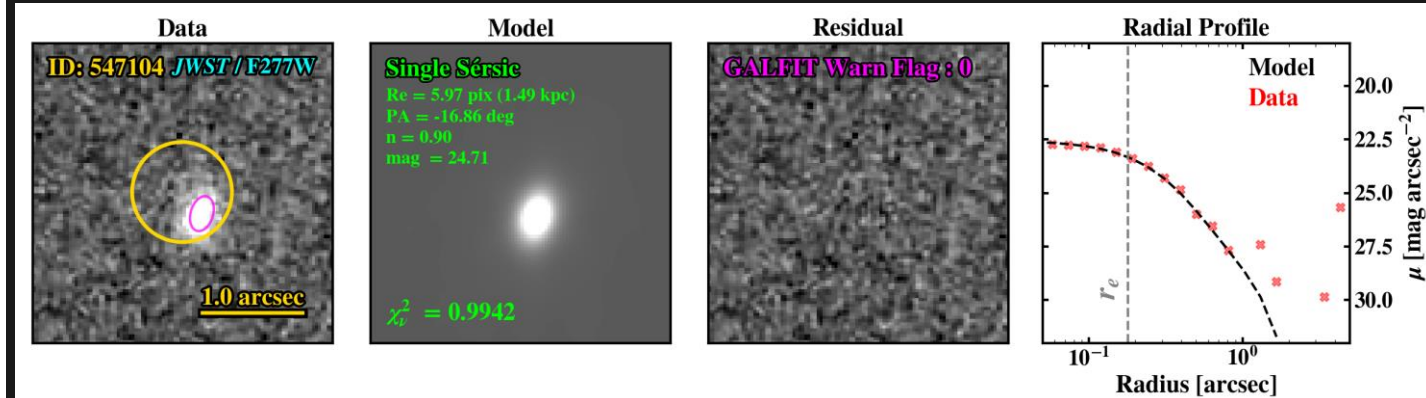
Lee, Kyoung-Soo, et al. (2023)

ODIN COSMOS LAEs ($z\sim 2.4$)



❖ With JWST/NIRCam, **the rest-frame optical morphologies for high- z LAEs** can be investigated!

(Figures provided by Sang Hyeok Im)



GALFIT results

Future Researches using JWST: with K-GMT

- ❖ **With Gemini F2:** spectroscopic observations of high- z quiescent galaxies
 - Fundamental plane (FP) relations: observation vs. simulation
 - **Great opportunity for making synergies with K-GMT program!**

(Figures provided by Priya Goyal)

